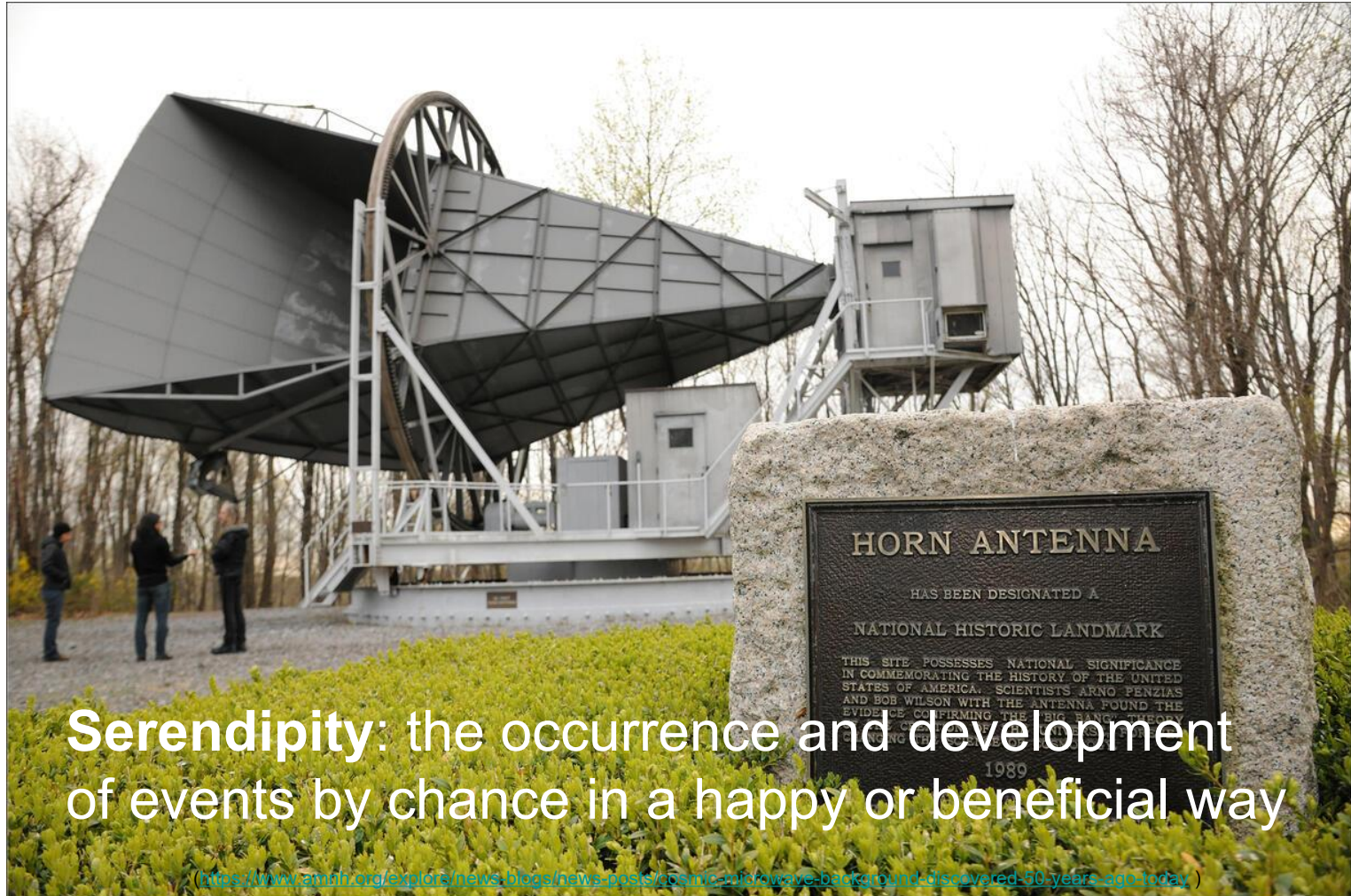


# 1960s: Decade of Serendipity (in Astronomy)



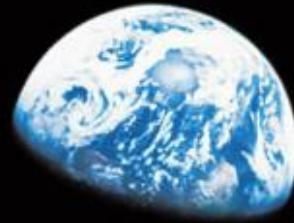
**Serendipity:** the occurrence and development of events by chance in a happy or beneficial way

**John Matthews**

Professor of Physics, U. of New Mexico

# 1960s: when was that?

“Oh, my God! Look at that picture over there! Here’s the Earth coming up. Wow, is that pretty!” Anders exclaimed, while orbiting the Moon.

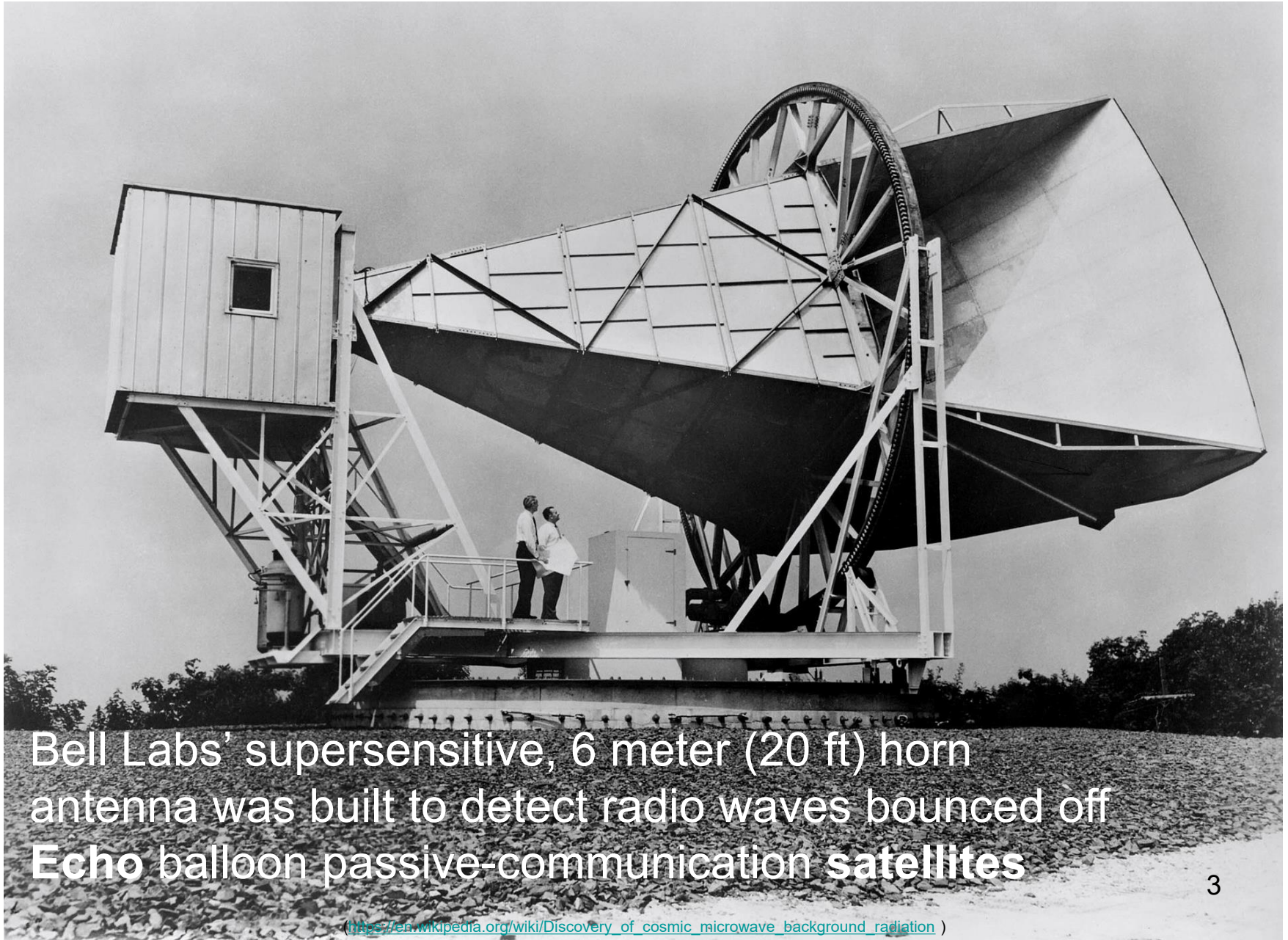


**Earthrise:** William Anders, from **Apollo 8** in **Lunar orbit**, 70 miles above the surface, December 24, 1968

(<http://100photos.time.com/photos/nasa-earthrise-apollo-8>)

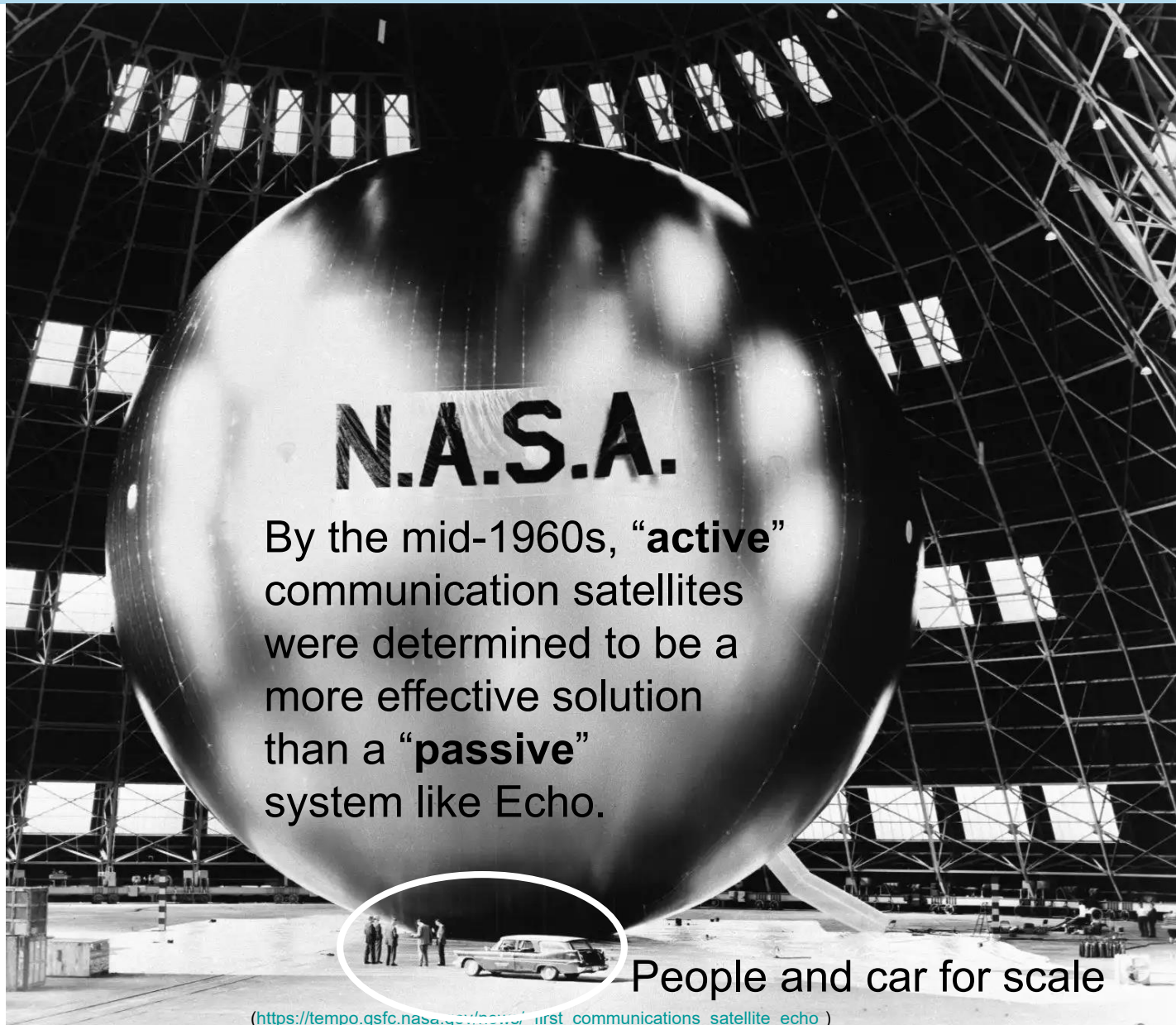


# Echo Project pioneered communications using satellites



Bell Labs' supersensitive, 6 meter (20 ft) horn antenna was built to detect radio waves bounced off **Echo** balloon passive-communication **satellites**

# Echo tried huge, reflective balloons for communications



People and car for scale

([https://tempo.gsfc.nasa.gov/news/first\\_communications\\_satellite\\_echo](https://tempo.gsfc.nasa.gov/news/first_communications_satellite_echo))



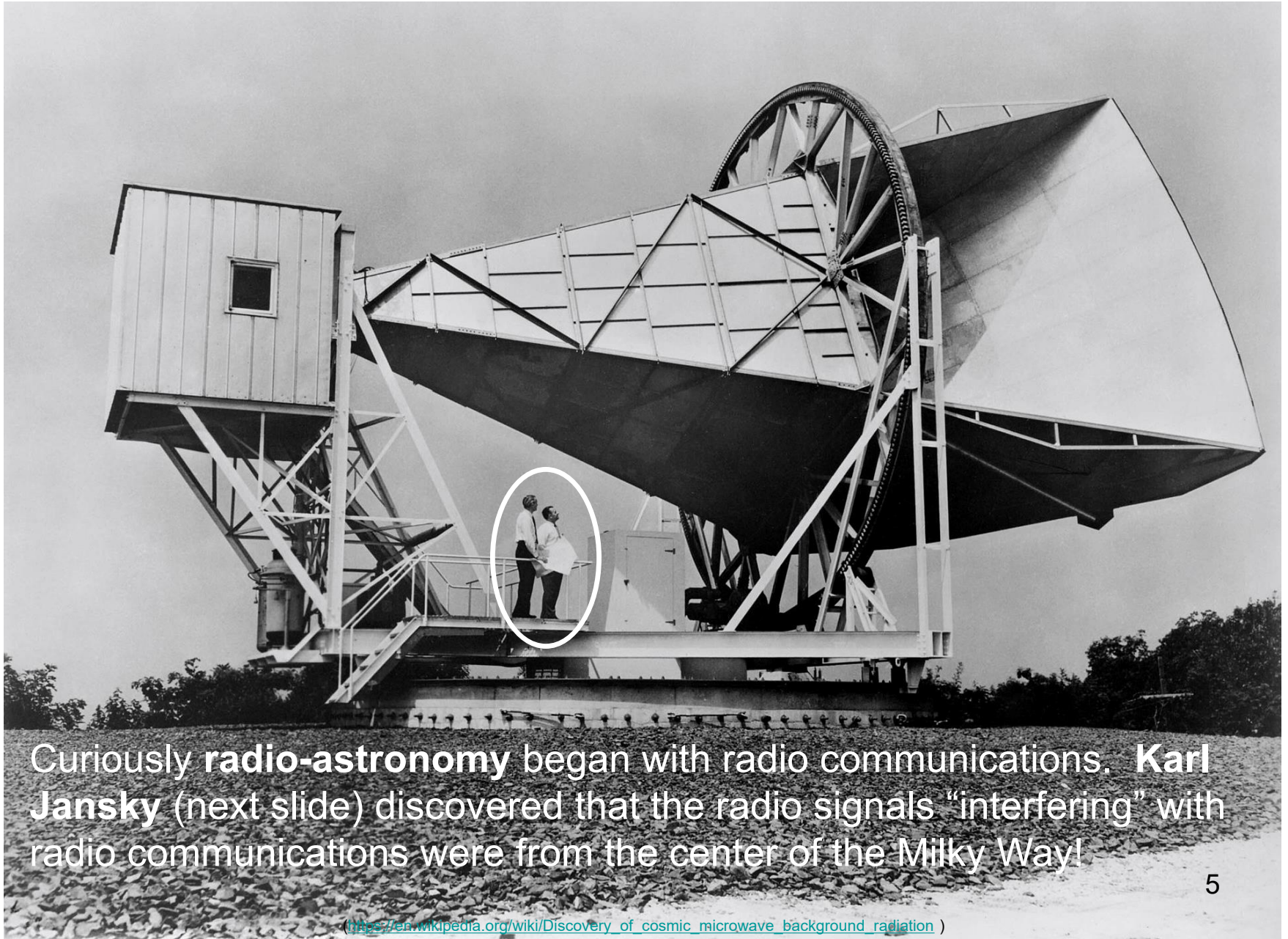
Static Inflation Test of 135 Ft Satellite In Weeksville, NC  
NASA Langley Research Center

6/28/1961

Image # EL-1996-00052

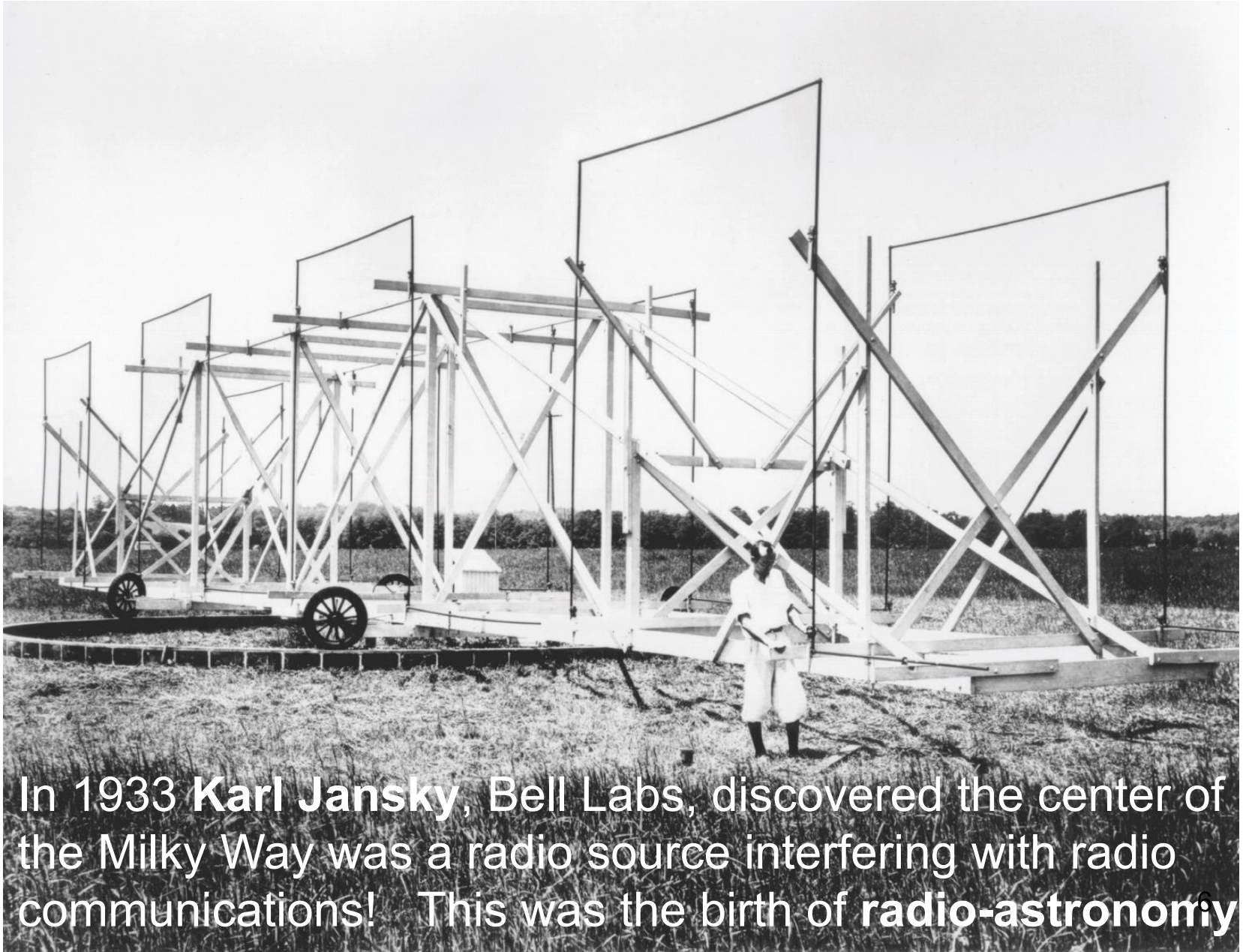


Not needed, “maybe we can do some radio-astronomy with it?”



Curiously **radio-astronomy** began with radio communications. **Karl Jansky** (next slide) discovered that the radio signals “interfering” with radio communications were from the center of the Milky Way!

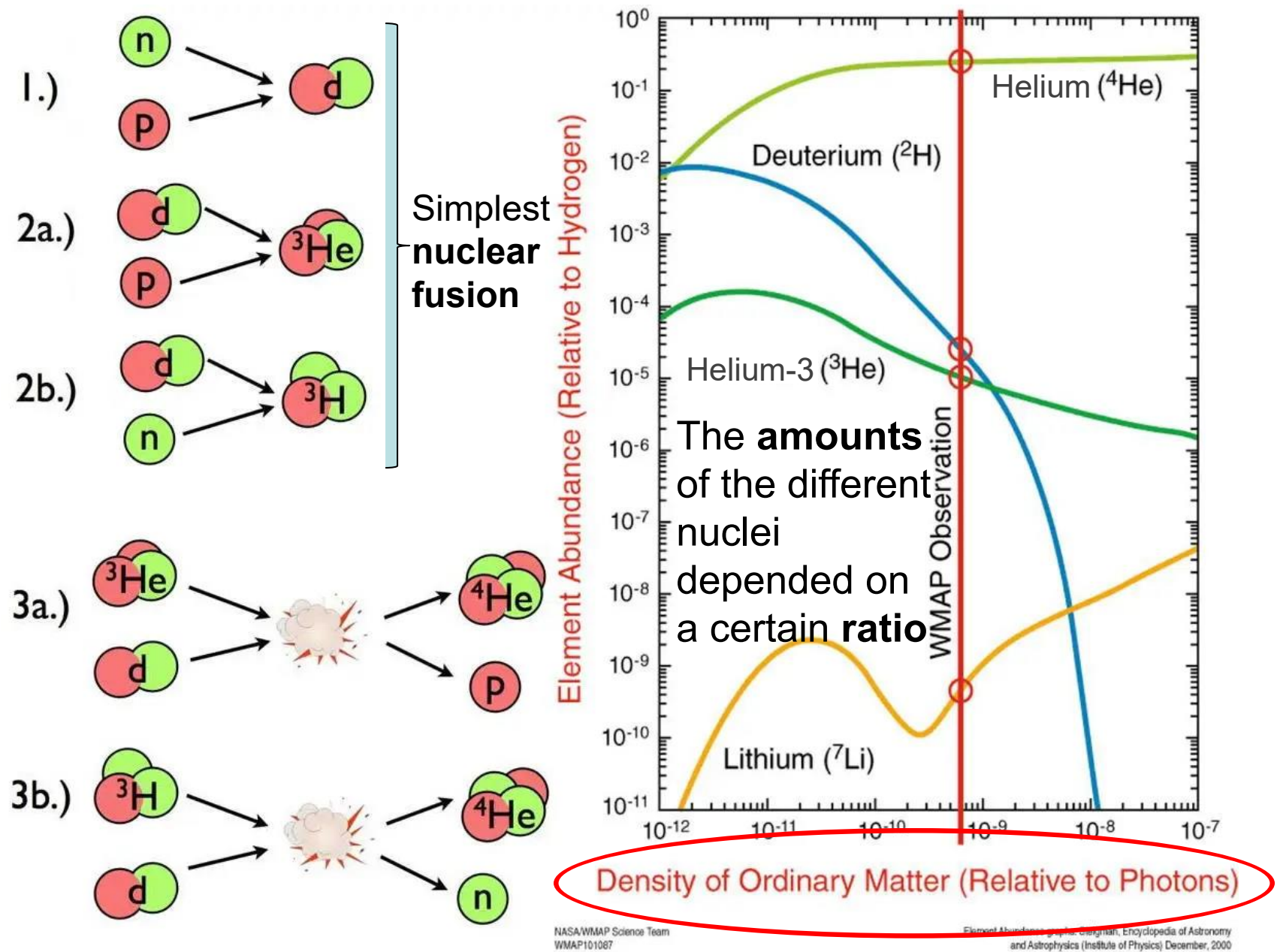
# Milky Way center was the first astronomical radio source



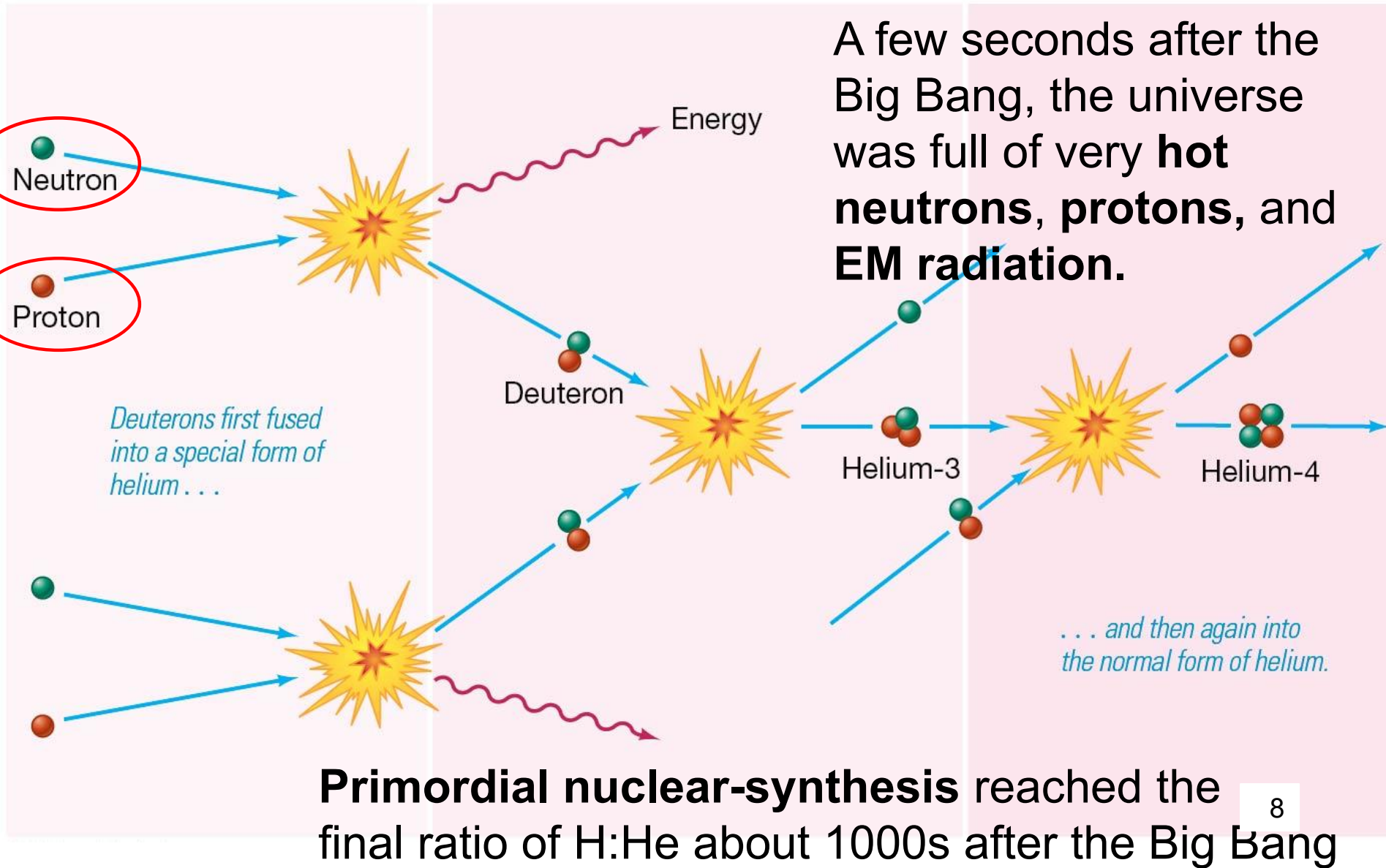
In 1933 **Karl Jansky**, Bell Labs, discovered the center of the Milky Way was a radio source interfering with radio communications! This was the birth of **radio-astronomy**



# About 13.7 billion years before the 1960s, the 1<sup>st</sup> nuclei formed

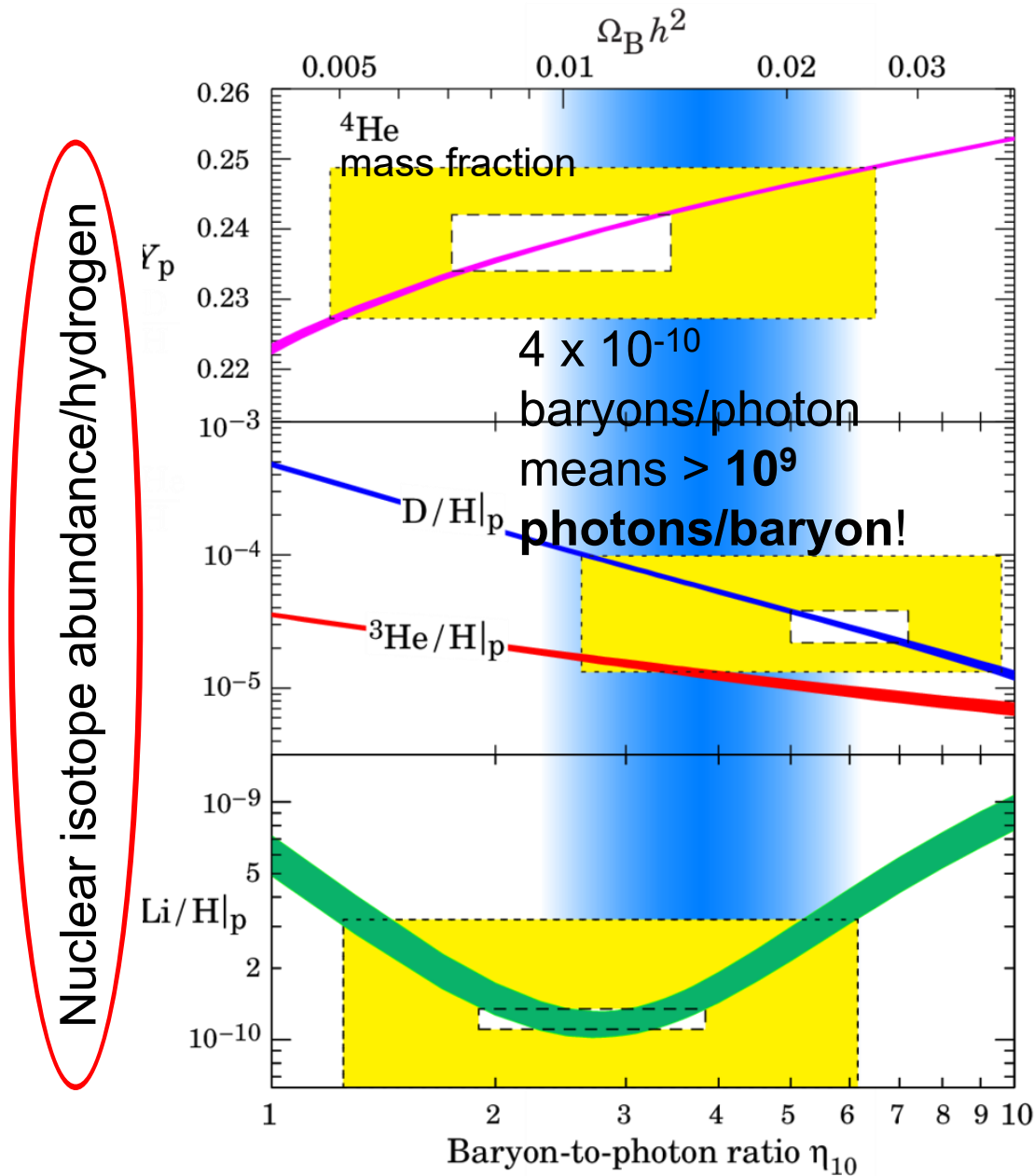


# For nuclear fusion the early universe had to be very hot





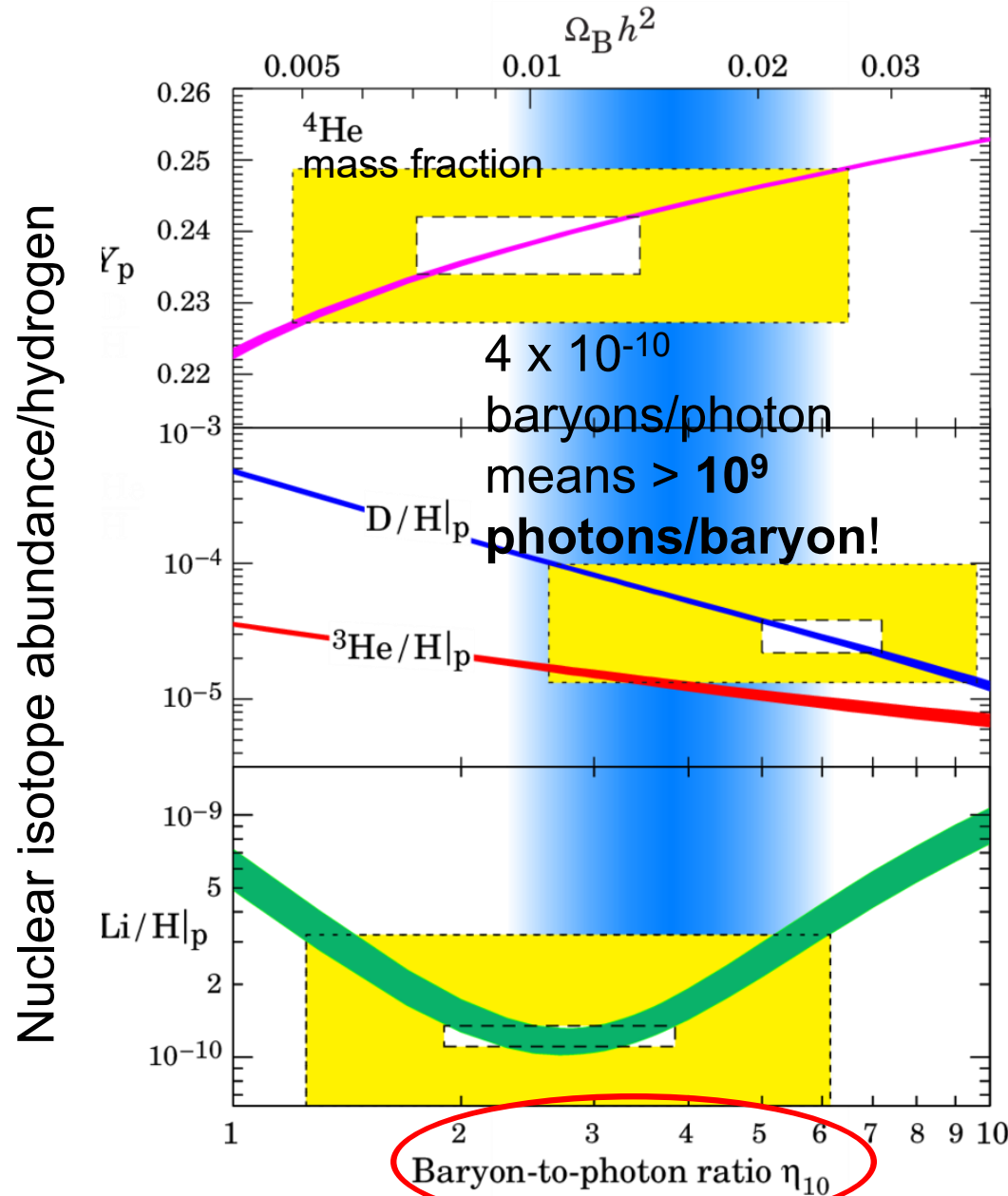
# For some reason primordial fusion mostly stopped at helium!



Curiously the 1st nuclei, and the universe today, are mostly hydrogen and helium (**H** and **He**).

Why is that?

# Theoretical isotope predictions needed a lot of EM radiation

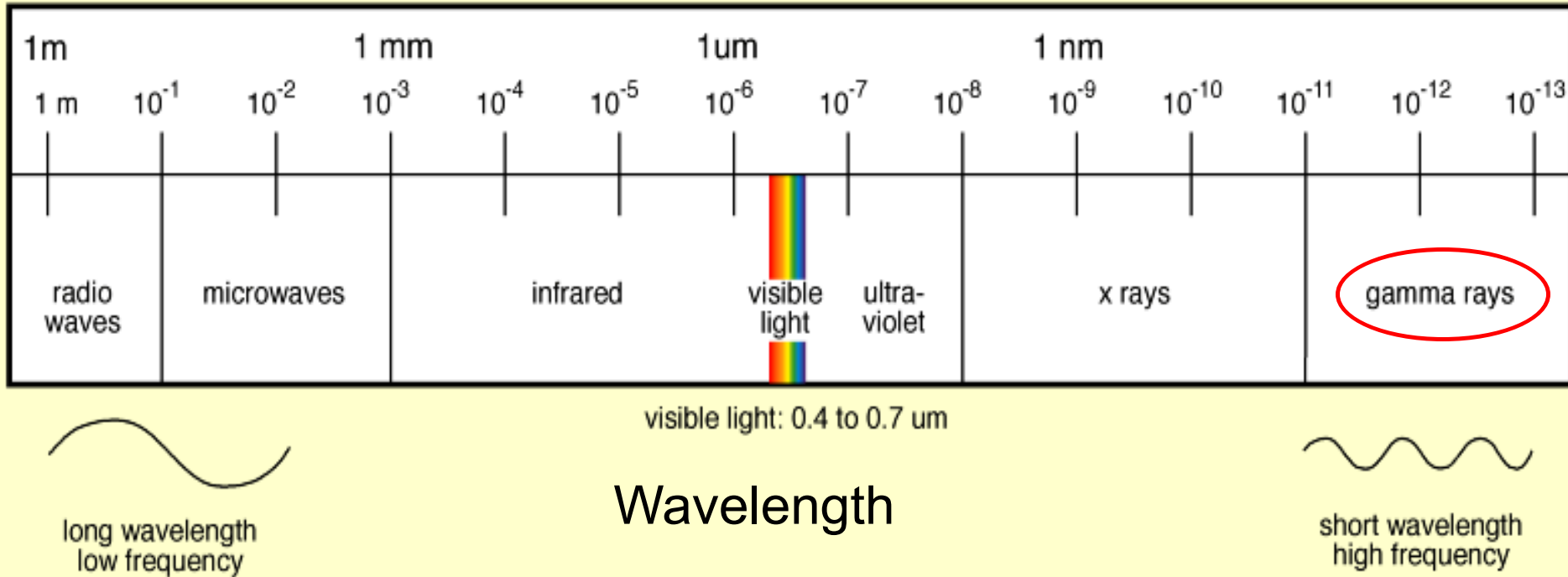


So called **Big Bang Nucleosynthesis**, agreement with measured nuclear abundances needed a “huge flux” of electromagnetic radiation (gamma ray photons).



# EM radiation is a type of wave

*The Electromagnetic Spectrum (range of wavelengths)*



**Gamma-rays** are the EM waves with the **shortest** wavelengths

**Our eyes see visible light, just a tiny part of the EM spectrum**

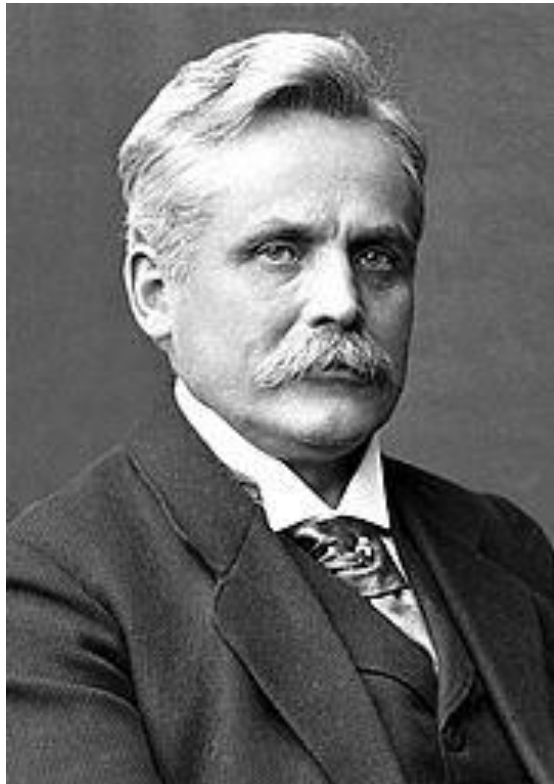
**Electromagnetic radiation (EM):** the transmission of energy through space without physical connection through varying **electric** and **magnetic** fields

Examples: Light from stars and the Sun

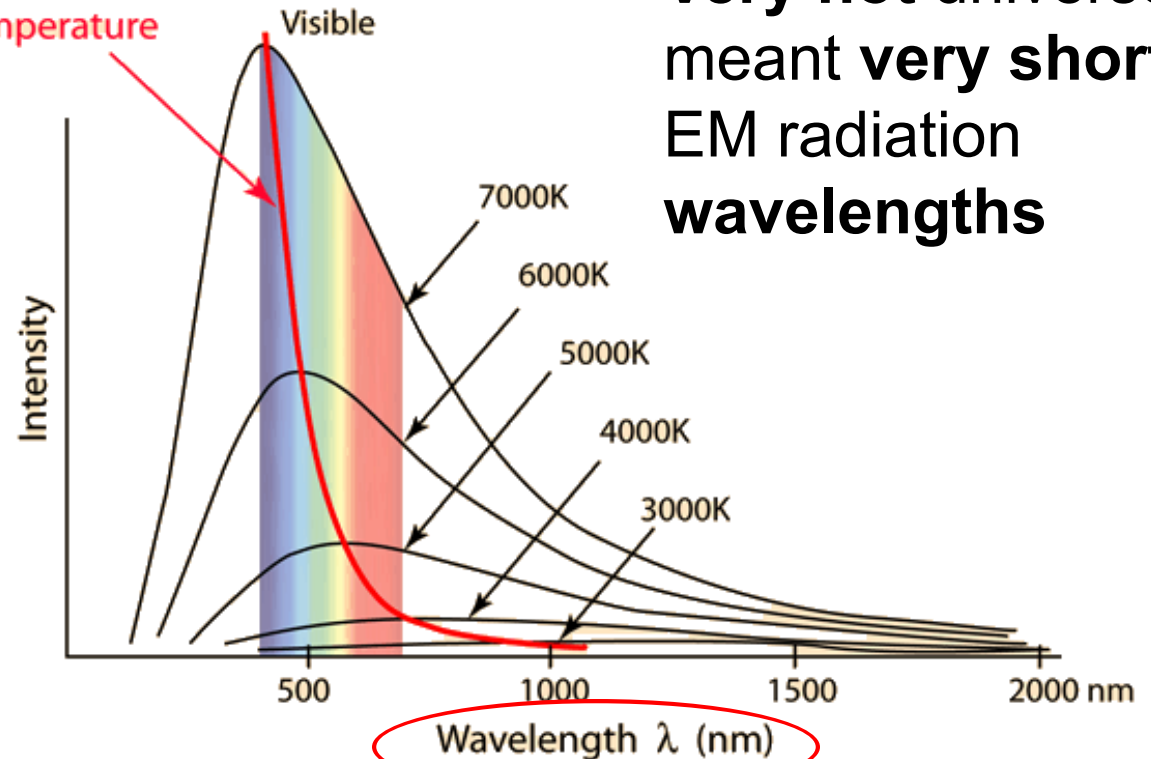


# Hot objects emit a spectrum of wavelengths

In 1893 Wien showed that the **wavelength of peak-emission**,  $\lambda_{\text{peak}}$ , was related to the **temperature** of the source,  $T(\text{K})$ : the **hotter** the source the **shorter** the wavelength.



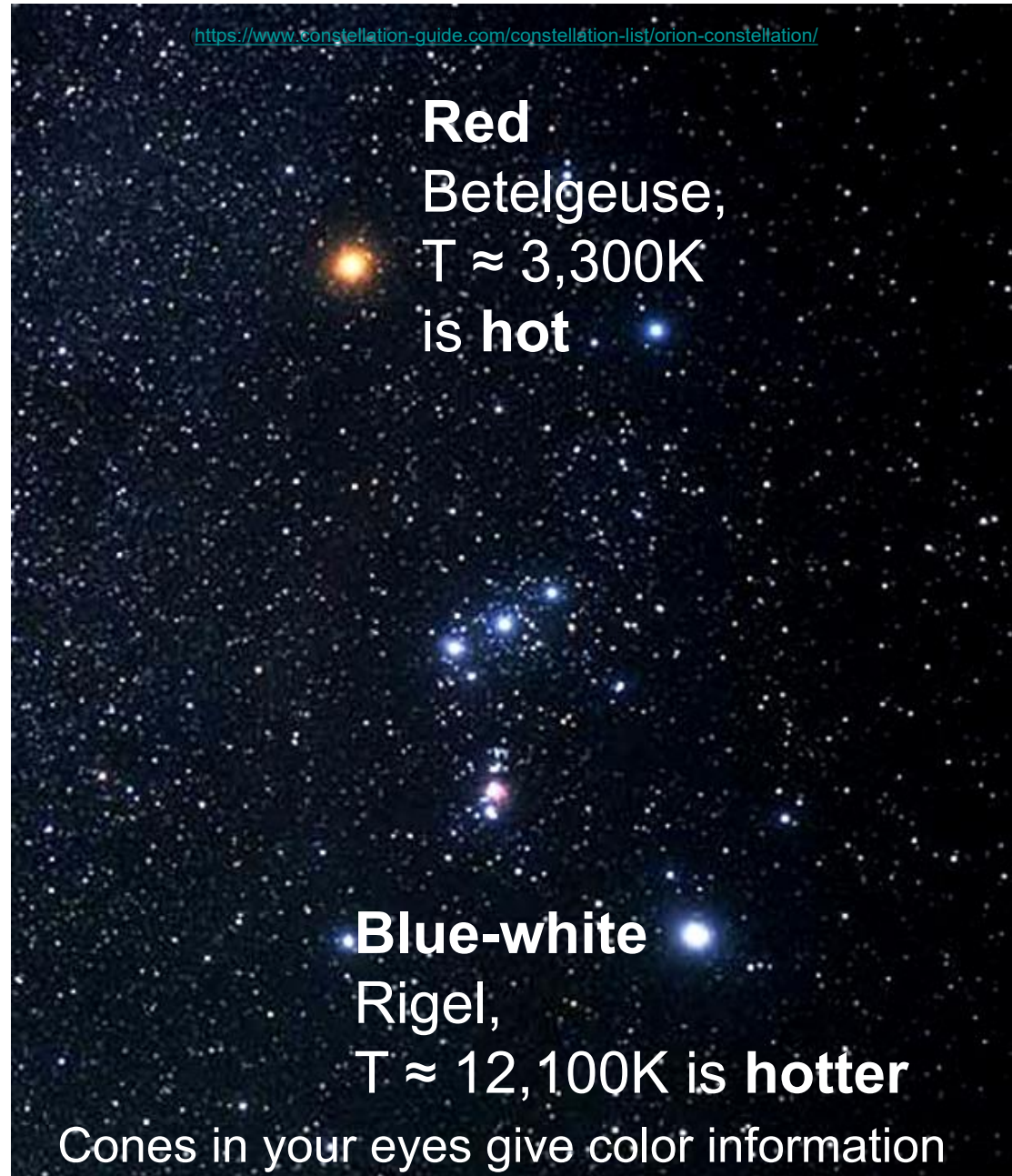
Decrease of  $\lambda_{\text{peak}}$   
with increase in  
temperature



**Very hot** universe  
meant **very short**  
EM radiation  
**wavelengths**

**Wilhelm Wien**, 1911 Nobel Prize in Physics

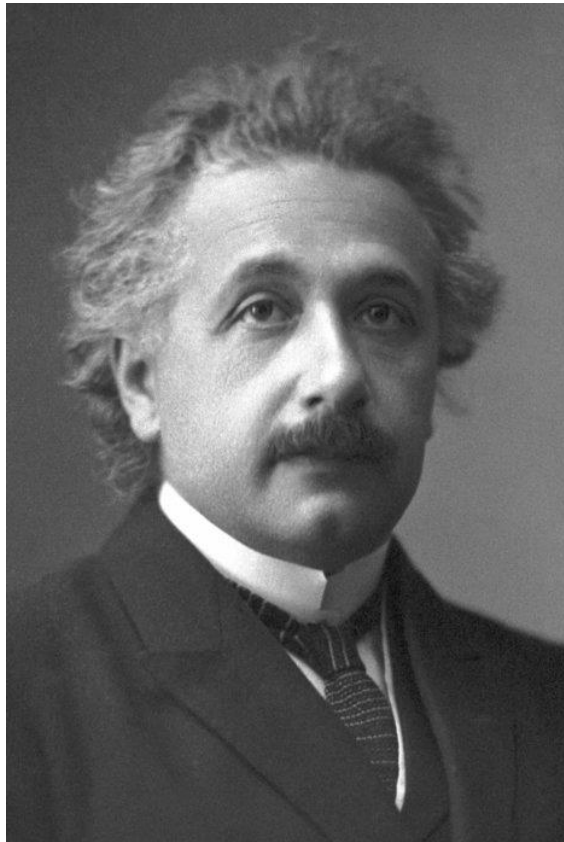
# You can experience this as the colors of bright stars





# EM radiation comes in packets called photons: crazy but true

Einstein realized that EM radiation has both wave characteristics (wavelength and frequency) and particle characteristics (well defined energy).



EM radiation are waves with e.g. well defined frequency, **f**, or equivalently wavelength,  $\lambda$ , and they come as **localized wave packets** called **photons**.

The energy of a photon depends on its wave property, frequency (or wavelength),  
 **$E_{\gamma} = h f = h c / \lambda$ .**

Thus the **shortest wavelength photons**, the gamma-rays, have the **highest energy**.

**Albert Einstein, 1921 Nobel Prize in Physics**

# Technology/medicine takes advantage of the EM spectrum

- **Gamma rays** with the shortest wavelengths have the **highest energies**.



(a) Mobile phone: radio waves



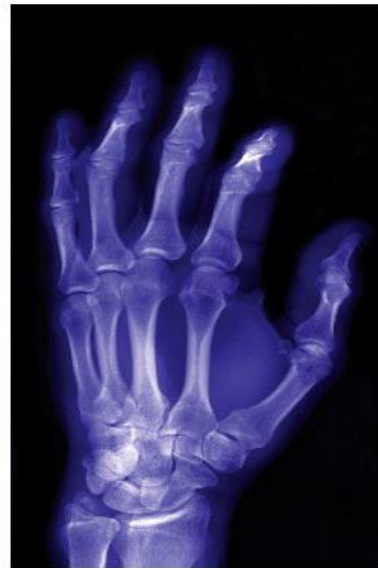
(b) Microwave oven: microwaves



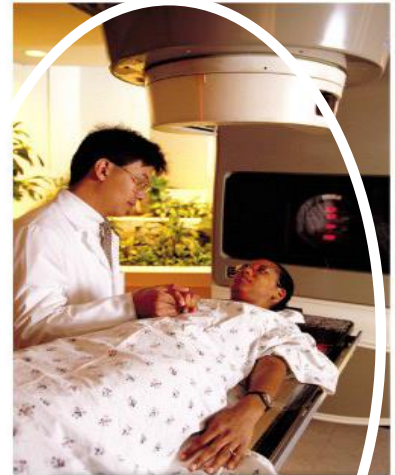
(c) TV remote: infrared light



(d) Tanning booth: ultraviolet light



(e) Medical imaging: X rays.



(f) Cancer radiotherapy: gamma rays

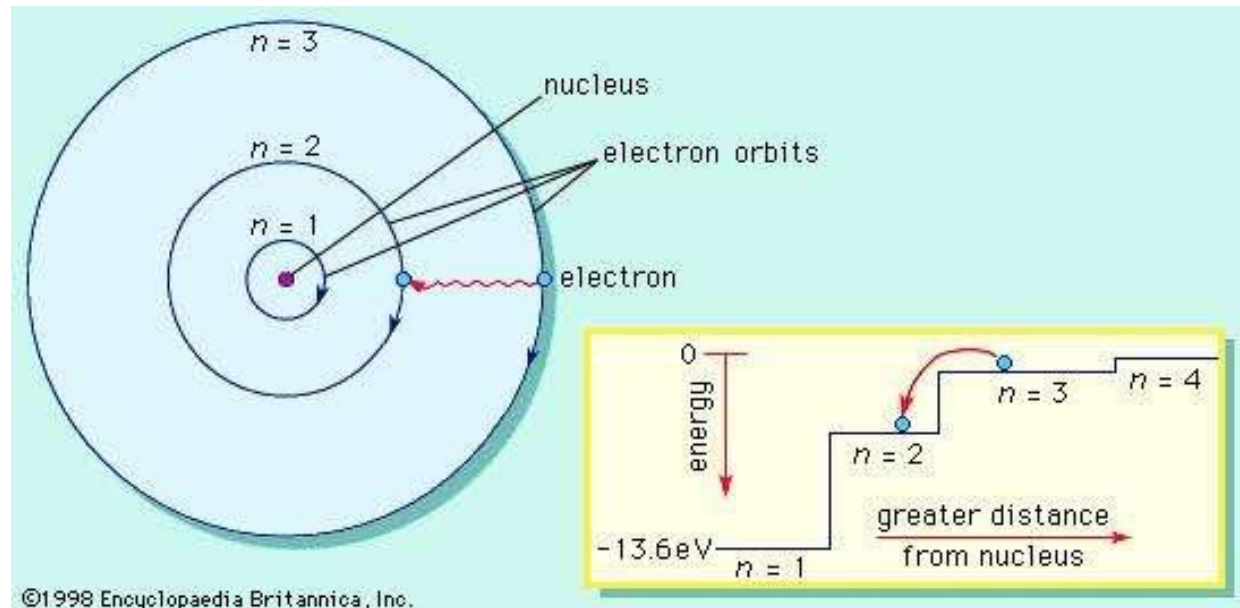


# Photons of just the right energy interact with atoms/nuclei

The **Bohr model** included two critical features: **quantized energy levels** and **emission/absorption** of **EM** radiation only during energy level transitions.



Schematic of hydrogen atom making a transition from a higher to lower energy level.

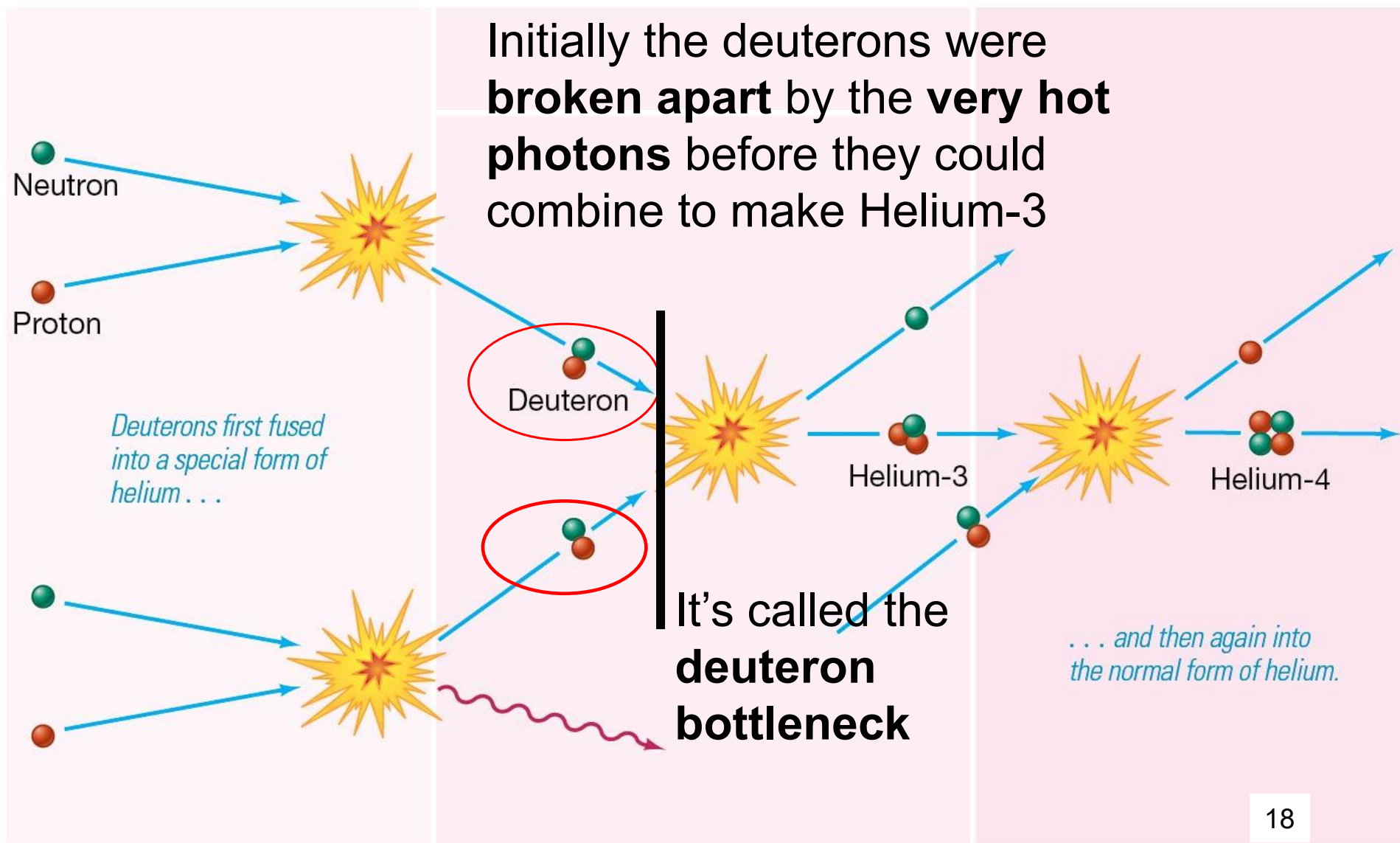


©1998 Encyclopaedia Britannica, Inc.

(<https://www.nobelprize.org/prizes/physics/1922/bohr/facts/> and <https://www.britannica.com/science/atom/Orbits-and-energy-levels> )

**Niels Bohr, 1922 Nobel Prize in Physics**

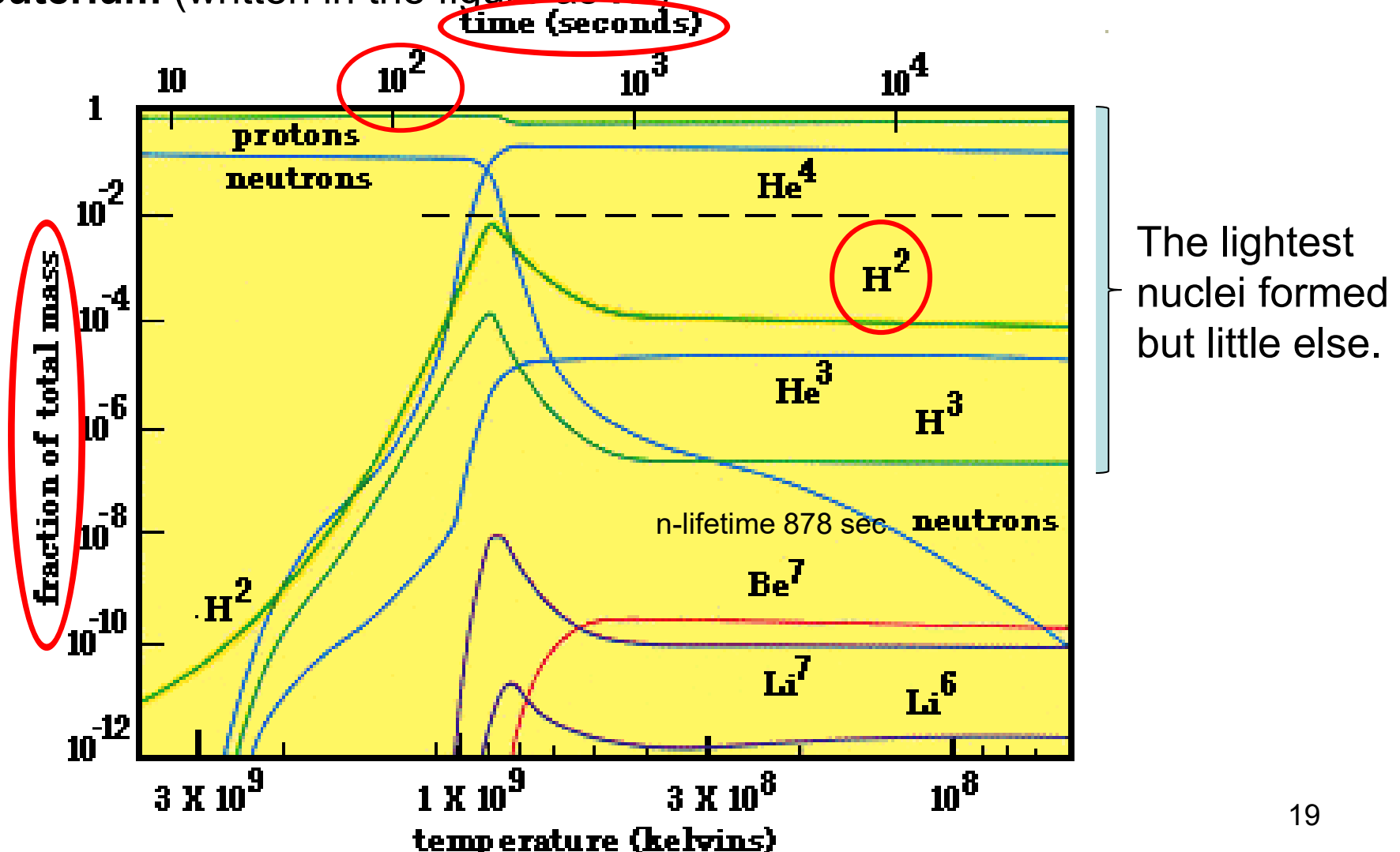
# Thanks to Bohr and Einstein, the deuterons were broken apart



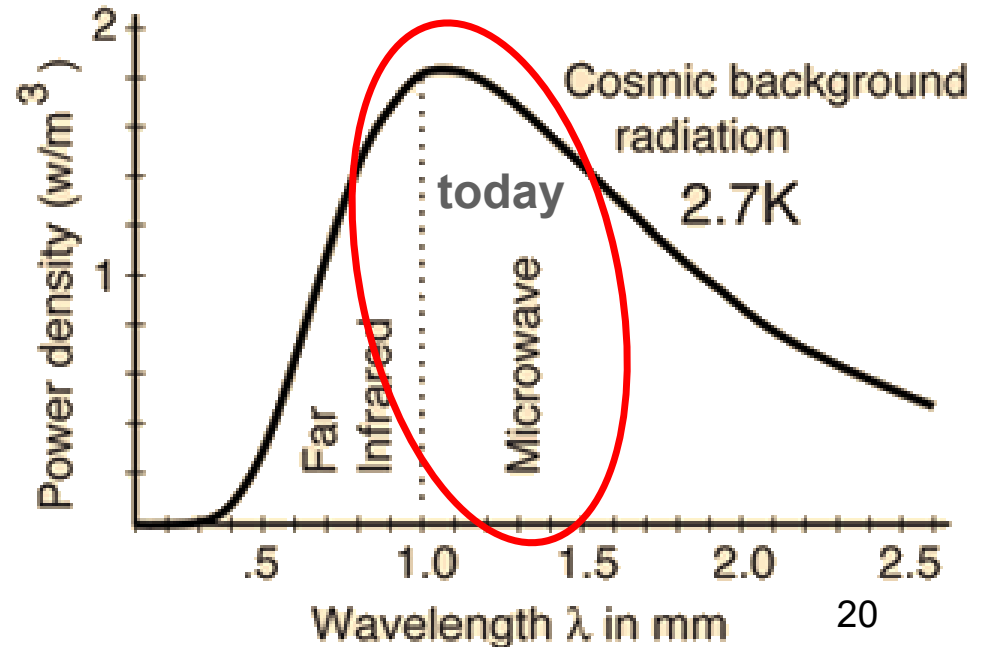
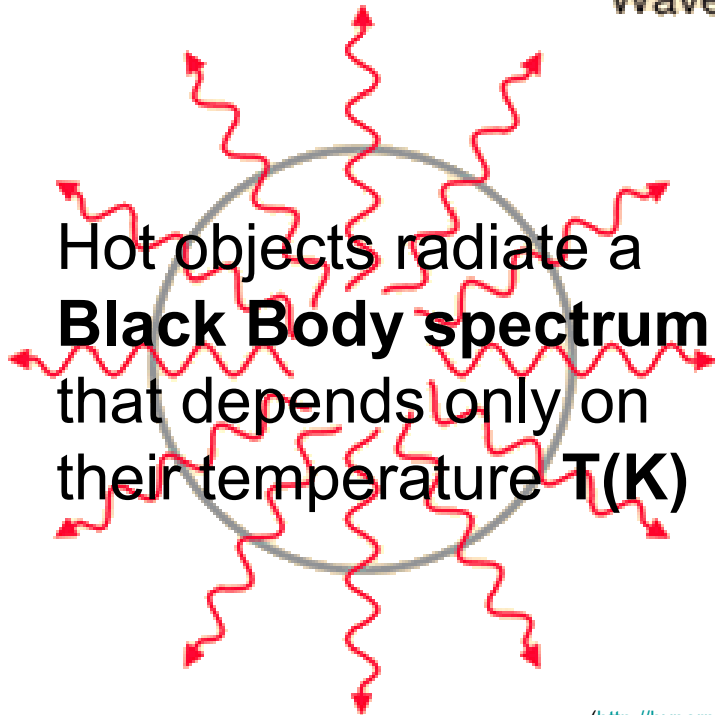
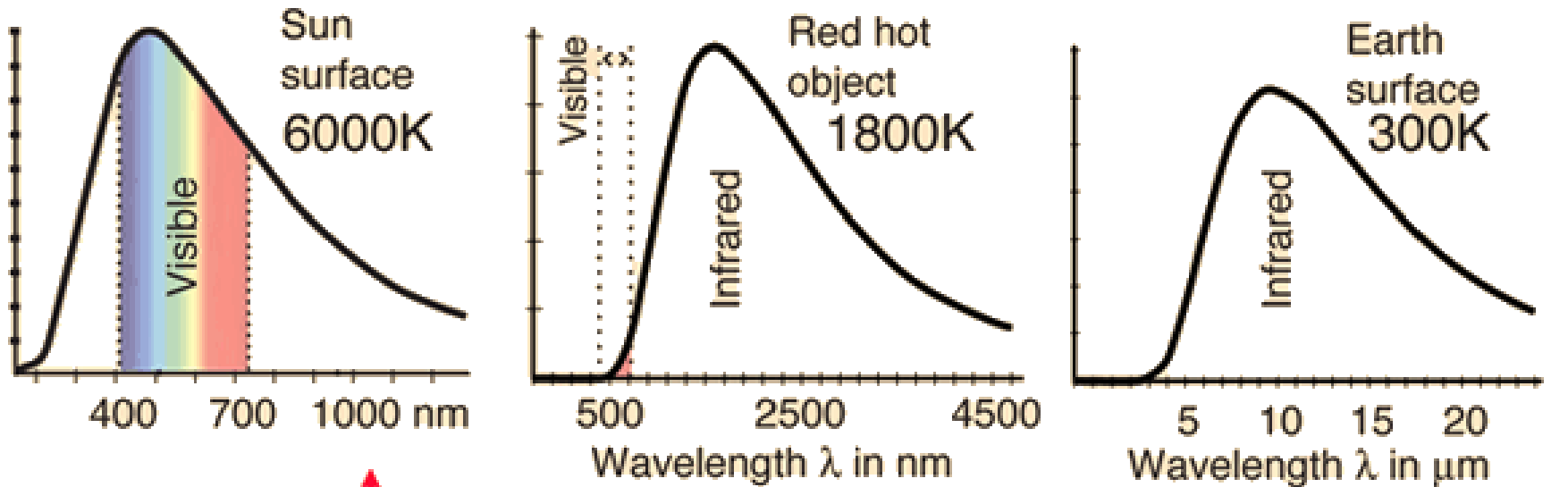


# Universe expansion cooled the universe “just in time”

Between  $10^2$  and  $4 \times 10^2$  seconds, the universe had cooled enough that the EM radiation photons were no longer energetic enough to dissociate the isotope **deuterium** (written in the figure as  $H^2$ )

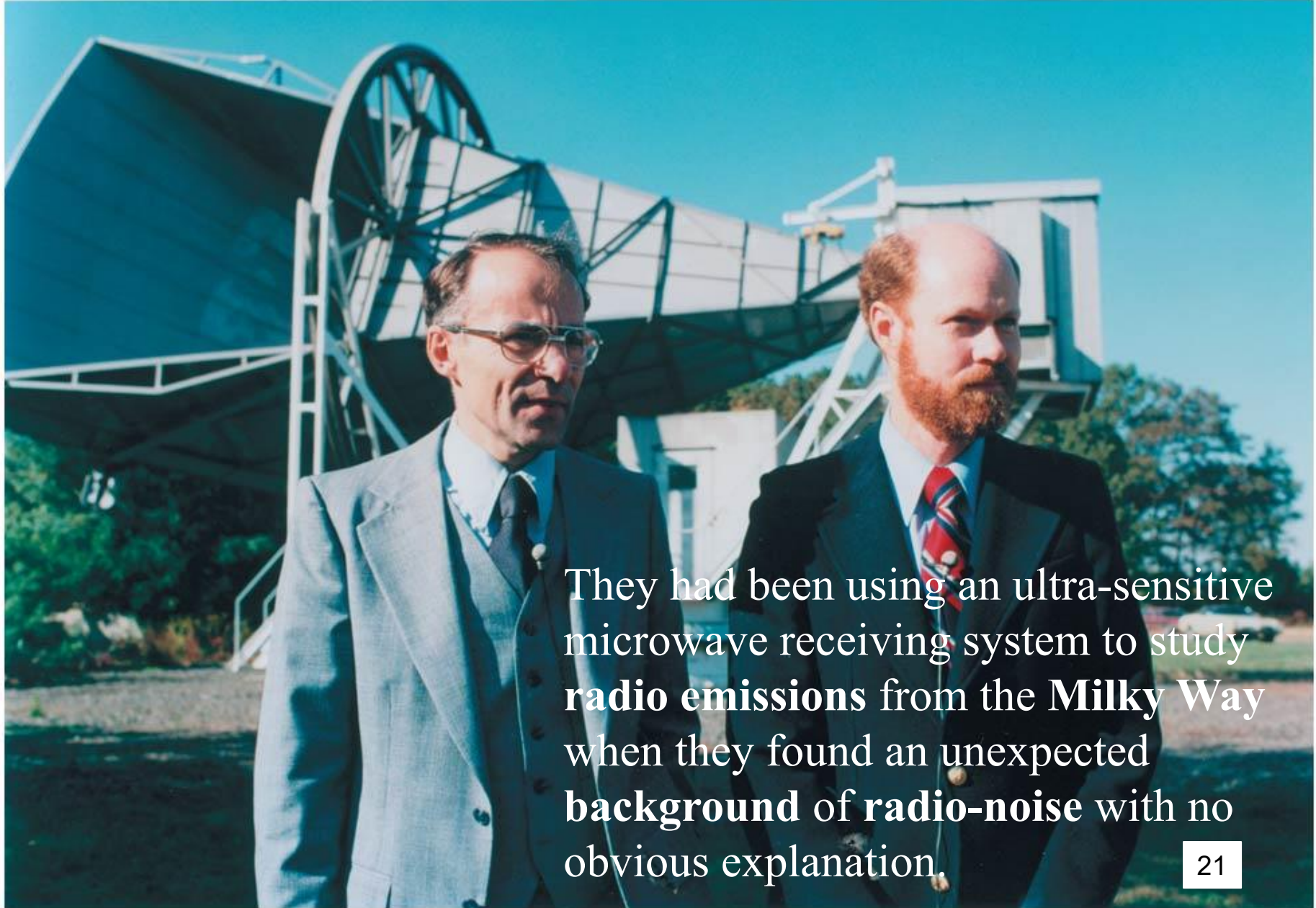


# Gamma rays were then stretched by the universe expansion





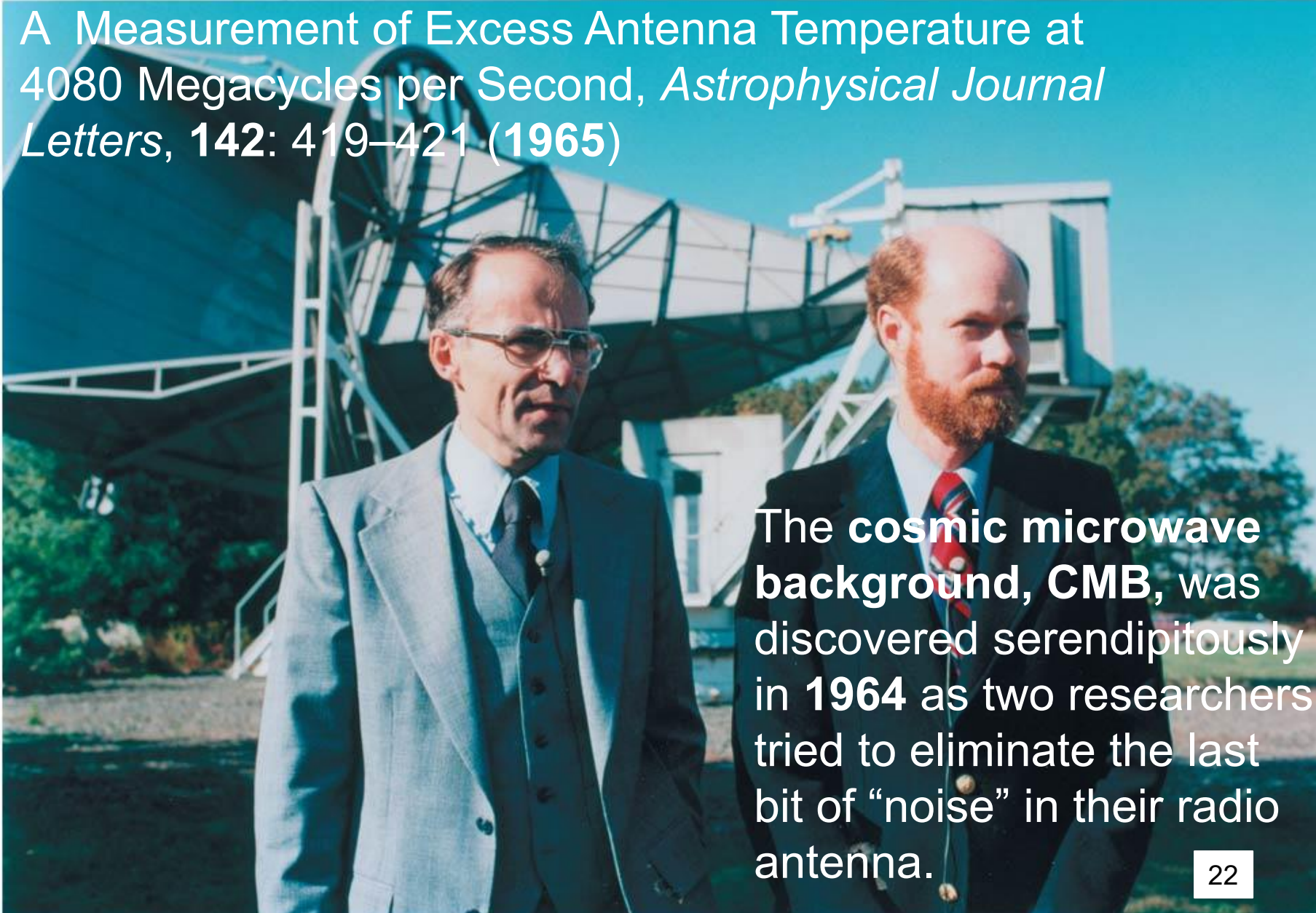
## The “cast off” radio antenna struggled with unexpected radio-noise



They had been using an ultra-sensitive microwave receiving system to study **radio emissions** from the **Milky Way** when they found an unexpected **background of radio-noise** with no obvious explanation.

# Antenna's radio-noise was the “huge flux” from the Big Bang

A Measurement of Excess Antenna Temperature at 4080 Megacycles per Second, *Astrophysical Journal Letters*, **142**: 419–421 (1965)

A color photograph of two men standing in front of a large, white, parabolic radio telescope dish. The man on the left is wearing a light blue suit, a white shirt, and a dark tie, and has glasses. The man on the right is wearing a dark suit, a white shirt, and a red and blue striped tie, and has a full red beard. The background shows the structure of the telescope and some greenery under a clear blue sky.

The **cosmic microwave background, CMB**, was discovered serendipitously in **1964** as two researchers tried to eliminate the last bit of “noise” in their radio antenna.



# CMB detection confirmed a Hot Big Bang prediction!

## The Nobel Prize in Physics 1978

(<https://www.nobelprize.org/prizes/physics/1978/summary/> )

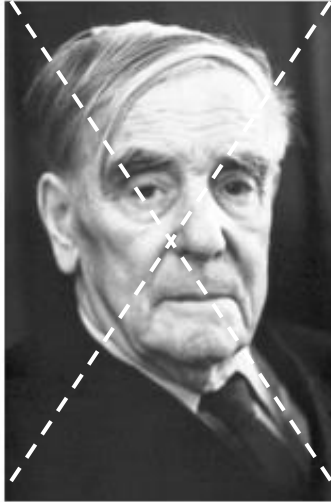


Photo from the Nobel  
Foundation archive.

**Pyotr Leonidovich  
Kapitsa**

Prize share: 1/2



Photo from the Nobel  
Foundation archive.

**Arno Allan Penzias**

Prize share: 1/4



Photo from the Nobel  
Foundation archive.

**Robert Woodrow  
Wilson**

Prize share: 1/4

The Nobel Prize in Physics 1978 was divided, one half awarded to Pyotr Leonidovich Kapitsa "for his basic inventions and discoveries in the area of low-temperature physics", the other half jointly to Arno Allan Penzias and Robert Woodrow Wilson "for their discovery of cosmic microwave background radiation"

# Curiously the 1960s had several serendipitous discoveries

1963 - Researchers realized that **Quasars** were at immense distances and must have an extreme energy source.

On May 20, 1964, American radio-astronomers Robert Wilson and Arno Penzias discovered the **cosmic microwave background** radiation (CMB)

1966 - Halton Arp published a "**catalog of peculiar galaxies**" (now known to be dominated by colliding galaxies).

1967 - **Pulsars** were discovered by Jocelyn Bell.

1968 - Thomas Gold proposed pulsars were spinning **neutron stars** which emit radiation similar to a rotating beacon.

1969~72 - **Sixteen gamma-ray bursts** of non-local origin were detected by the Vela military satellites. In 1997, i.e. 28 years post-discovery, gamma-ray bursts were found to be at huge distances requiring some extreme energy source.

# Precision radio-astronomy requires interferometers



After WW II, **Martin Ryle**, Antony Hewish and colleagues of the Cambridge Astrophysics Group developed a series of radio interferometers. Interferometers combine the signals from multiple telescopes.

(<https://physicsworld.com/a/martin-ryle-an-energy-visionary/>)



# Radio-astronomy benefitted from WW II radar instrumentation

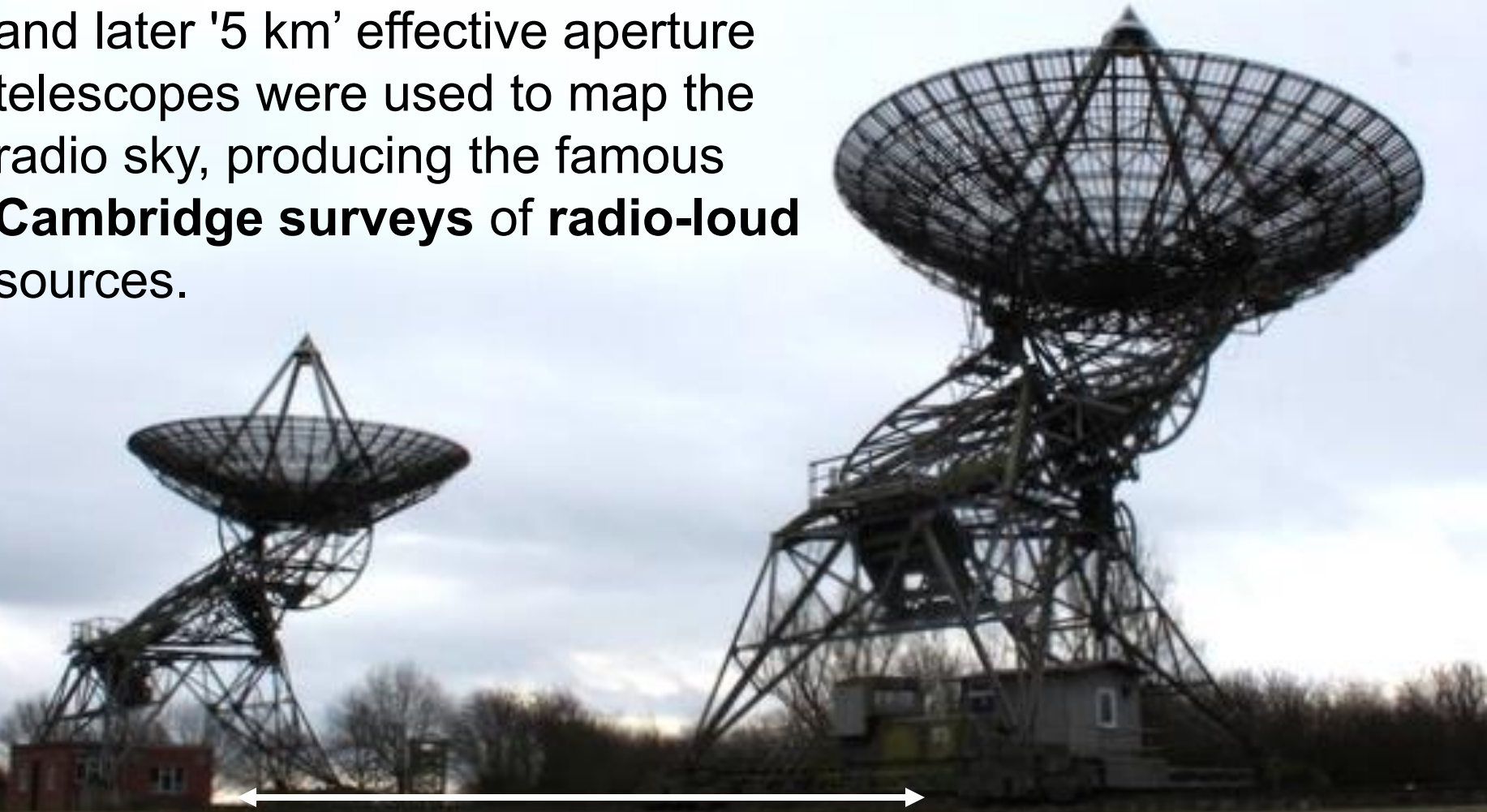
## Mullard EL34 Power Pentode



- Mullard produced 40% of vacuum tubes used by Britain during WW II.
- 1955 - Mullard Company donated £100,000 for an additional radio telescope and associated facilities.
- **Mullard Observatory** opened on July 27, 1957.

# Cambridge 2C and 3C radio surveys

By the late 1950s the 'One-Mile' and later '5 km' effective aperture telescopes were used to map the radio sky, producing the famous **Cambridge surveys of radio-loud sources**.



Telescope separation is pivotal: 1 mile or 5 km

# What were those “radio-loud” sources?



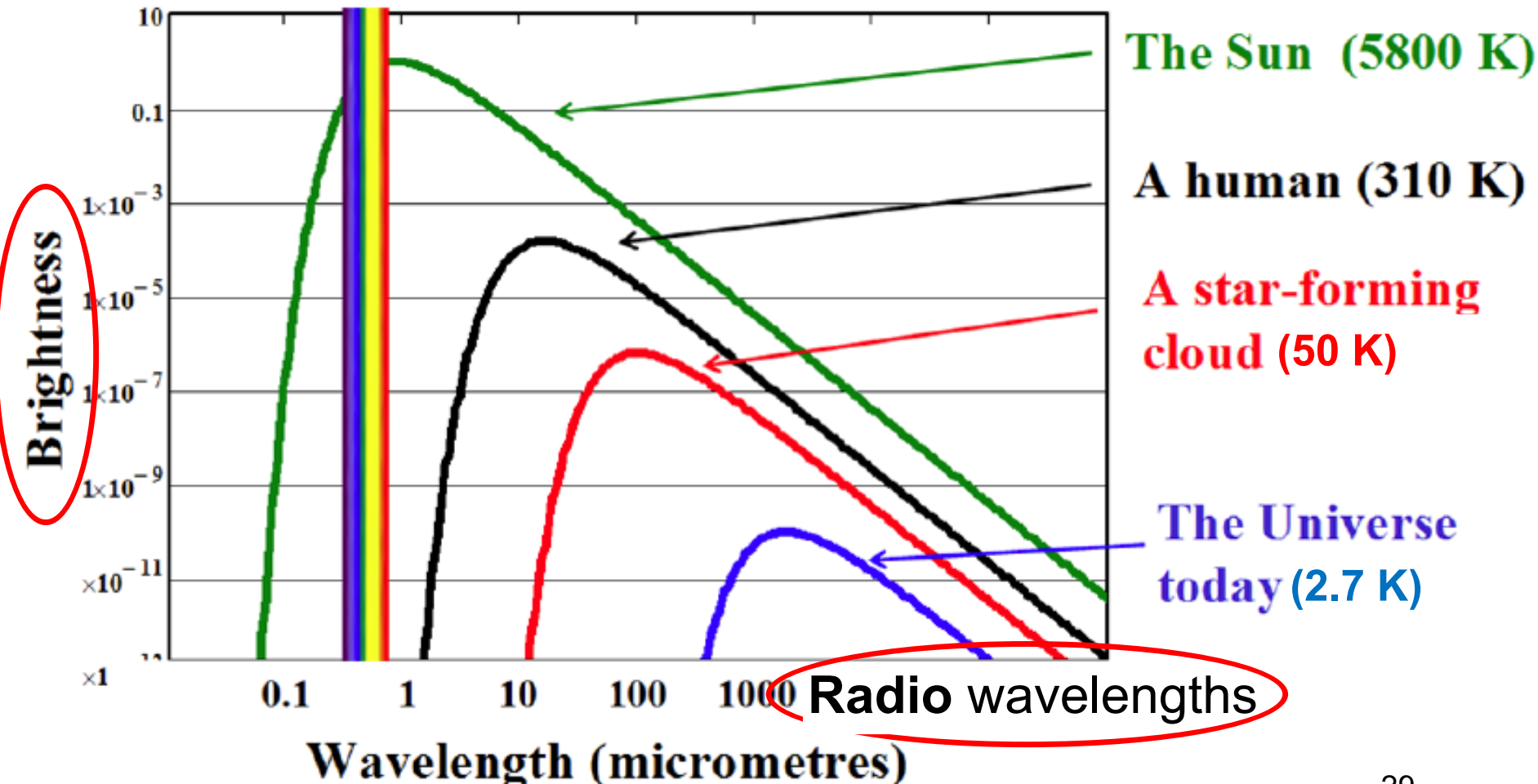
**Radio-loud sources** – are those with high luminosity at **radio wavelengths**.

**Stars** – which are bright at much shorter wavelengths, are **radio-quiet**.



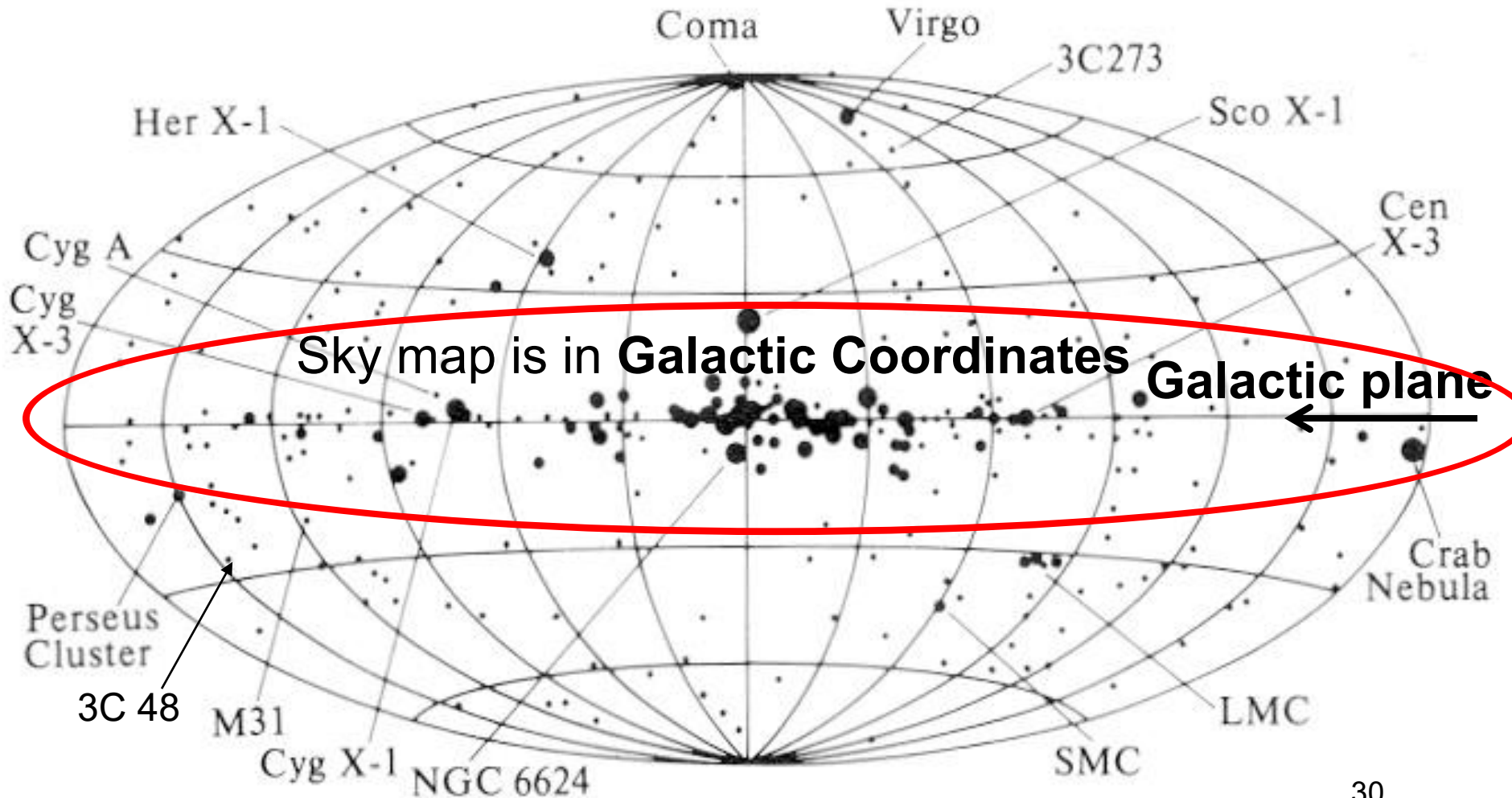
# Conventional sources are not radio-loud

**Normal** sources of electro-magnetic radiation have almost no emission (brightness) at **radio wavelengths**!  
So **radio-loud** is totally **unexpected**.

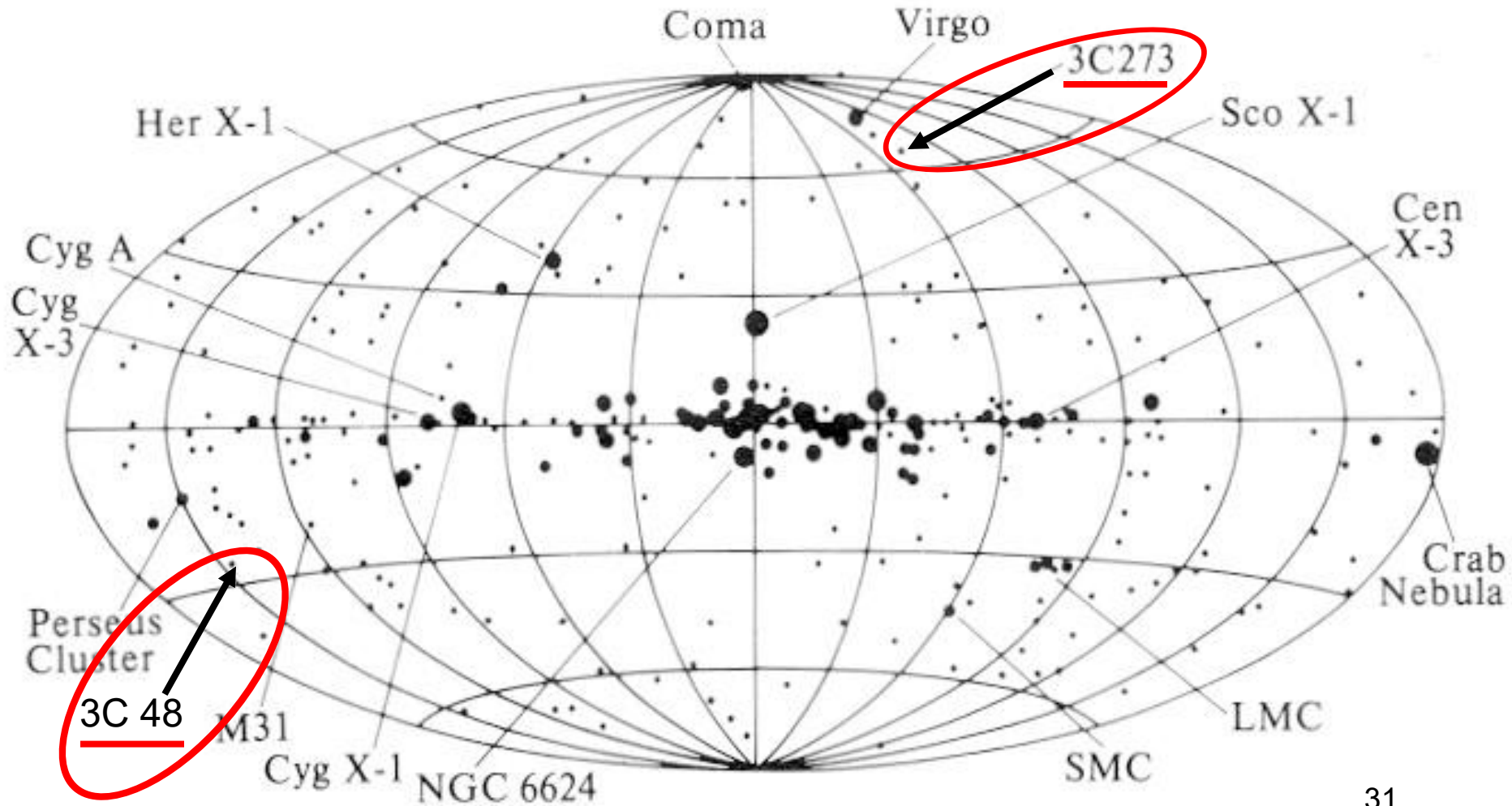


# Many of the radio sources were likely within the Milky Way

The Cambridge **2C** and **3C** surveys of **radio-loud sources** resulted in a map, the **black dots**, on the sky like this:



# Particularly unexpected were the likely extra-galactic sources





# The hunt was on for optical counterparts

**Optical counterparts**  
of the Cambridge 3C  
**catalog sources**  
should help unravel the  
*physics* of these radio-  
loud objects.

Large light gathering  
power of the Hale  
200" telescope at the  
Mount Palomar  
Observatory was  
essential for **spectral  
analysis!** [More on  
spectra soon] Source  
spectra can tell us the  
distance to the  
sources.

# 1<sup>st</sup> breakthrough was from improved radio source location

1960 - a star-like object was associated with the radio source 3C 48 by Thomas **Matthews** and analyzed by Allan **Sandage**.

A modern radio telescope image of 3C 48

Called a ***point source*** because of its small size.  
Green arrow is **0.5 arcsec**

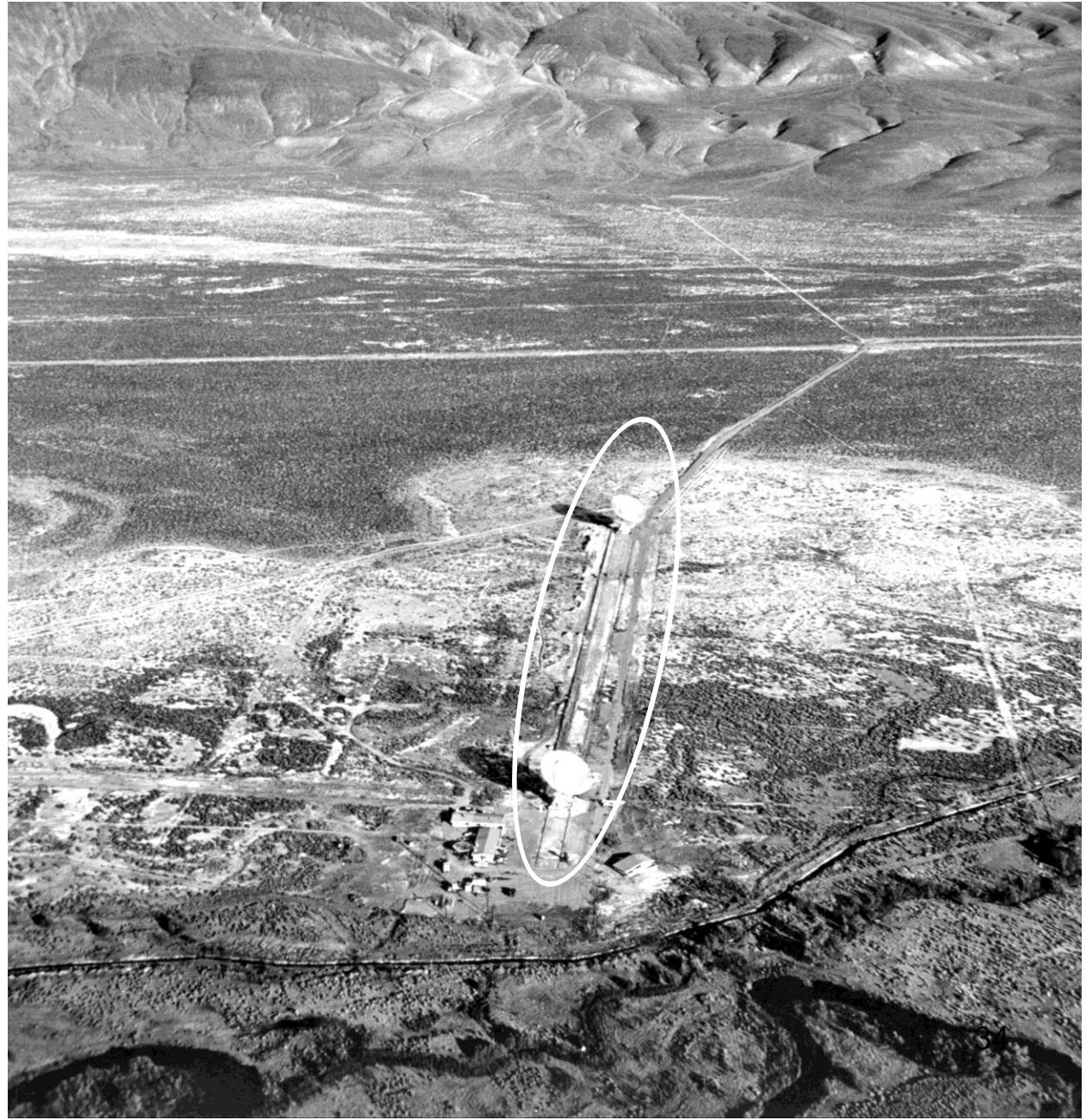
([https://en.wikipedia.org/wiki/3C\\_48](https://en.wikipedia.org/wiki/3C_48))



# The position uncertainty of 3C 48 needed to be reduced

To link the **3C 48** radio source with an optical object required improved **radio-position** information.

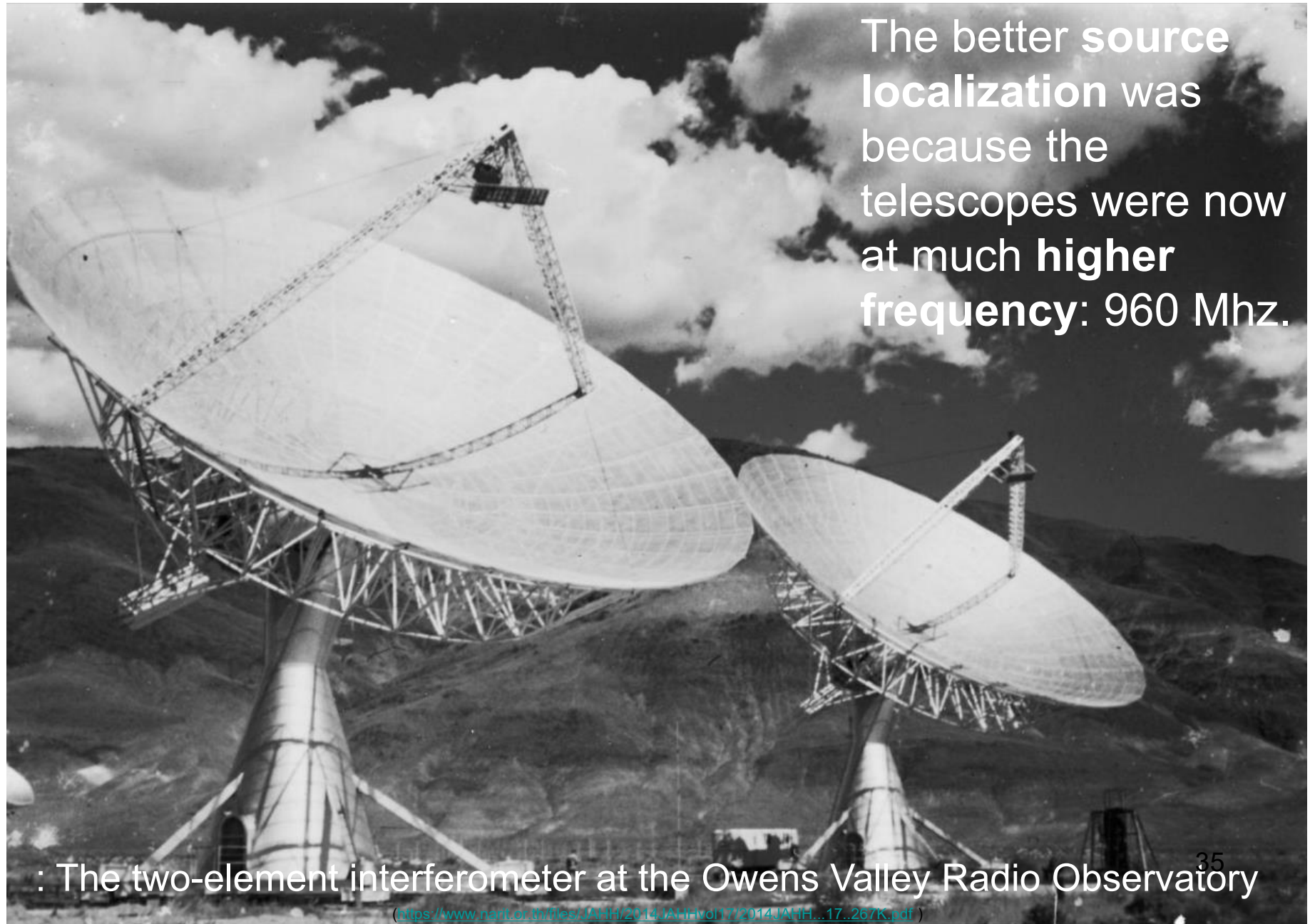
This was **Matthews'** roll using the Owen's Valley Radio Observatory interferometer.





# Improved source localization was the game changer

The better **source localization** was because the telescopes were now at much **higher frequency**: 960 Mhz.



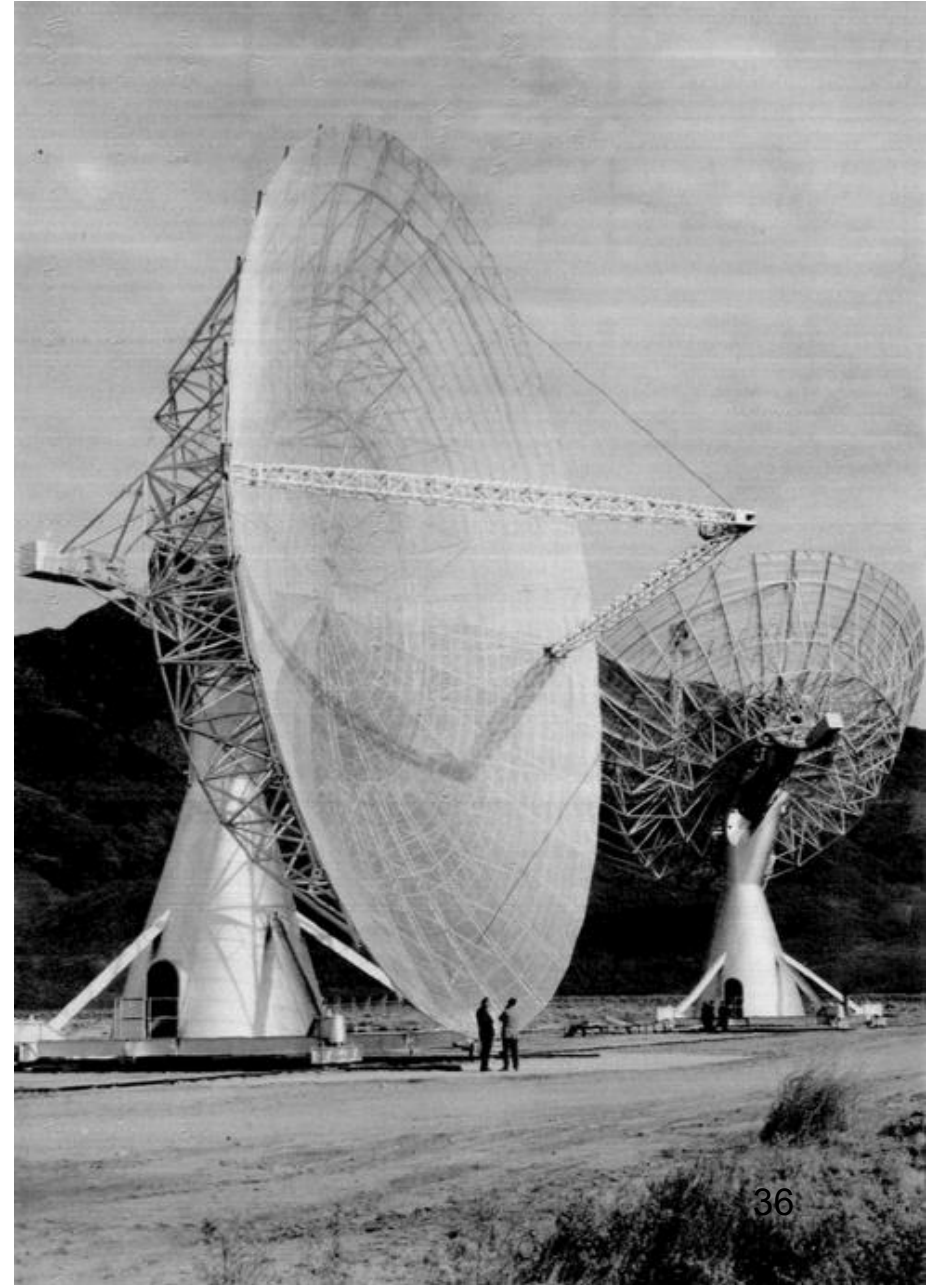
: The two-element interferometer at the Owens Valley Radio Observatory

# High frequency means shorter wavelengths critical for pointing

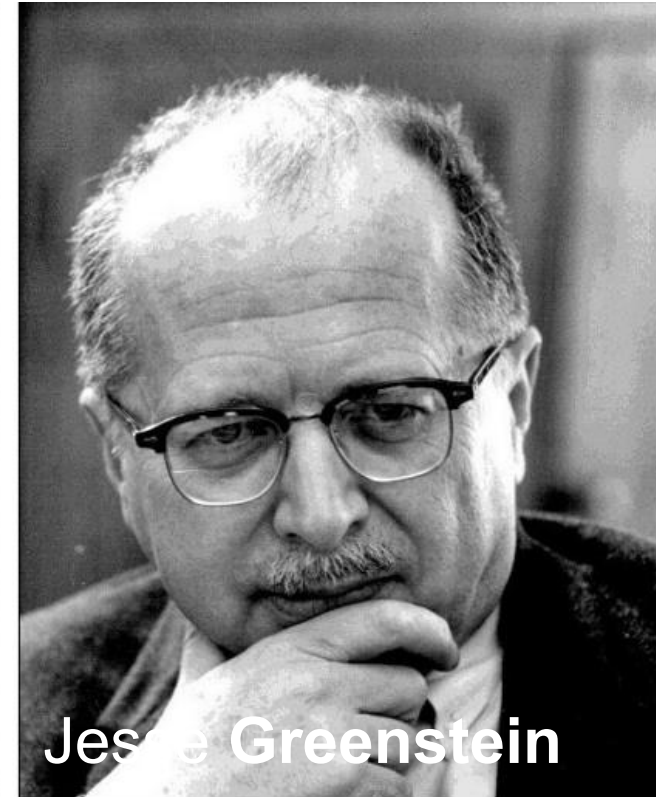
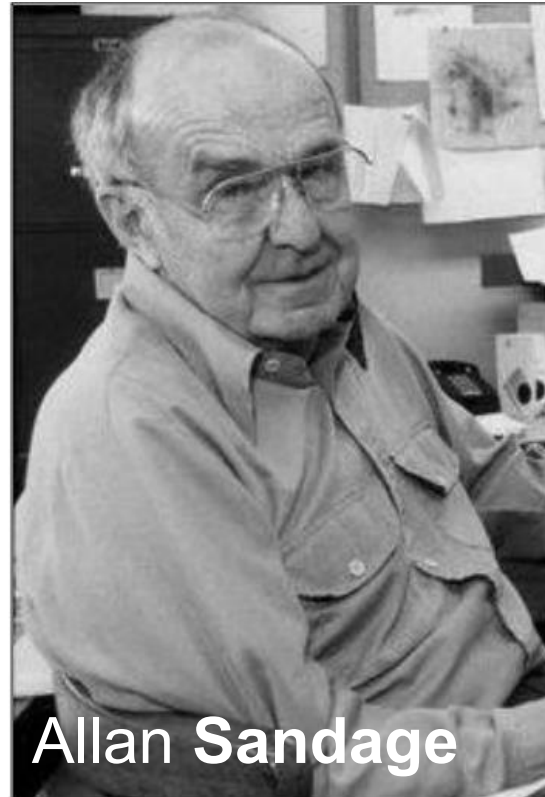
Owen's  
Valley Radio  
Observatory  
radio source  
positions  
were  
accurate  
up to 10-  
seconds  
of arc (**Jesse  
Greenstein**  
1961)



**Jesse Greenstein, here in 1957 with a more famous Porter drawing, was one of the prime movers behind the establishment of radio astronomy at Caltech.**



Matthews identified 3C 48 with what appeared to be a 16th magnitude star.



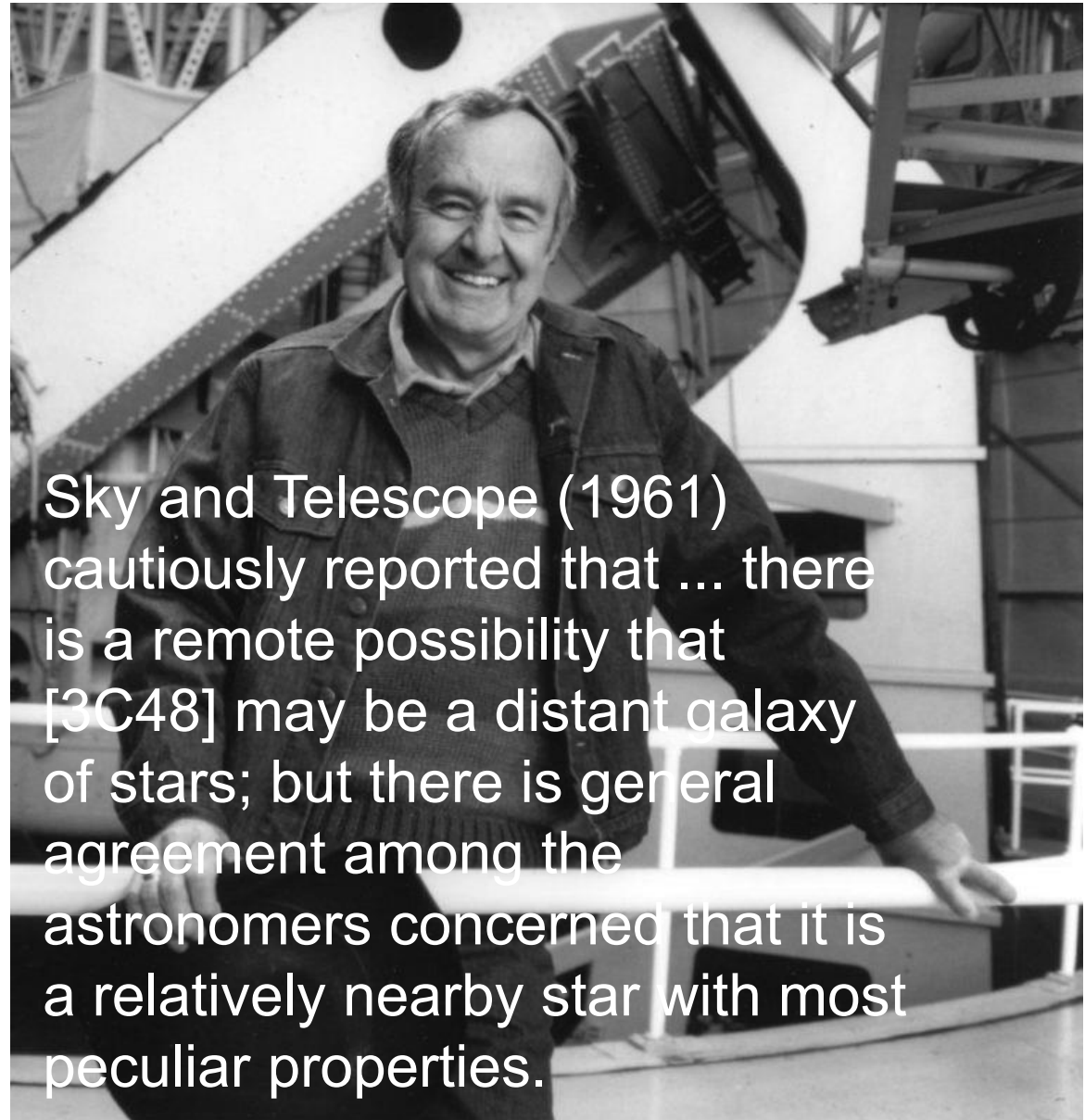
Fortunately a few years earlier, Jesse **Greenstein** had revolutionized the measurement of radio source positions with the **Owen's Valley Radio Observatory**.



# Frustratingly having a spectrum of the source didn't help

**Sandage's** roll was to obtain a spectrum with the state-of-the-art 200" Hale telescope.

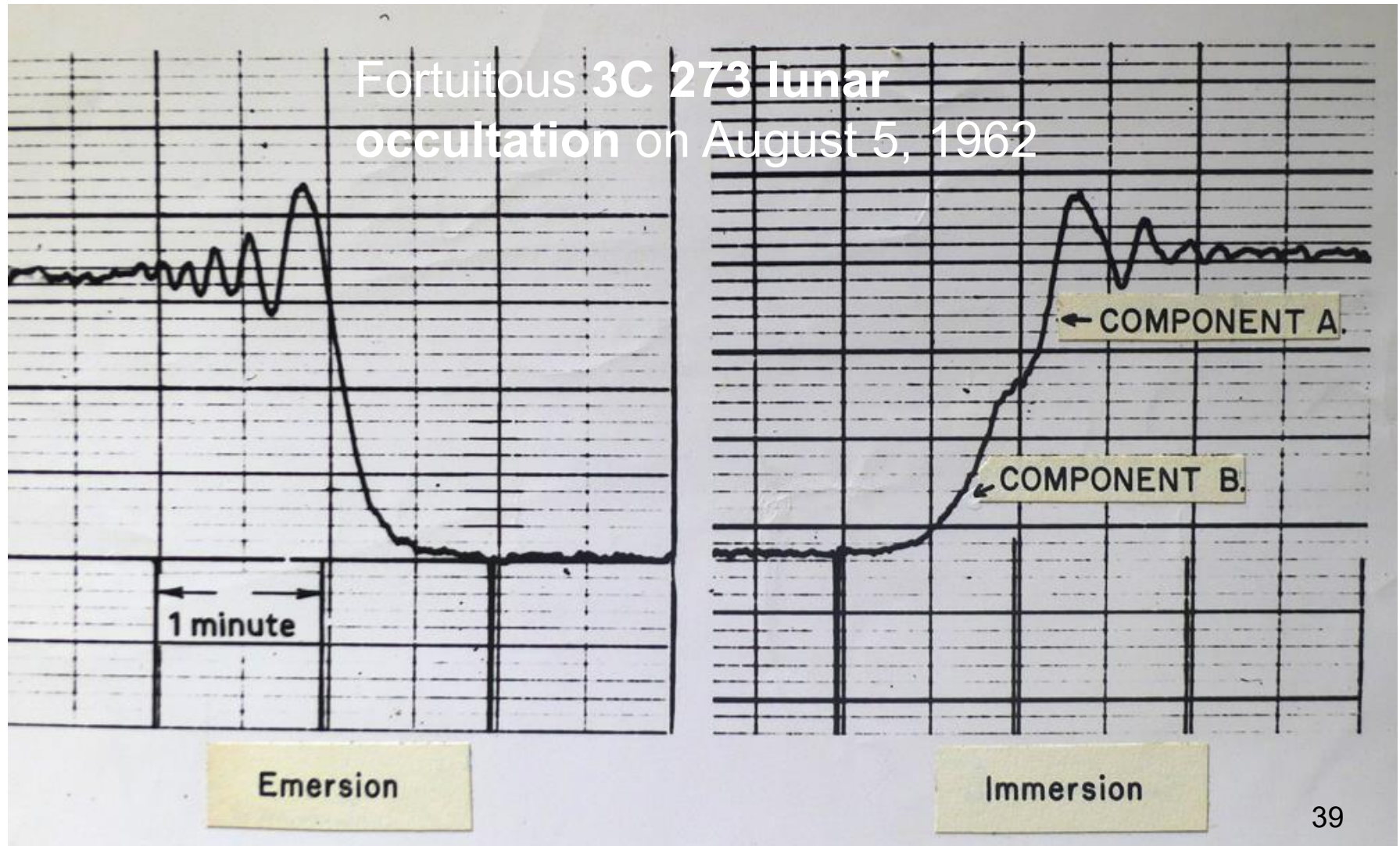
However the **spectrum** of the faint blue star contained many **unknown broad emission lines**.



Sky and Telescope (1961) cautiously reported that ... there is a remote possibility that [3C48] may be a distant galaxy of stars; but there is general agreement among the astronomers concerned that it is a relatively nearby star with most peculiar properties.

## 2<sup>nd</sup> breakthrough was from a serendipitous source occultation

In 1962/3 the location of another radio source, **3C 273**, was pinpointed using the Parkes Radio Telescope.



# The Moon's motion determined the position of 3C 273

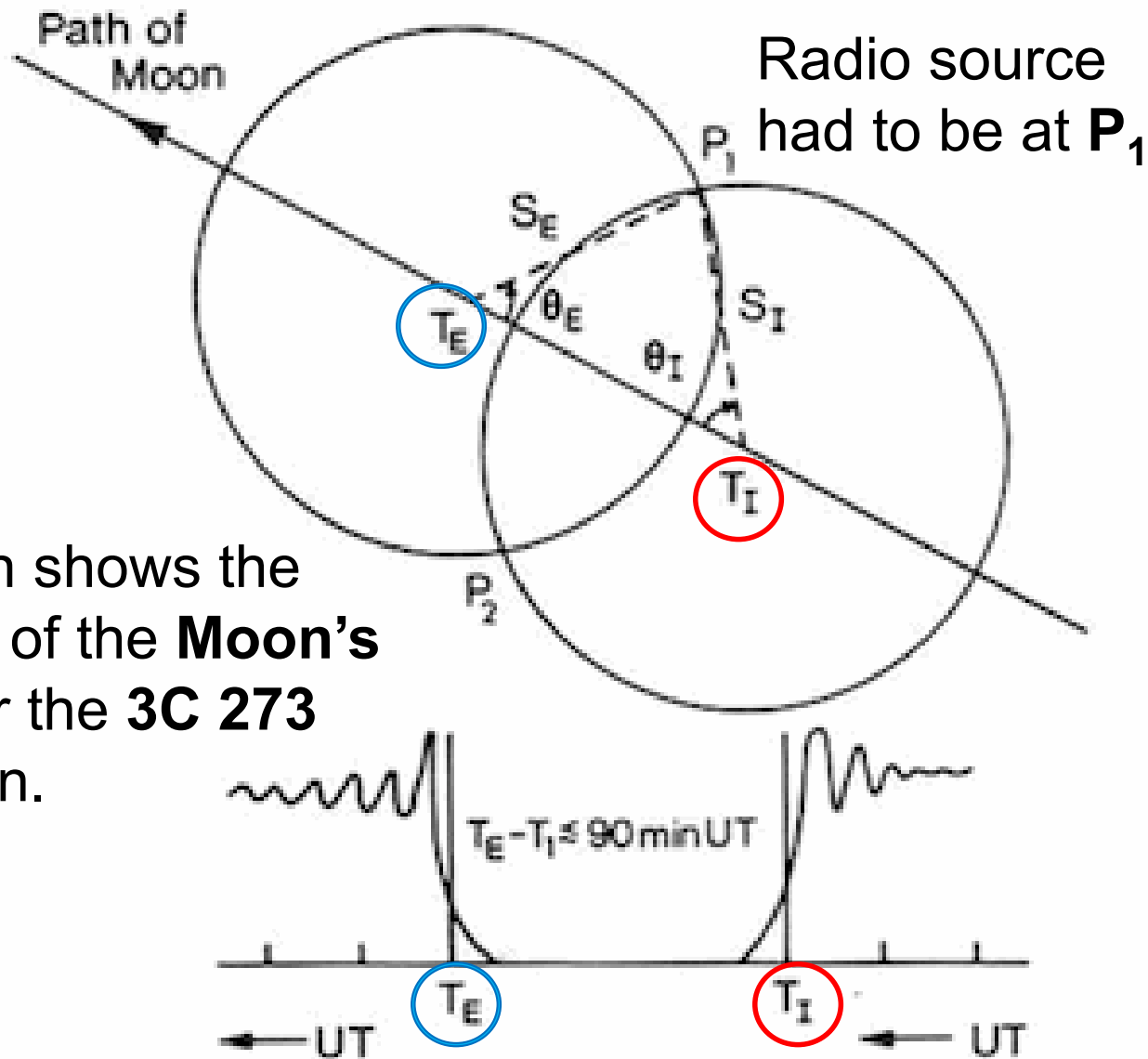


Illustration shows the geometry of the **Moon's** motion for the **3C 273** occultation.



## Sometimes being second is an advantage

**3C 273** was then associated with its optical counterpart, and **Maarten Schmidt** obtained an optical spectrum also using the 200" Hale Telescope.

This spectrum revealed the same ***strange emission lines*** as **3C 48**.



# Spectra are like *fingerprints* uniquely identifying each atom

Atoms (molecules) emit/absorb light only at certain **wavelengths** which are **distinctive** to each **atom**.

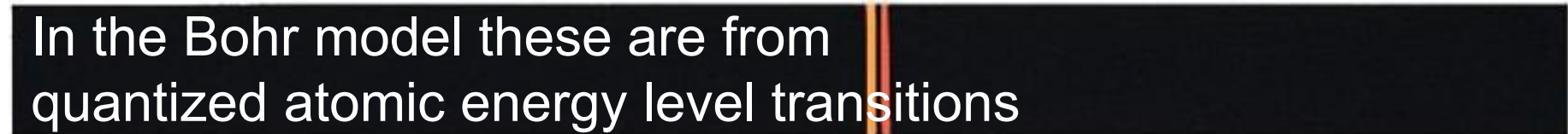
HYDROGEN



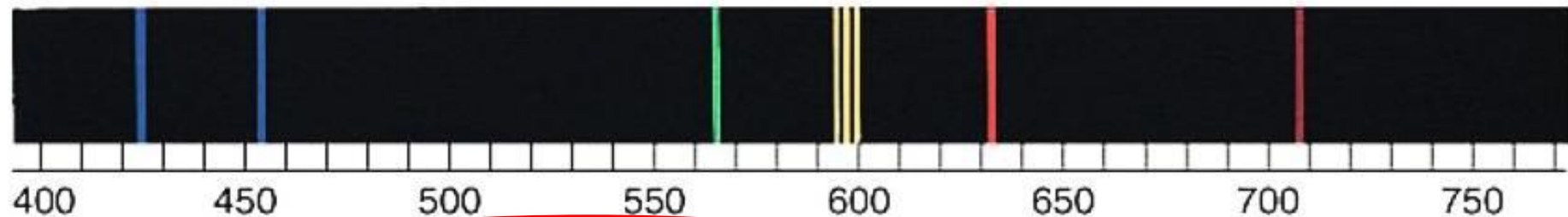
HELIUM



SODIUM

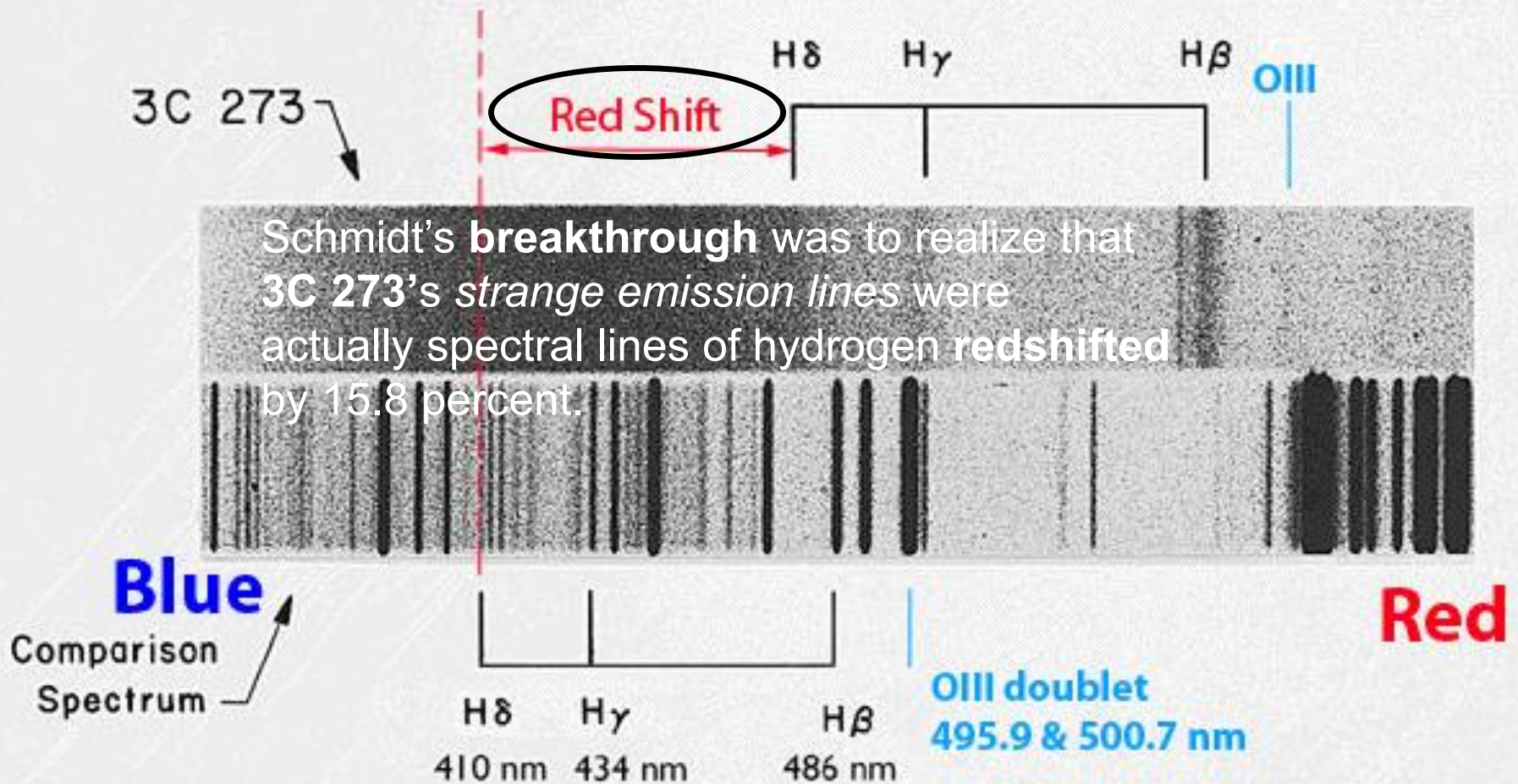


MERCURY



Wavelength of light ( Nanometers )

# 3C 273 *strange spectrum* was just an unexpected redshift





## HST image of 3C 273



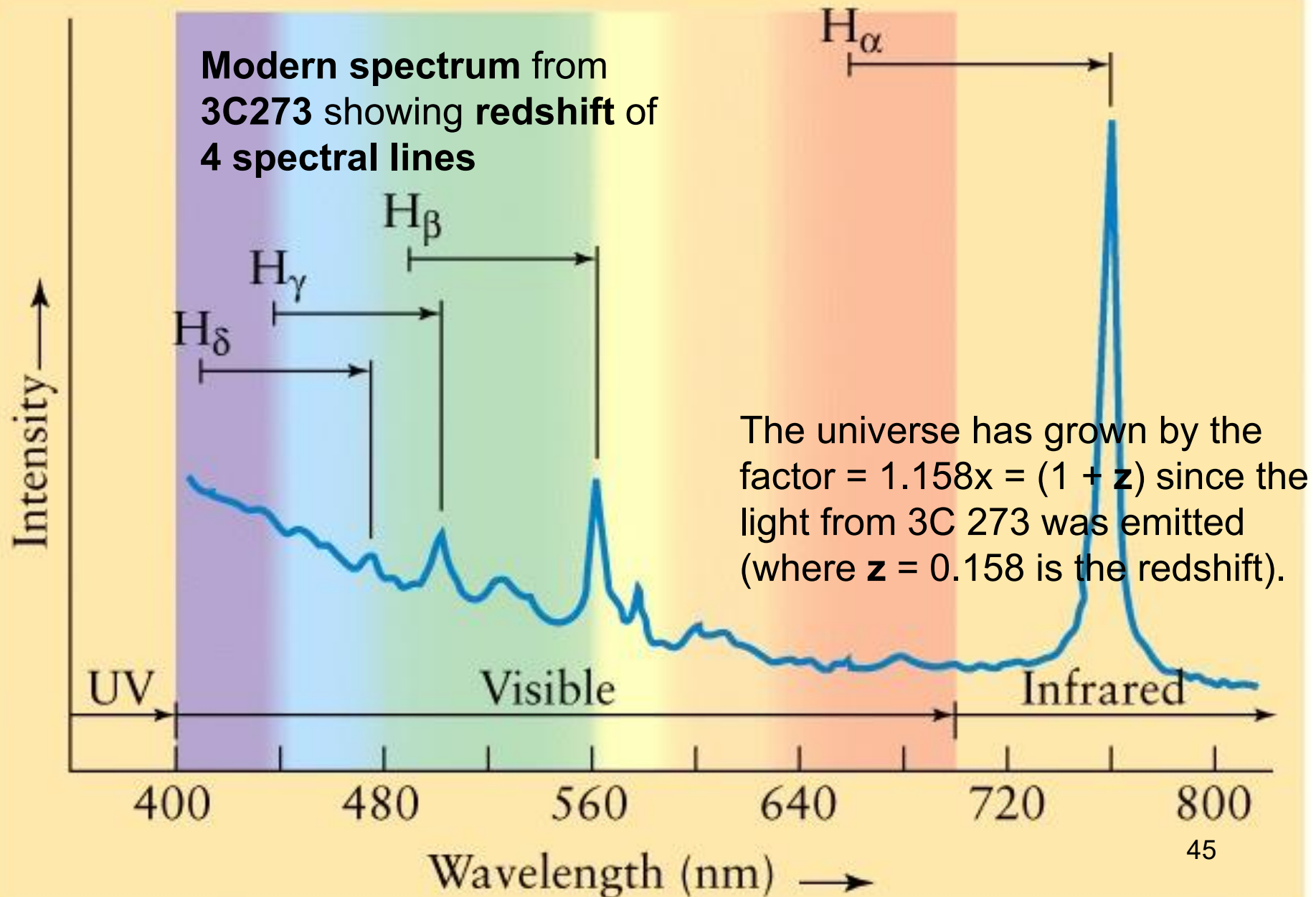
**Maarten Schmidt**, Nature **197**, 1040 (1963)

Schmidt also noted: “the close correlation between the radio structure (i.e. 3C 273) and the **star** (i.e. the **Quasar**) with the **jet** is suggestive and intriguing.”

We know **today** that 3C 273 is a **supermassive black hole** and associated **jets** are commonplace!

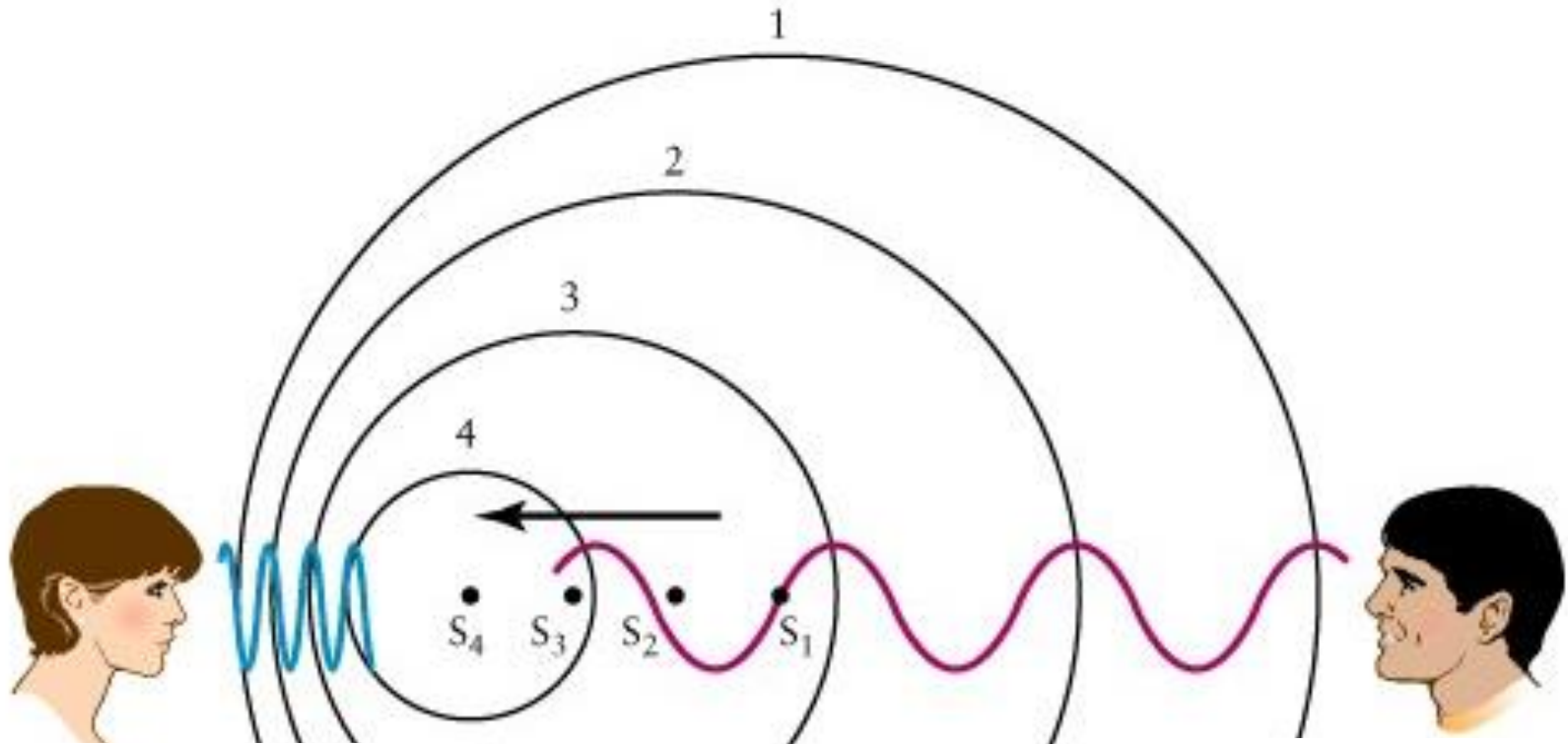
[https://en.wikipedia.org/wiki/3C\\_273](https://en.wikipedia.org/wiki/3C_273)

# Unexpected redshift was from the expansion of the universe!



# For waves, source motion changes what you observe

You may know this as a **Doppler shift**



When the source **approaches**, the observer sees a **wavelength shorter** than the source wavelength. This is labelled a **blueshift**.

When the source **recedes**, the observer sees a **wavelength longer** than the source wavelength. This is labelled a **redshift**.



# Term “redshift” is based on observed VS source wavelength

The terms **blueshift/redshift** are used instead of saying shorter or longer wavelengths than at the source.

Interestingly redshifts dominate, and it is defined as:

$$\text{redshift} = \mathbf{z} \equiv (\lambda_{\text{observed}} - \lambda_{\text{rest}})/\lambda_{\text{rest}}$$

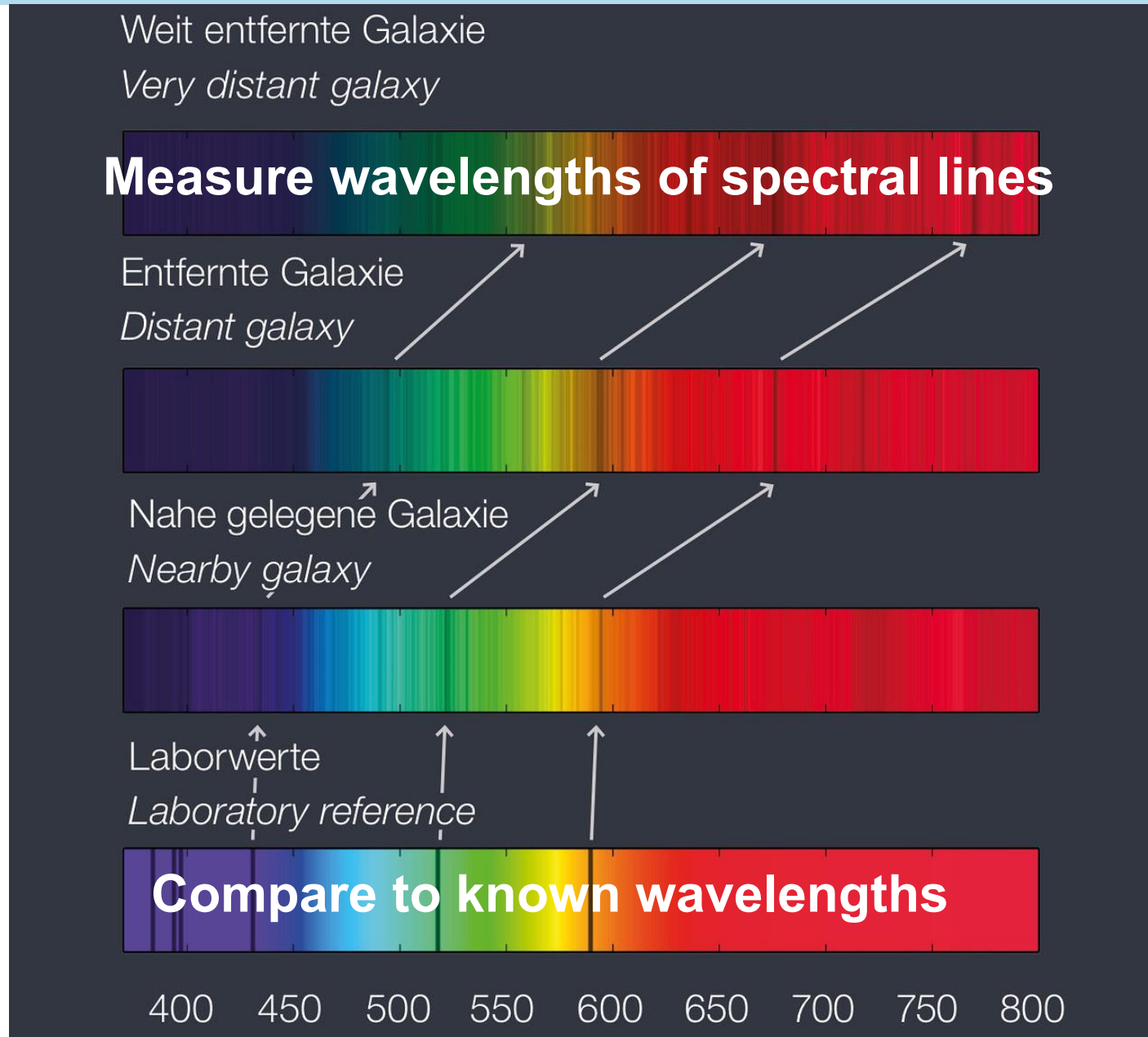
$\lambda_{\text{observed}}$  = observed **wavelength** (of light)

$\lambda_{\text{rest}}$  = **wavelength** (of light) at the source

For sources that are **receding** from us with velocity,  $v_{\text{recession}}$ , much less than the speed of light,  $c$ :

$$\mathbf{z} = v_{\text{recession}}/c$$

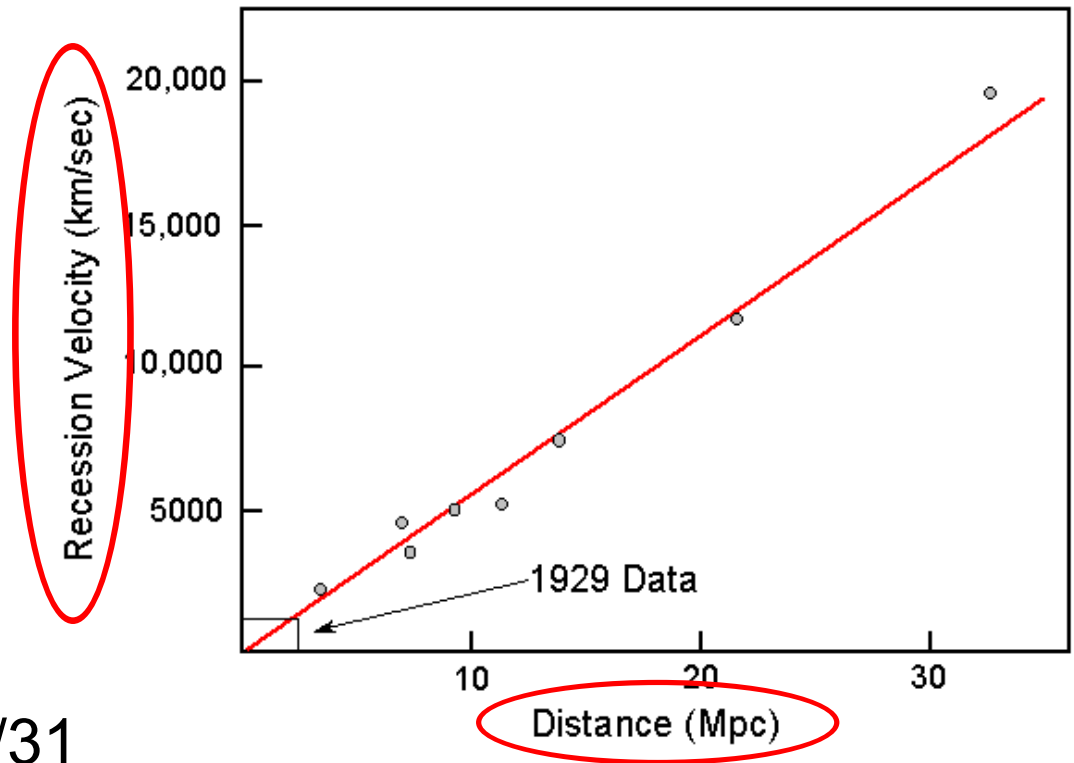
# How do astronomers determine redshift?



# Hubble observed that redshift $\propto$ to source distance

We now measure the distance to the most remote stars/galaxies with **Hubble's Law** (graph). Hubble saw that source **Recession Velocity** (speed) is highly correlated with source **Distance**.

## Hubble & Humason (1931)

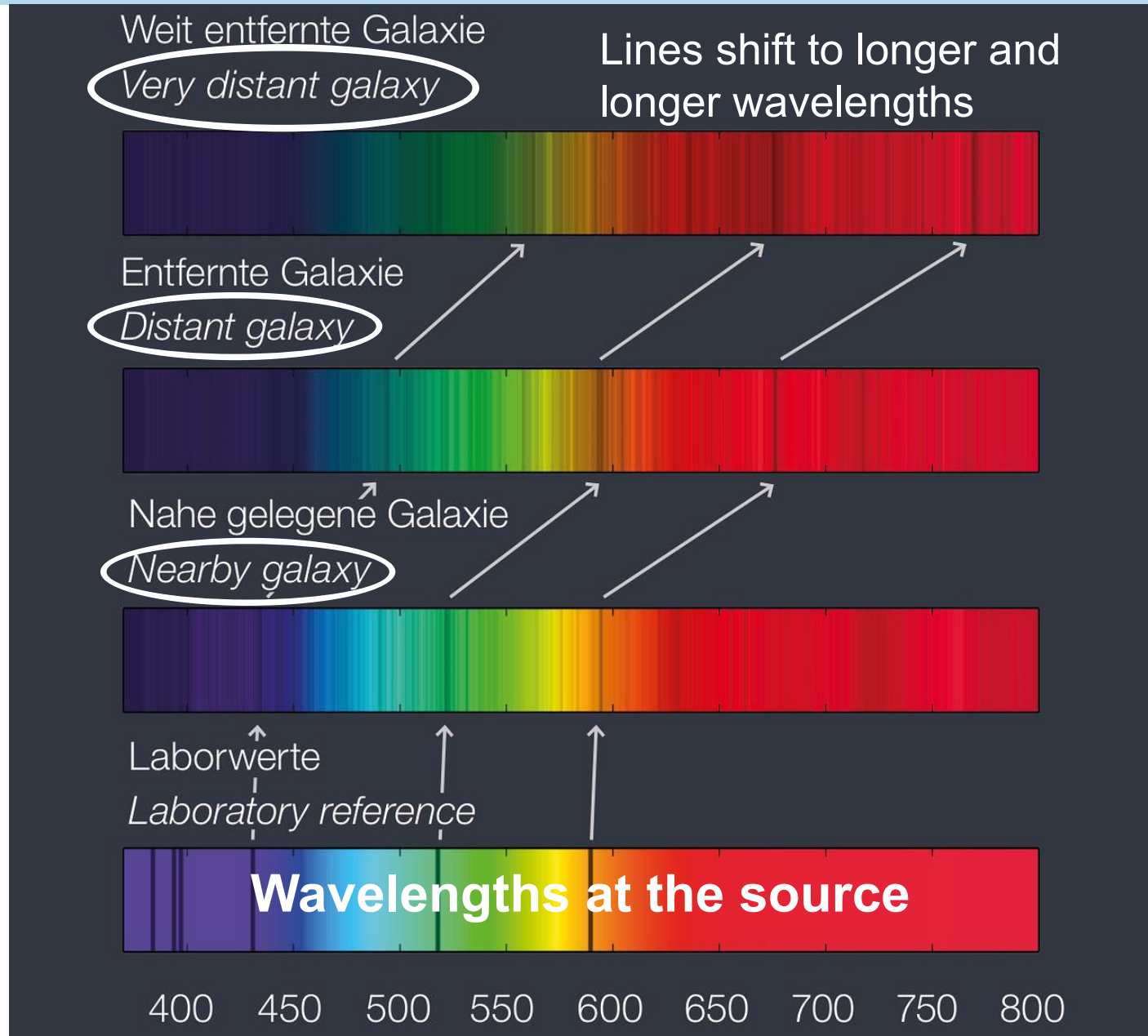


**Edwin Hubble, 1929/31**

**All distant galaxies are speeding away from us!**



# All spectral lines move to longer and longer wavelengths



## Hubble's Law: likely because the universe is expanding

- This linear relationship is called **Hubble's Law**. The slope of the line is characterized by **Hubble's constant**  $H_0$ :

$$\text{recessional velocity} = H_0 \times \text{distance}$$

- The value of Hubble's constant is approximately 70 (km/s)/Mpc. **Hubble's law** relates how **fast** galaxies are moving away from us at **different distances**. A larger value for  $H_0$  implies a faster expansion rate.
- The **simplest explanation** for Hubble's observation was that the **entire universe is expanding**. If so, then at a time in the past it should have started from a single point - an idea known today as the **Big Bang**.

# Expanding started from a “Big Bang”, coined by Fred Hoyle

Hoyle used "**Big Bang**" for the first time in his “Nature of the Universe” talks, but **with derision**, to describe a theory on the origin of the universe that he didn't accept.



Fred Hoyle was a regular on BBC radio programs in the 1940s and 1950s, and one of the most influential figures in the field of stellar nucleosynthesis

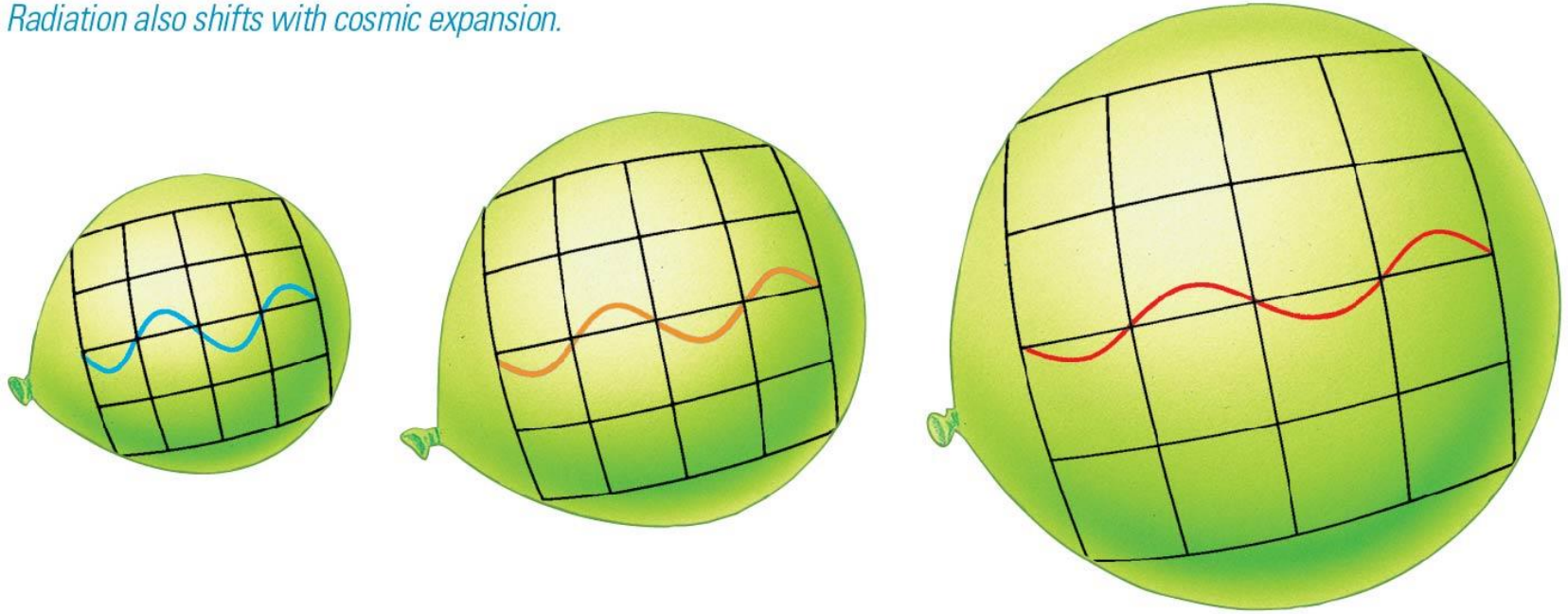
(<https://bigthink.com/starts-with-a-bang/big-bang-meaning/> )



# Universe expansion is recorded by the light that reaches Earth

- As the universe expands so do the wavelengths of the light travelling through it!

*Radiation also shifts with cosmic expansion.*

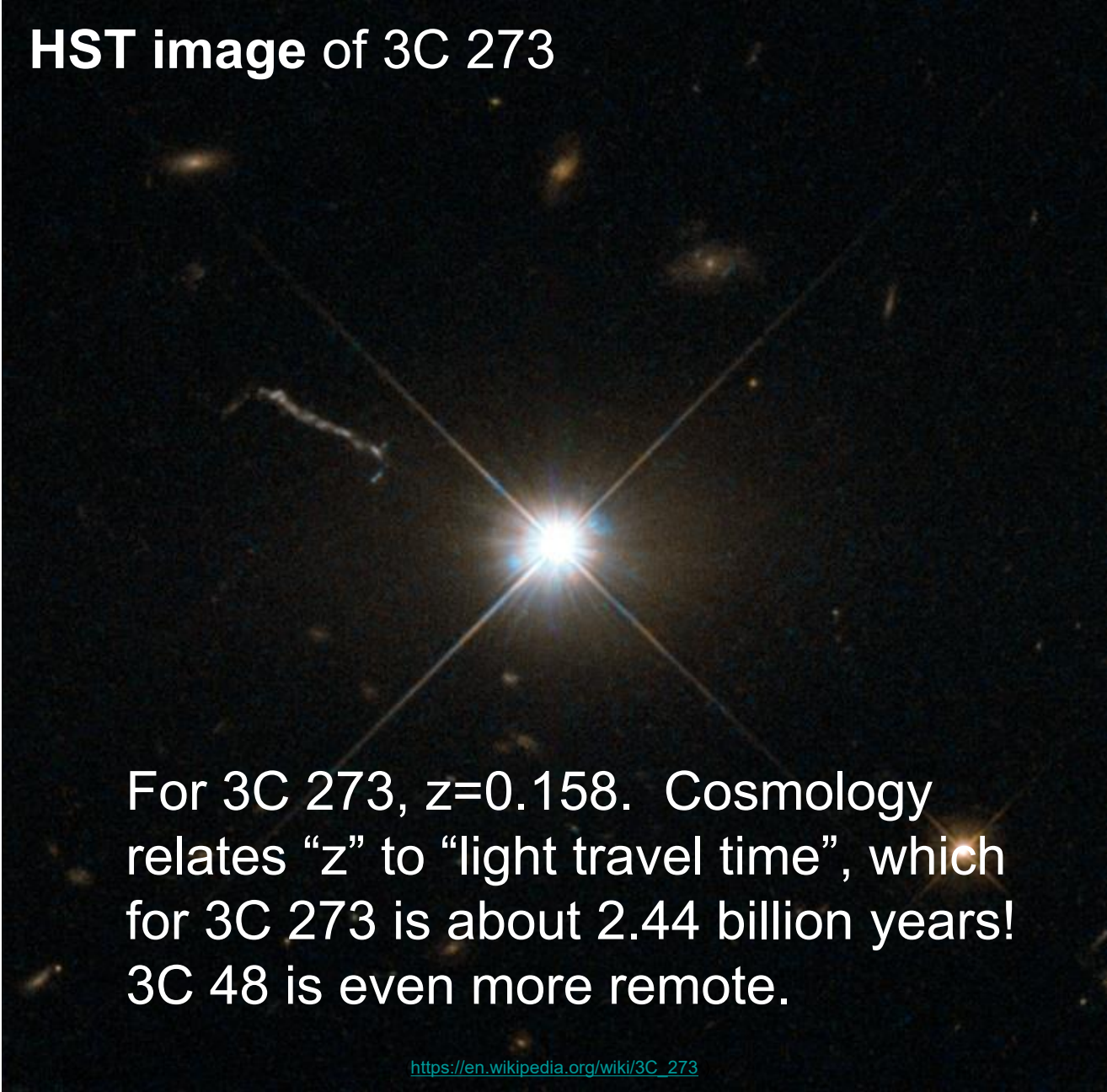


Cosmologically the factor  $\lambda_{\text{observed}} / \lambda_{\text{rest}} = 1 + z$  is not a Doppler effect but rather the fractional stretching of light wavelengths since leaving their source.

Because the **universe expansion scale factor** is also  $1 + z$ , the universe (and the wavelengths of light) have grown by a factor of  $1 + z$  since this light was emitted!

## 3C 48 and 3C 273 are at extreme distances

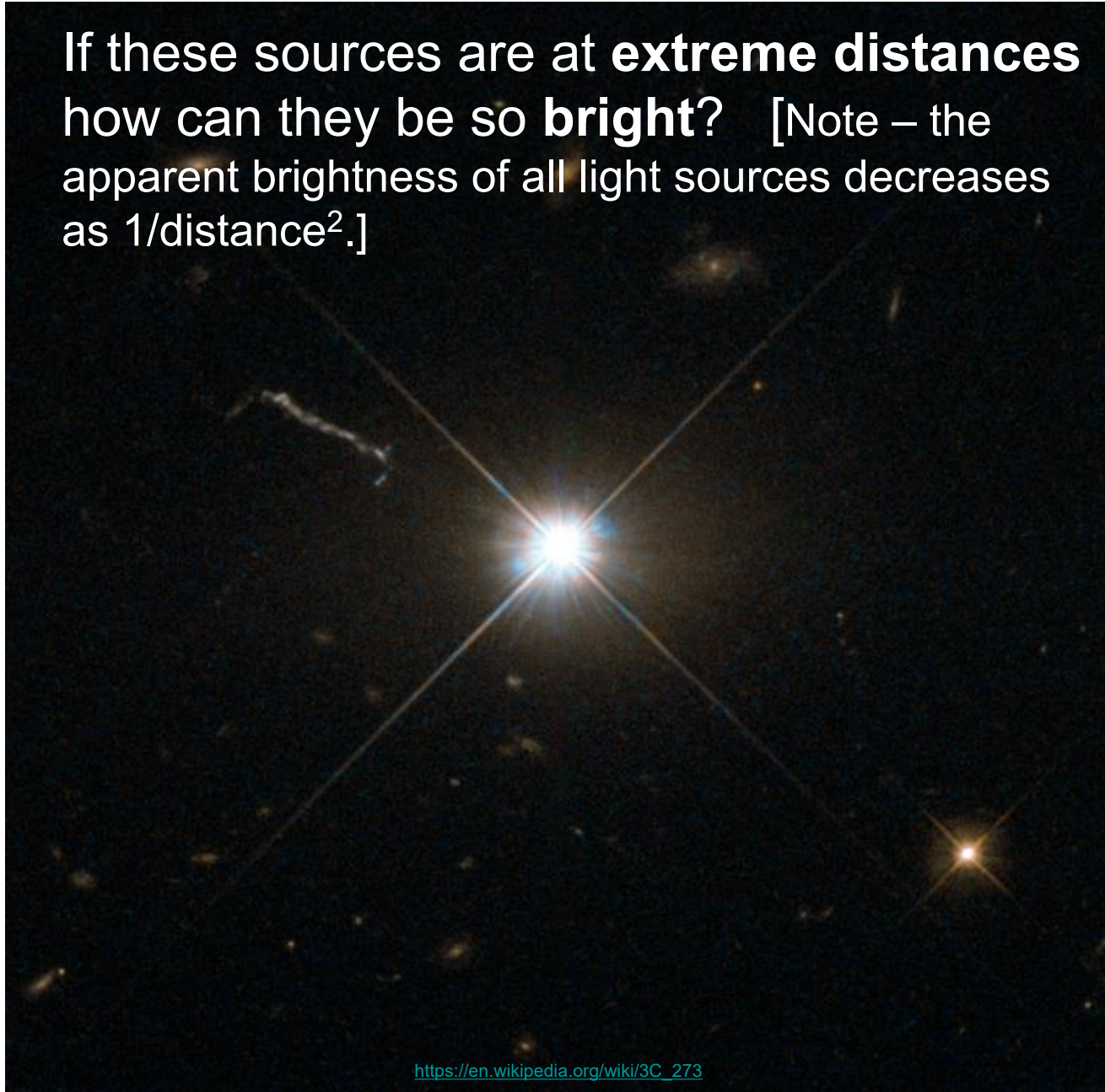
### HST image of 3C 273



For 3C 273,  $z=0.158$ . Cosmology relates “ $z$ ” to “light travel time”, which for 3C 273 is about 2.44 billion years! 3C 48 is even more remote.

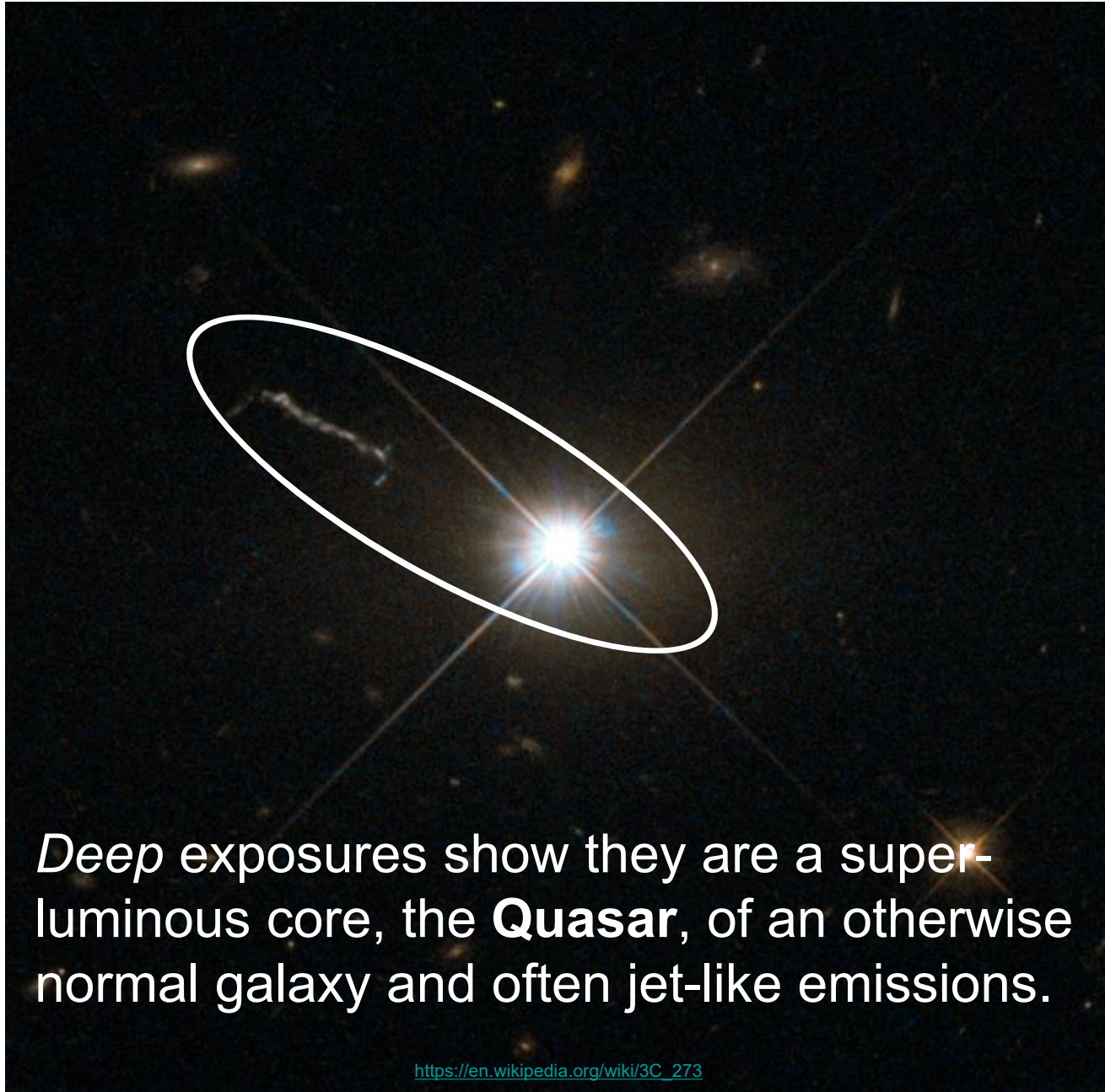
## 3C 48 and 3C 273 intrinsic luminosities were unimaginable

If these sources are at **extreme distances** how can they be so **bright**? [Note – the apparent brightness of all light sources decreases as  $1/\text{distance}^2$ .]



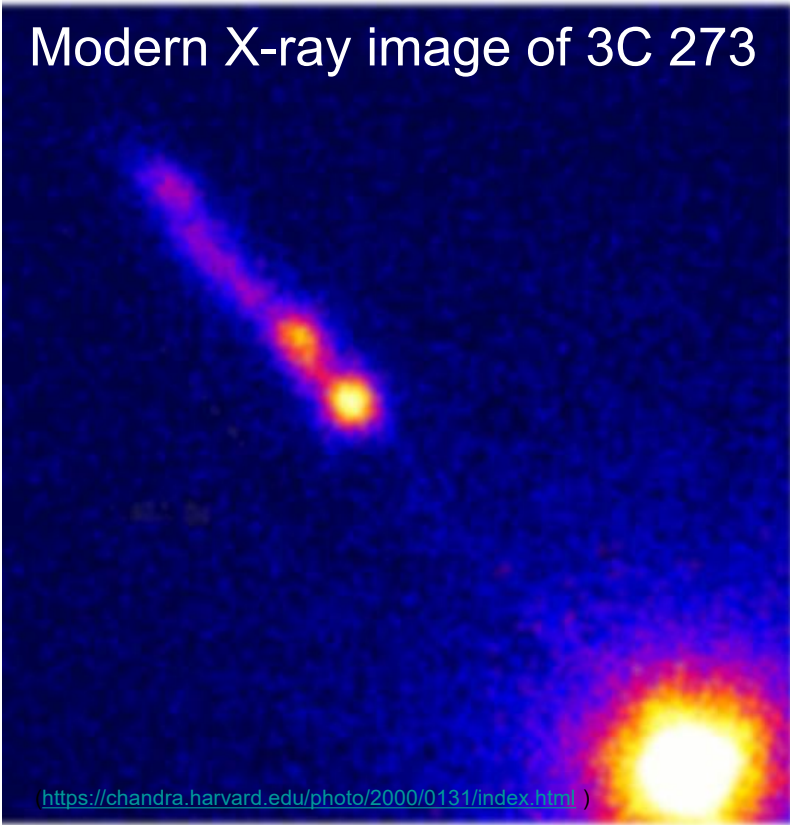


## 3C 48 and 3C 273 had other, non-star-like features



# Schmidt's “star-like object” is one of the closest Quasars

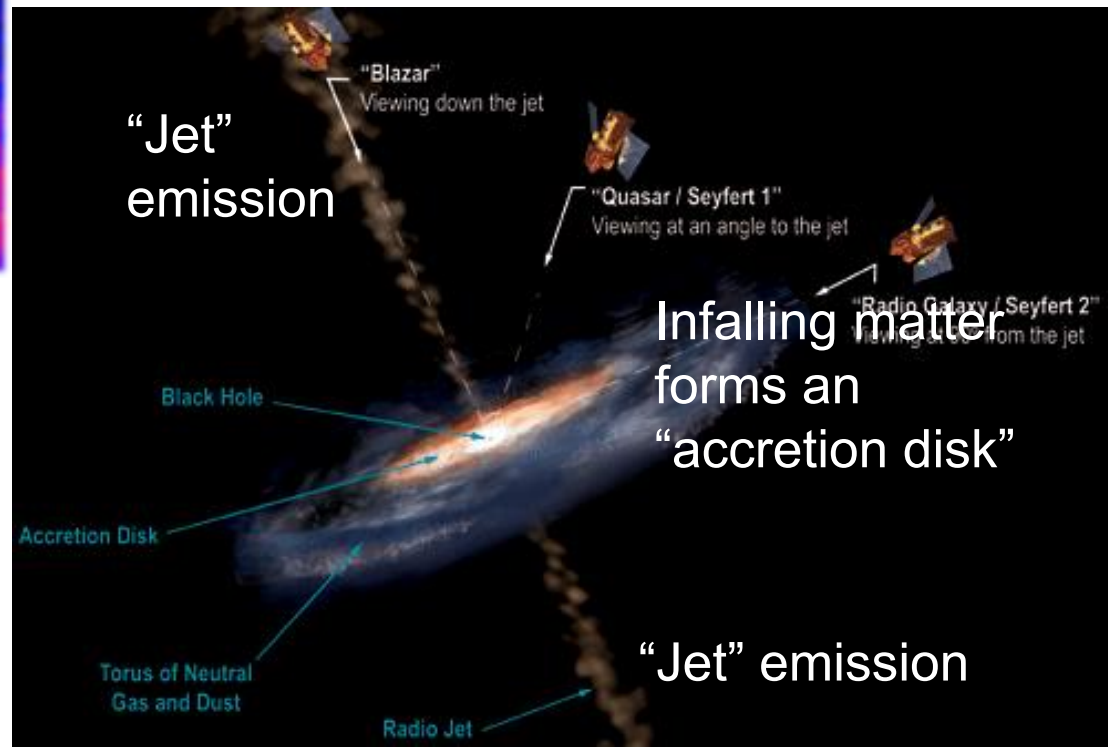
Modern X-ray image of 3C 273



Quasars are thought to be super-massive Black Holes in a period of “peak” mass acquisition ...

(Top left): X-ray image of 3C 273

(Bottom right): artist sketch of typical model of **super-massive Black Holes**

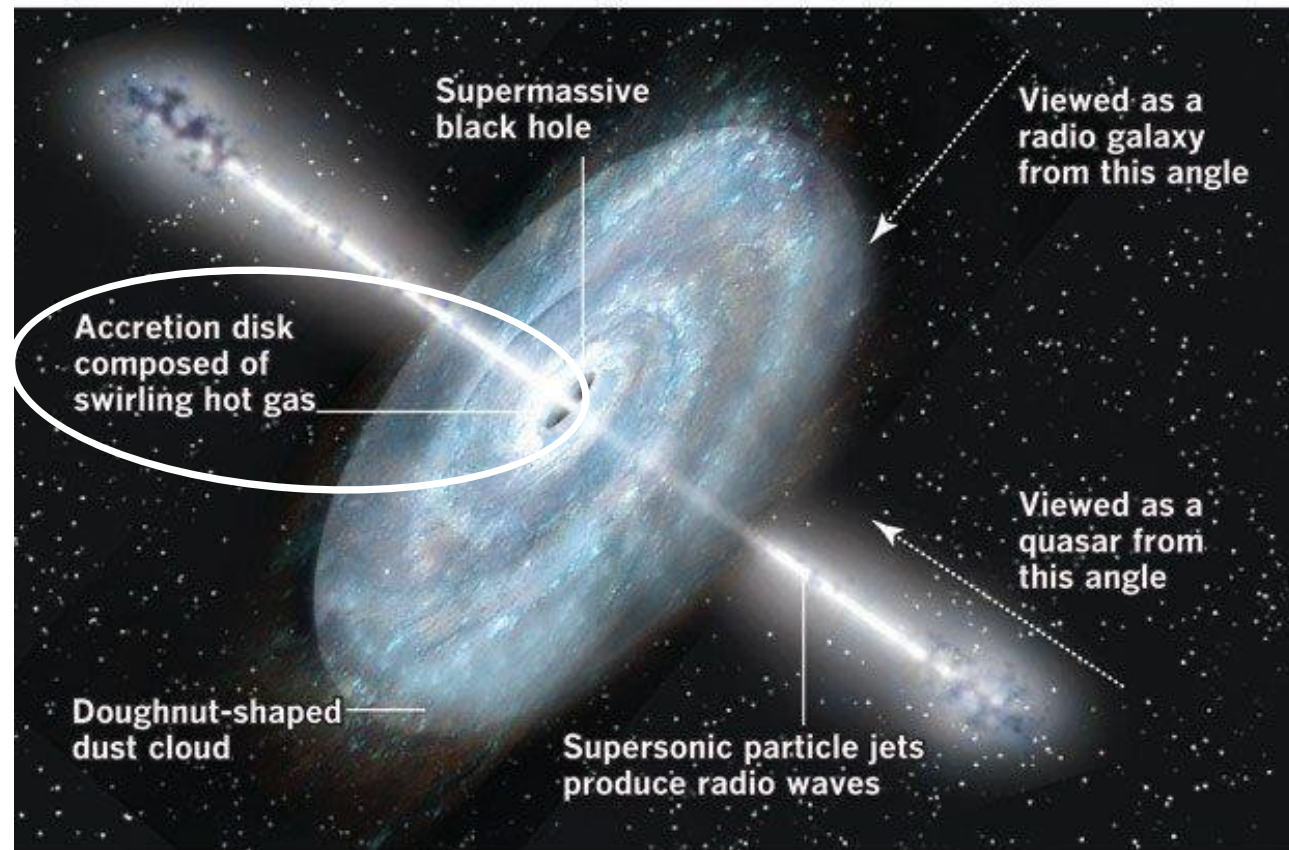


# Quasars are mostly early-universe super-massive black holes

**Quasars** are likely the **most luminous** objects in the universe!

Quasars are **super-massive Black Holes** actively **accreting matter** found predominantly in the **early universe!**

The luminous glow of a quasar is powered by the accumulation of matter into a giant black hole at the center of a galaxy.



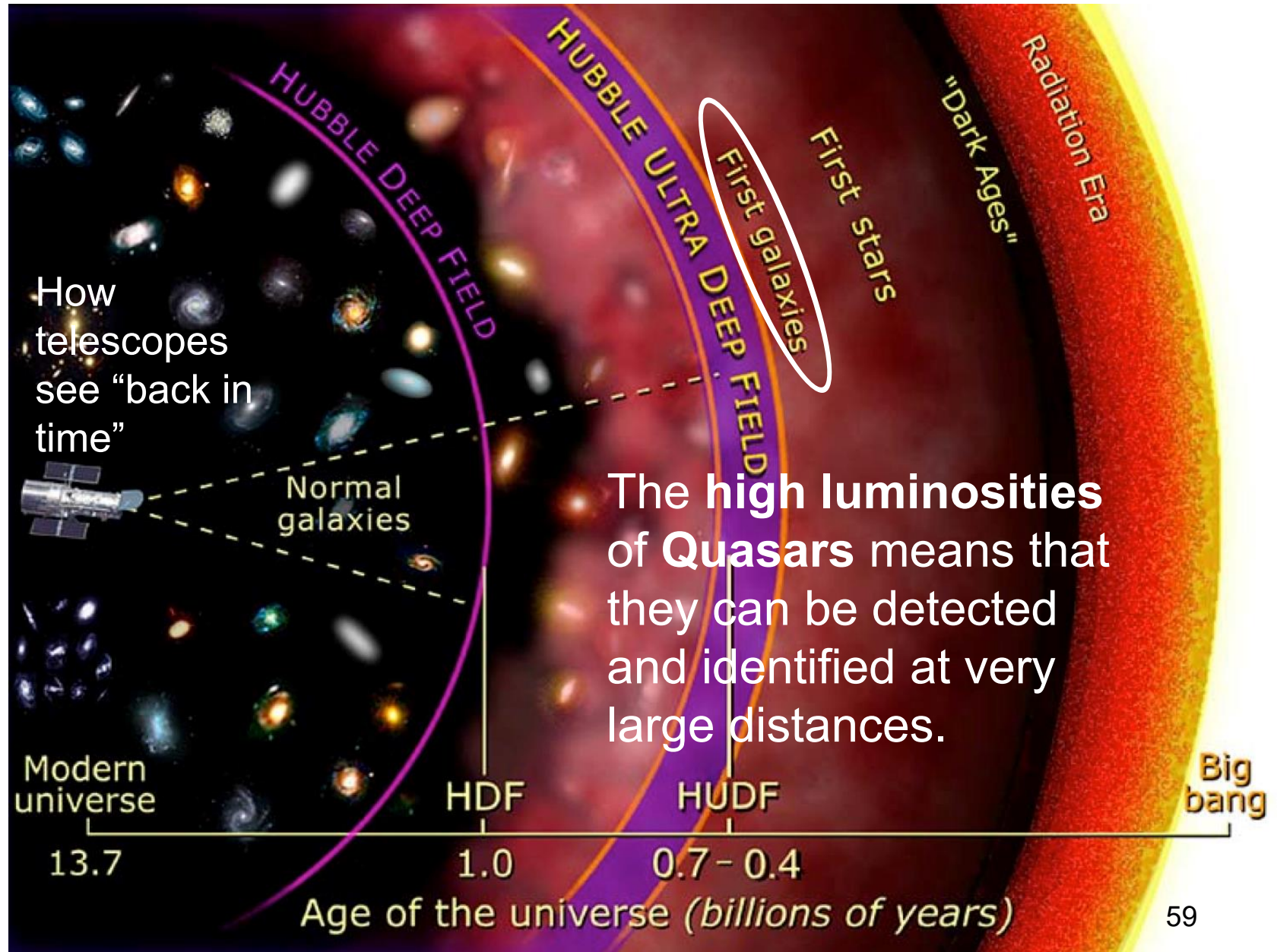
Source: Nature. Graphics reporting by **BRADY MACDONALD**

**LORENA INIGUEZ ELEBEE** Los Angeles Times

The **accretion disk** is very, very, very luminous! 58



# Quasars should provide insights into the early universe



# Scientists are competitive: the hunt to find Quasars

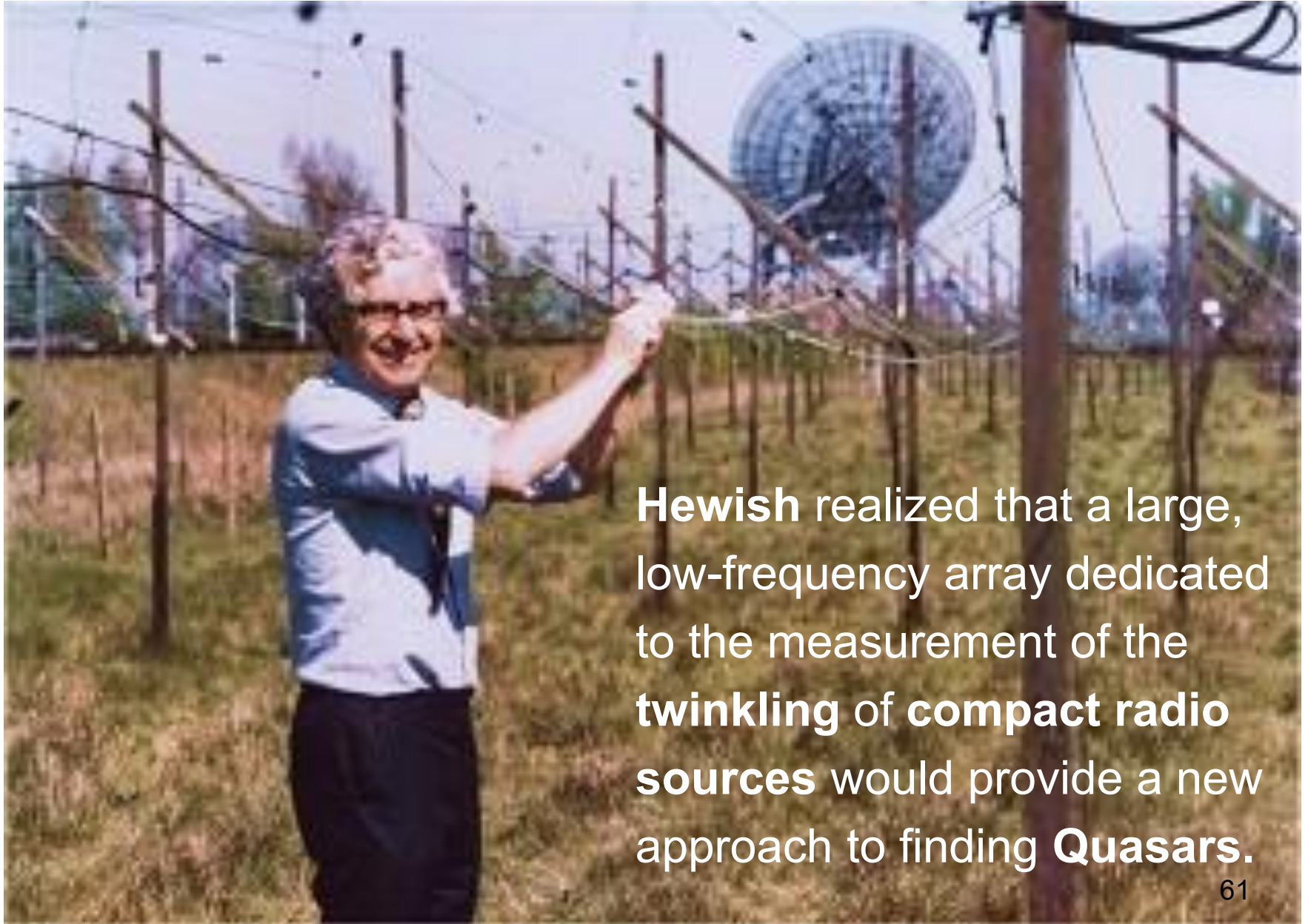
By 1964 a number of radio **Quasars** were known, some of these were **point-like**, *i.e.* had small angular sizes.

**Antony Hewish** showed that point-like radio sources would ***twinkle*** as the radio waves were slightly modified by small inhomogeneities in the solar wind flowing out from the Sun.





# Hewish's “Interplanetary Scintillation Array”



**Hewish** realized that a large, low-frequency array dedicated to the measurement of the **twinkling of compact radio sources** would provide a new approach to finding **Quasars**.

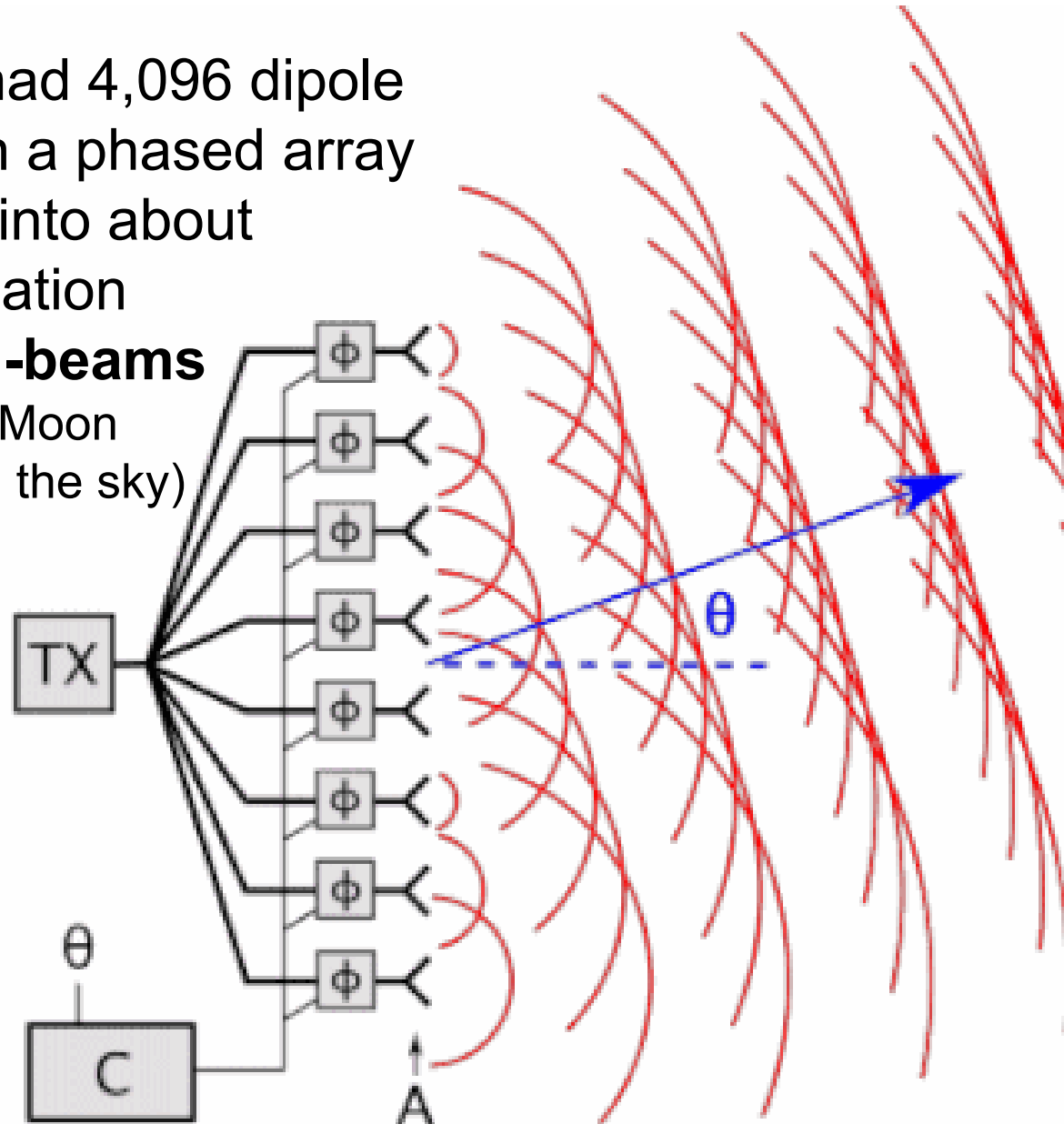


In 1965 Hewish received a grant of £17,286 to construct it

The array had 4,096 dipole antennas in a phased array configured into about **5°-in-declination**

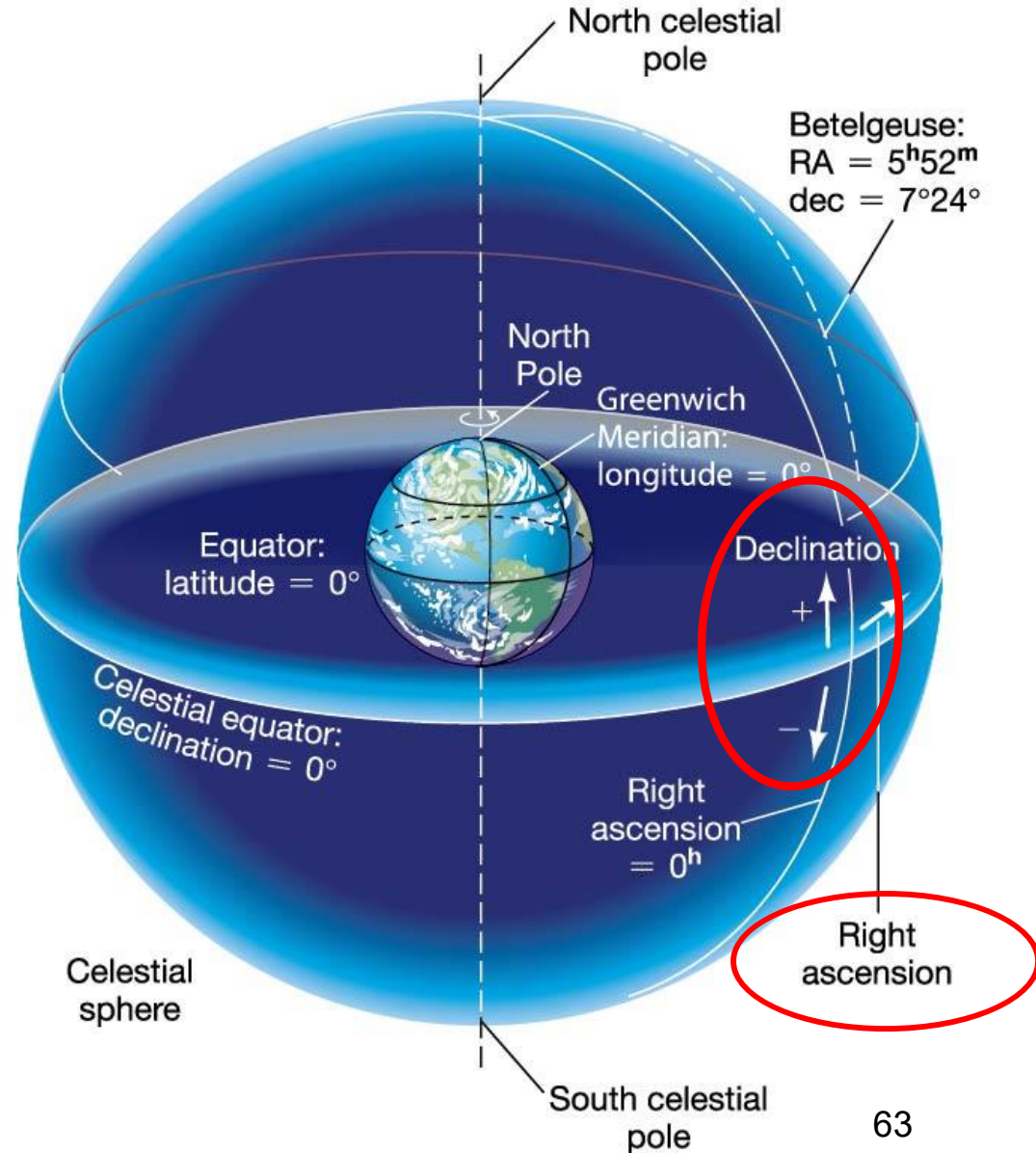
**directional-beams**

(FYI: 5° = 10-Moon diameters on the sky)



# Earth: latitude/longitude. Stars: declination/right ascension.

- **Declination** (like latitude): degrees north(+) or south(-) of the Celestial Equator.
- **Right Ascension** (like longitude): but measured in hours, minutes and seconds with  $1 \text{ hour} = 15^\circ$ .





# The “phased array” does not look like a radio telescope



The instrument originally covered about 4 acres (16,000 m<sup>2</sup>) and was mowed by sheep.



# It took a couple of years to bring the array online



**Jocelyn Bell** joined the project as a graduate student in October 1965

(<https://sciencesprings.wordpress.com/2019/03/19/from-national-radio-astronomy-observatory-astronomers-find-cannonball-pulsar-speeding-through-space/dame-susan-jocelyn-bell-burnell-at-work-on-first-pulsar-chart-1967-pictured-working-at-the-four-acre-array-in-1967-image-courtesy-of-mullard-radio-astronomy-observatory/> )

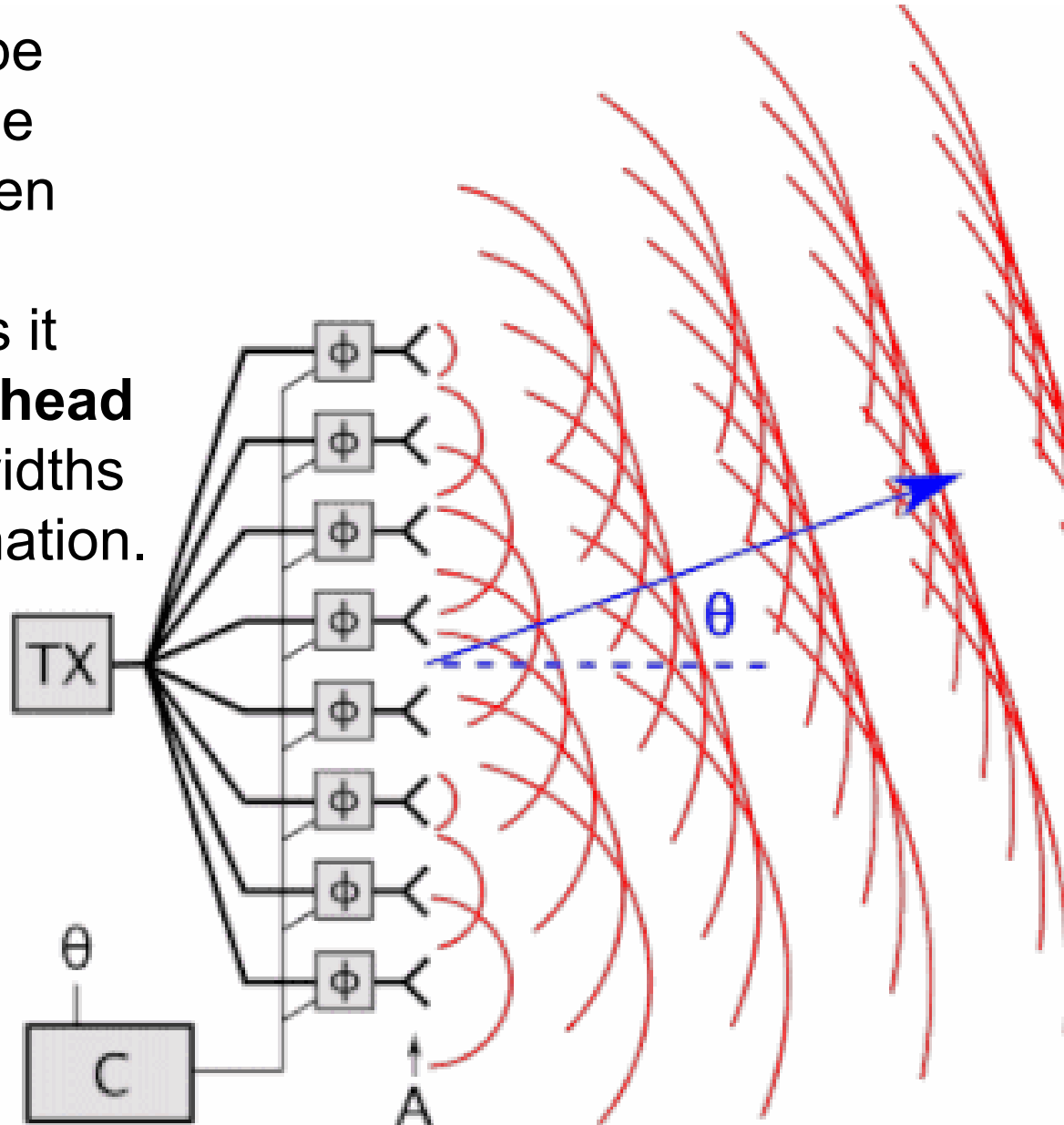
# The telescope was commissioned during July 1967



The key aspect was to measure the **twinkling (time variations of brightness)** of the radio sources.

## 4 of the declination-beams were recorded at any given time

The telescope **observed** the “sky”, between declinations  $-8^\circ$  to  $44^\circ$ , as it **drifted overhead** with beam widths  $\pm 3^\circ$  in declination.





## Fortunately data were not digital and computers inaccessible

**Bell recalls:** “The data stream was configured with four beams. The output appeared on **four 3-track pen recorders**, and produced 96 feet of chart paper every day (1foot/hour/beam). The charts were analyzed by-hand by me. We decided initially not to computerize the output. Until we were familiar with the behavior of our telescope and receivers, we thought it better to inspect the data visually. A **human** can recognize **signals of different character** whereas it is difficult to program a computer to do so.

After the first few hundred feet of chart analysis I could recognize the **twinkling sources**, and I could recognize radio interference (**noise**).”

# What is that *scruff* in the chart recording?

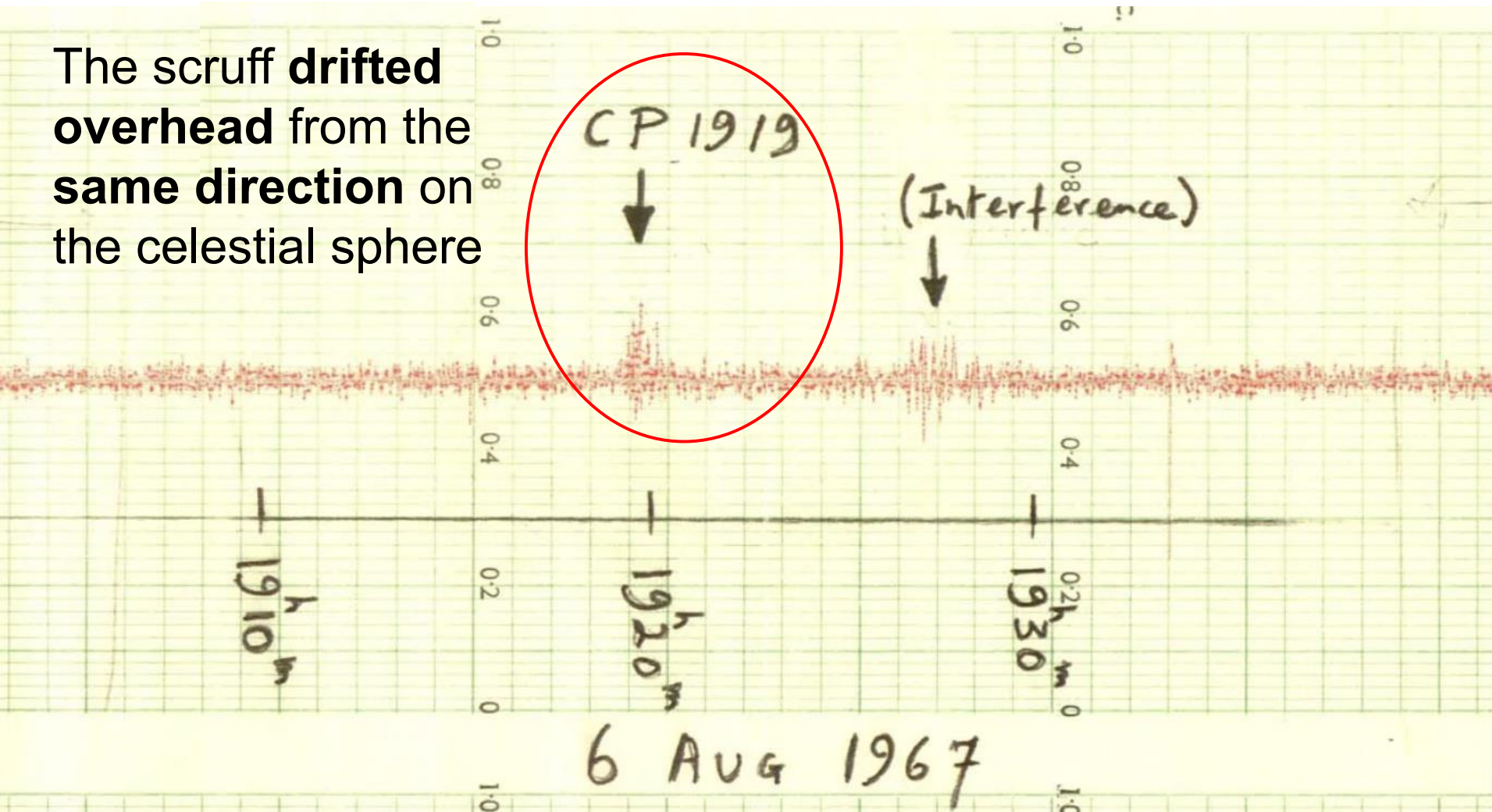
**Bell recalls:** “Six or eight weeks after starting the survey I became aware that on occasions there was a bit of **scruff** on the records, which did not look exactly like a twinkling source, and yet did not look exactly like man-made interference either.

Furthermore I realized that we had seen the **scruff** previously on the same part of the records *viz.* from the same patch of sky.” (**right ascension 1919**) [in constellation Velpecula]



# To Bell, scruff looked different from interference

The scruff **drifted**  
**overhead** from the  
**same direction** on  
the celestial sphere



Dark lines are spaced by 1 minute =  $\frac{1}{4}$  - degree in RA  
CP 1919 tells us pointing resolution likely about  $\frac{1}{4}$  -degree!<sup>70</sup>



## To interpret the scruff required faster chart recorders

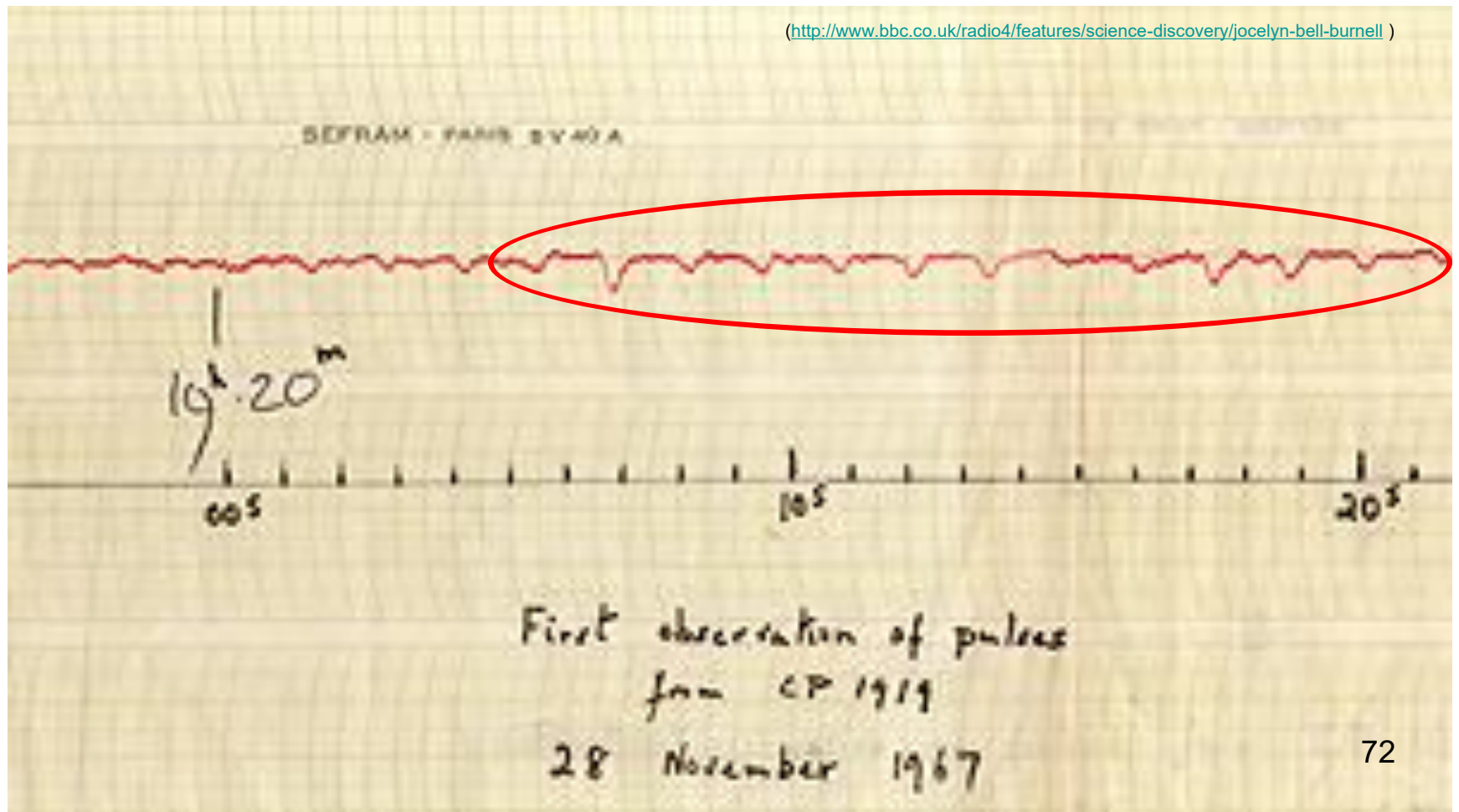
**Bell recalls:** “The source was transiting during the night — a time when solar-wind twinkling should be at a minimum, and **one idea** we had was that it was a **point source**. So we decided that it deserved closer inspection, and that this would involve making **faster chart recordings** as it transited.

Towards the end of October 1967, when we had finished doing some special tests on 3C 273 and when we had at last our full complement of receivers and recorders, I started going out to the observatory each day to make the fast recordings.

They were useless. For weeks I recorded nothing but receiver noise. Then one day I skipped the observations to go to a lecture, and on the next day on my normal recording I saw that the **scruff** had been there."

**“At the end of November 1967 I got it on the fast recording”**

**Bell recalls:** “As the chart flowed under the pen I could see that the signal was a series of pulses, and my suspicion that they were equally spaced was confirmed as soon as I got the chart off the recorder. They were 1 1/3 seconds apart.”



## Bell soon identified 3 more pulsing sources!

**Bell recalls:** “Some days later I was analyzing a recording of a completely different part of the sky, I thought I saw some **scruff**. I rapidly checked through previous recordings of that part of the sky, and on occasions there was **scruff** there.”

Ultimately two 1968 publications in Nature by Hewish, Bell, et al. reported **4 rapidly pulsing radio sources**.



# First paper in Nature on CP1919

Bell's pulsar discoveries were appendices in her thesis.

*(Reprinted from Nature, Vol. 217, No. 5130, pp. 709-713, February 24, 1968)*

## Observation of a Rapidly Pulsating Radio Source

by

A. HEWISH

S. J. BELL

J. D. H. PILKINGTON

P. F. SCOTT

R. A. COLLINS

Mullard Radio Astronomy Observatory,  
Cavendish Laboratory,  
University of Cambridge

Unusual signals from pulsating radio sources have been recorded at the Mullard Radio Astronomy Observatory. The radiation seems to come from local objects within the galaxy, and may be associated with oscillations of white dwarf or neutron stars.

# Reprint of second paper from Bell's thesis

*(Reprinted from Nature, Vol. 218, No. 5137, pp. 126-129, April 13, 1968)*

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Ph.D. Dissertation 6567

## MULLARD RADIO ASTRONOMY OBSERVATORY

# Observations of some further Pulsed Radio Sources

by

J. D. H. PILKINGTON

A. HEWISH

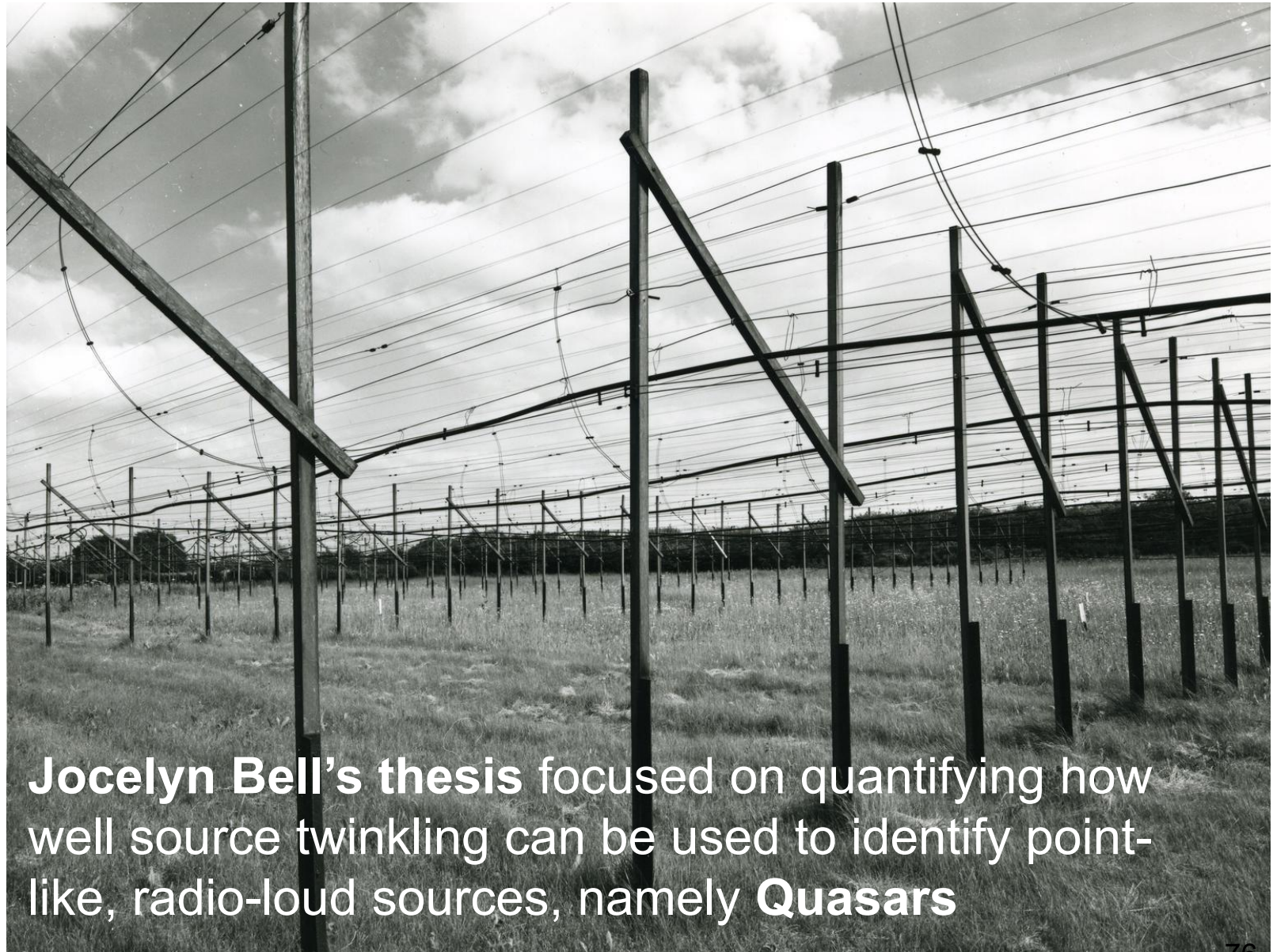
S. J. BELL

T. W. COLE

Mullard Radio Astronomy Observatory,  
Cavendish Laboratory,  
University of Cambridge

Details are now given of three of the four pulsating radio sources  
discovered at Cambridge.

# If not pulsars, what was Jocelyn Bell's thesis topic?



**Jocelyn Bell's thesis** focused on quantifying how well source twinkling can be used to identify point-like, radio-loud sources, namely **Quasars**

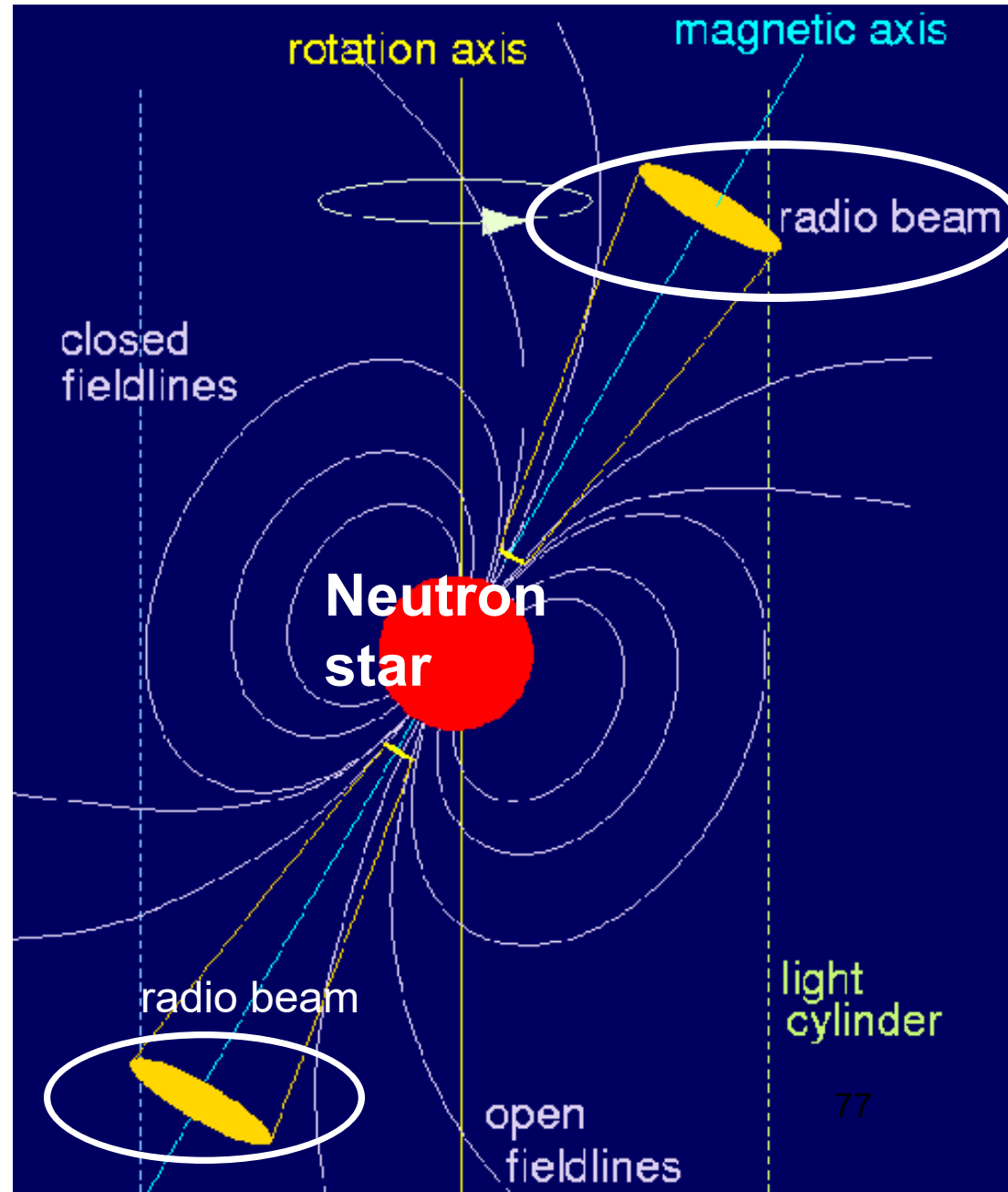


# Pulsar's were explained one year later

Thomas Gold's 1968 Nature paper proposed *Rotating Neutron Stars as the Origin of the Pulsing Radio Sources*.

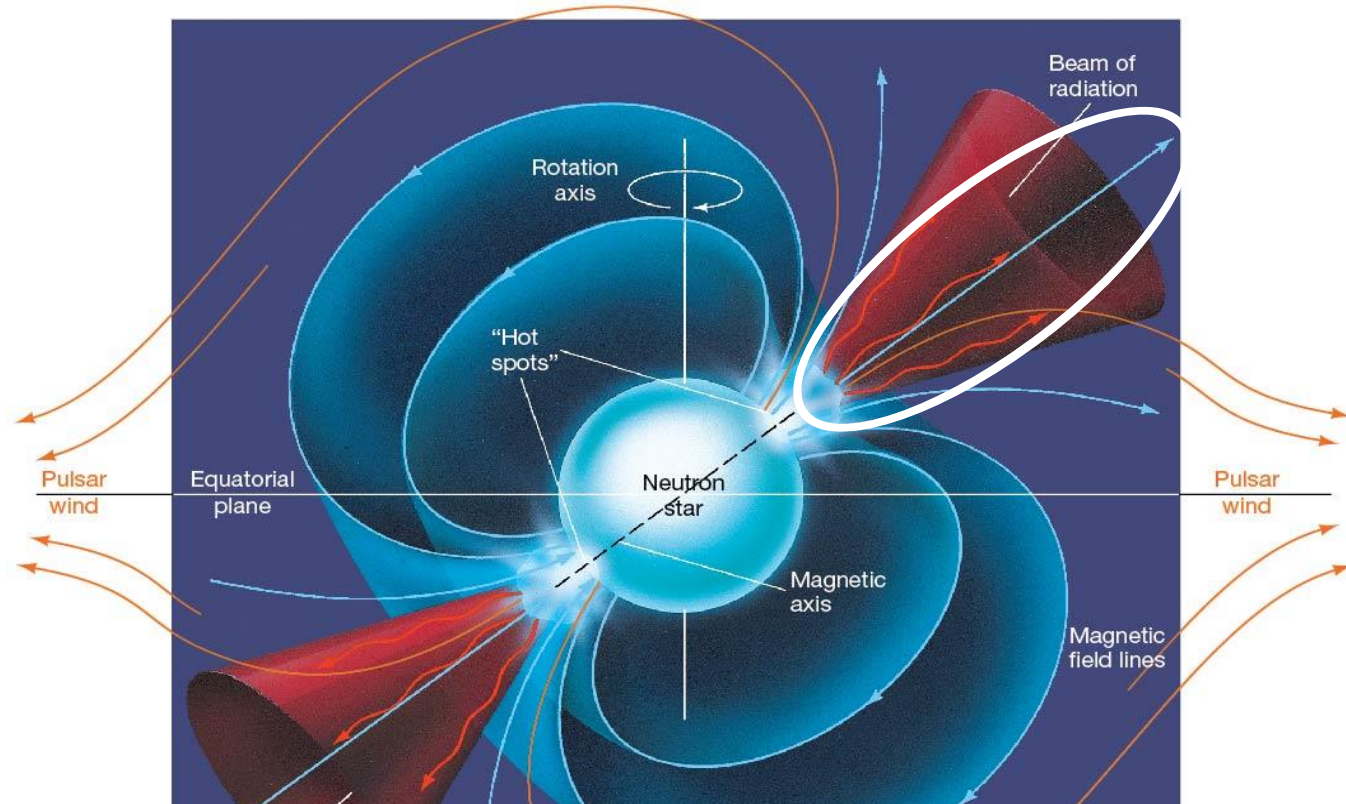
Neutron stars' intense, rotating, magnetic fields result in beamed radio emission along the magnetic axis of the neutron star.

These **searchlight-like beams** cause a pulse when they sweep across the Earth.



# Rotating lighthouse beacon appears to flash on and off

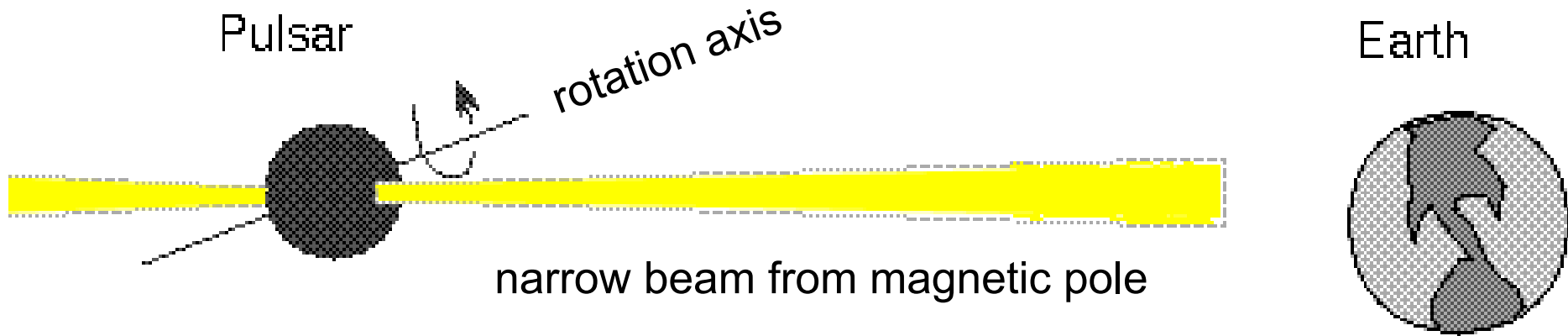
- But why would a neutron star flash on, and off?
- This figure illustrates the analogy with a **lighthouse beacon**.



A lighthouse beacon  
is a good analogy for  
a rotating pulsar.



# From scruff, to pulsar, to neutron star



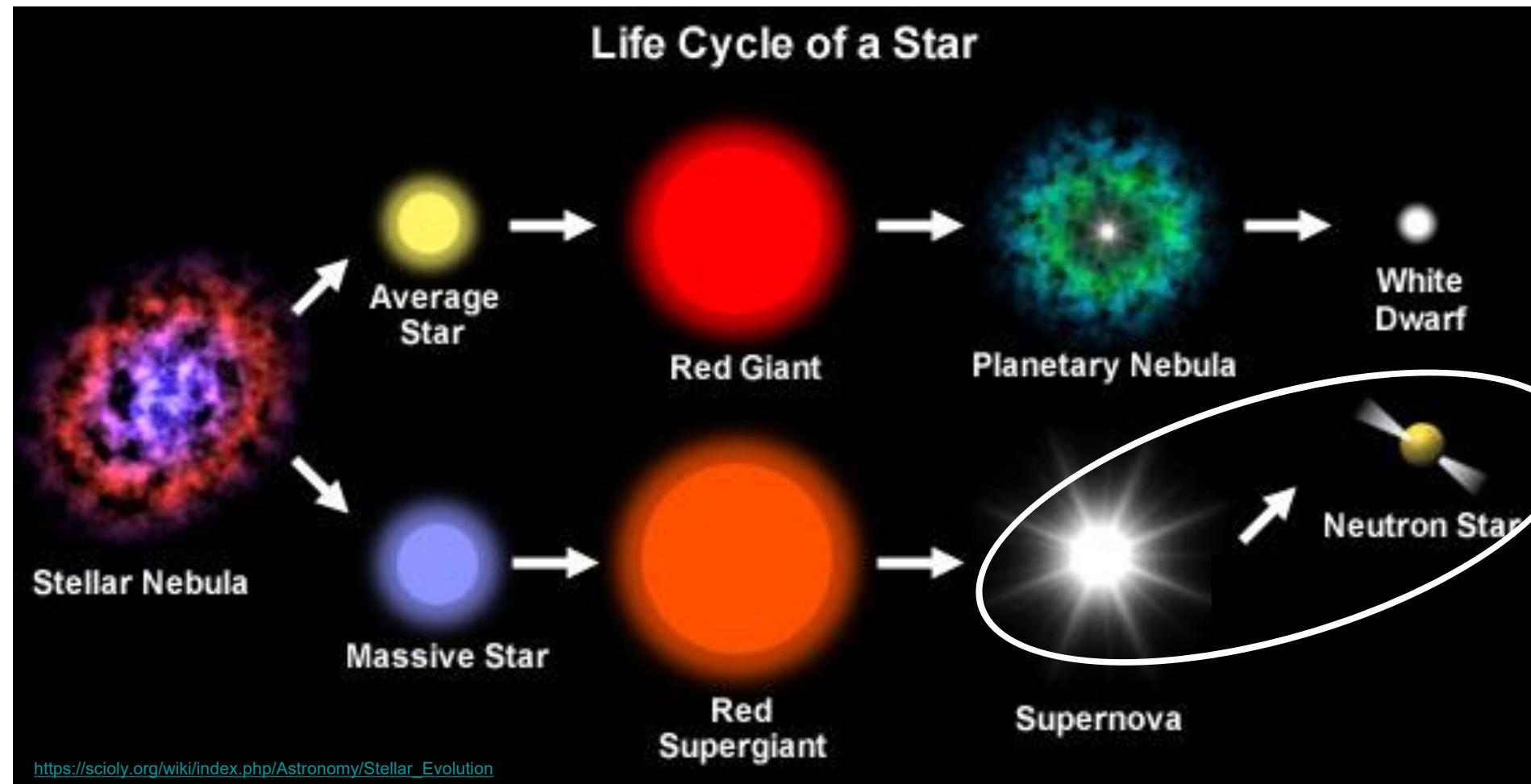
If the narrow synchrotron beam passes over the Earth, we see the neutron star flash on and off like a lighthouse beam does for ships at sea.

This lucky mechanism is almost the only way to observe the **neutron star** remnants of **Core Collapse supernovas!**



# Conjectured supernova remnants were proposed in 1934

- **Neutron stars** form in the *death* of 10 ~ 20 solar-mass stars as **Core Collapse supernovas**. They have masses of 1 to 3 solar-masses but radii of about 10km!
- It took 34 years for **evidence** of their existence.



# This was the 1<sup>st</sup> Nobel Prize in astronomy/astrophysics

## The Nobel Prize in Physics 1974

(<https://www.nobelprize.org/prizes/physics/1974/summary/> )



Photo from the Nobel  
Foundation archive.

**Sir Martin Ryle**

Prize share: 1/2



Photo from the Nobel  
Foundation archive.

**Antony Hewish**

Prize share: 1/2

The Nobel Prize in Physics 1974 was awarded jointly to Sir Martin Ryle and Antony Hewish "for their pioneering research in radio astrophysics: Ryle for his observations and inventions, in particular of the aperture synthesis technique, and Hewish for his decisive role in the discovery of pulsars"

# Nobel prize for Hewish and Ryle but not Bell

Jocelyn Bell received many significant awards over her career:

- **Order of the British Empire** (1999)
- Dame Commander **OBE** (2007)
- **Special Breakthrough Prize in Fundamental Physics** (2018)

*“The \$3 million award recognizes not only Bell Burnell's 1967 discovery of the weird, fast-spinning stellar corpses known as pulsars but also her scientific leadership in the 50 years since then.”*





# President John F. Kennedy Signs the Nuclear Test Ban Treaty

Treaty Banning **Nuclear Weapon Tests** in the Atmosphere, in Outer Space and Under Water, signed by Kennedy July 10, 1963

Treaty was also initially signed by the U.K. and the Soviet Union. By the end of 1963, more than 120 other nations had signed the treaty.



# Energetic gamma-rays are a *weapons* signature

Vela (velar: to *keep vigil over*) was the name of a group of satellites developed at Los Alamos and Sandia Labs to **monitor compliance** by the Soviet Union with the **1963 Partial Test Ban Treaty**.

Vela started in 1959 as a small-budget research program.



## In 1967 our knowledge of the universe changed

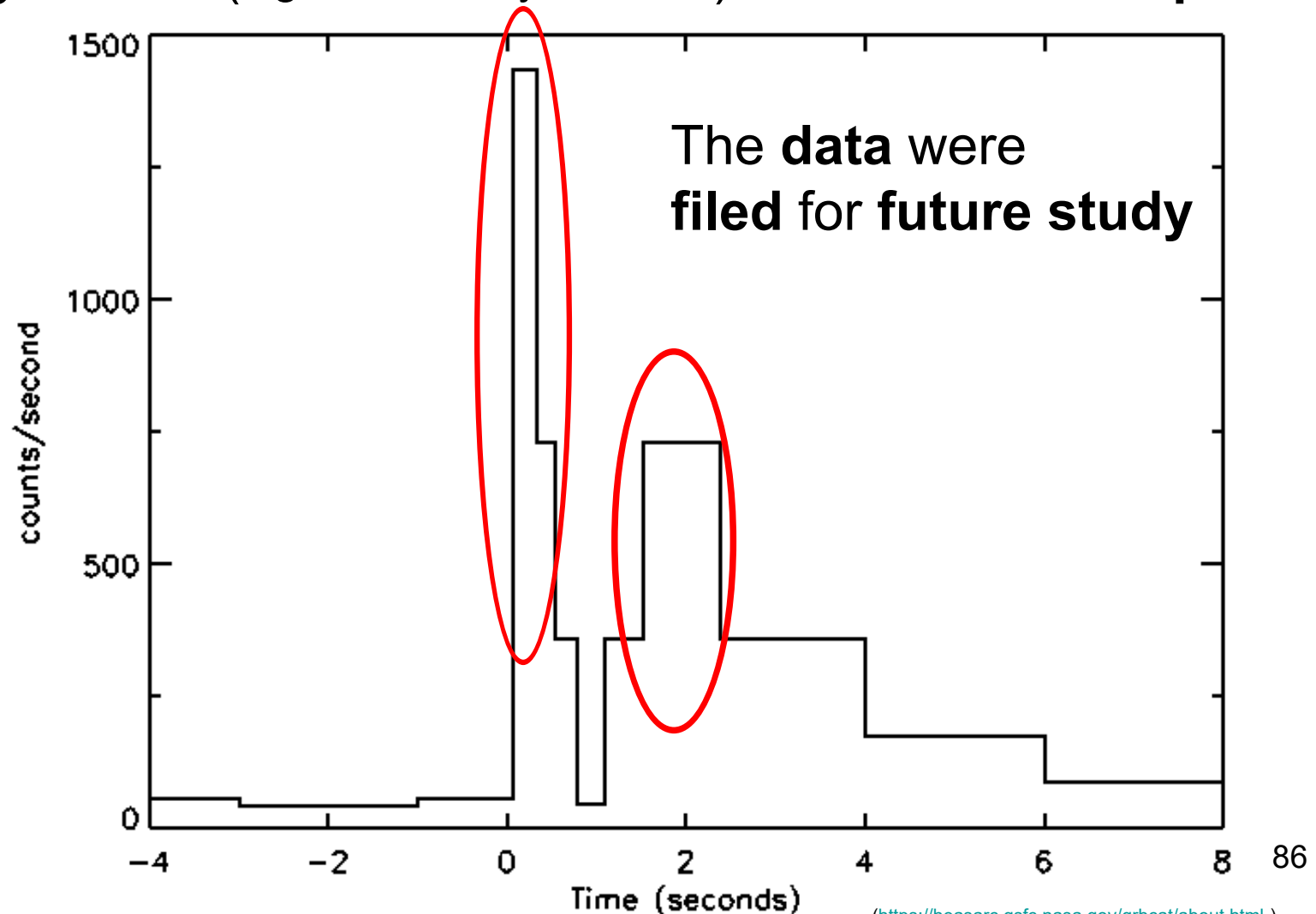
- On **July 2, 1967**, at 14:19 UTC, the Vela 4 and Vela 3 satellites detected a **burst of gamma radiation** that was **unlike any known nuclear weapons signatures**.
- Nuclear bombs produce a very brief, intense burst of gamma rays: less than one millionth of a second. The radiation then steadily fades as the unstable nuclei decay.



# The observed burst of gamma-rays lasted several seconds

The signal detected by the Vela satellites had neither the intense initial flash nor the gradual fading.

The light curve (signal intensity VS time) had **two distinct peaks**.

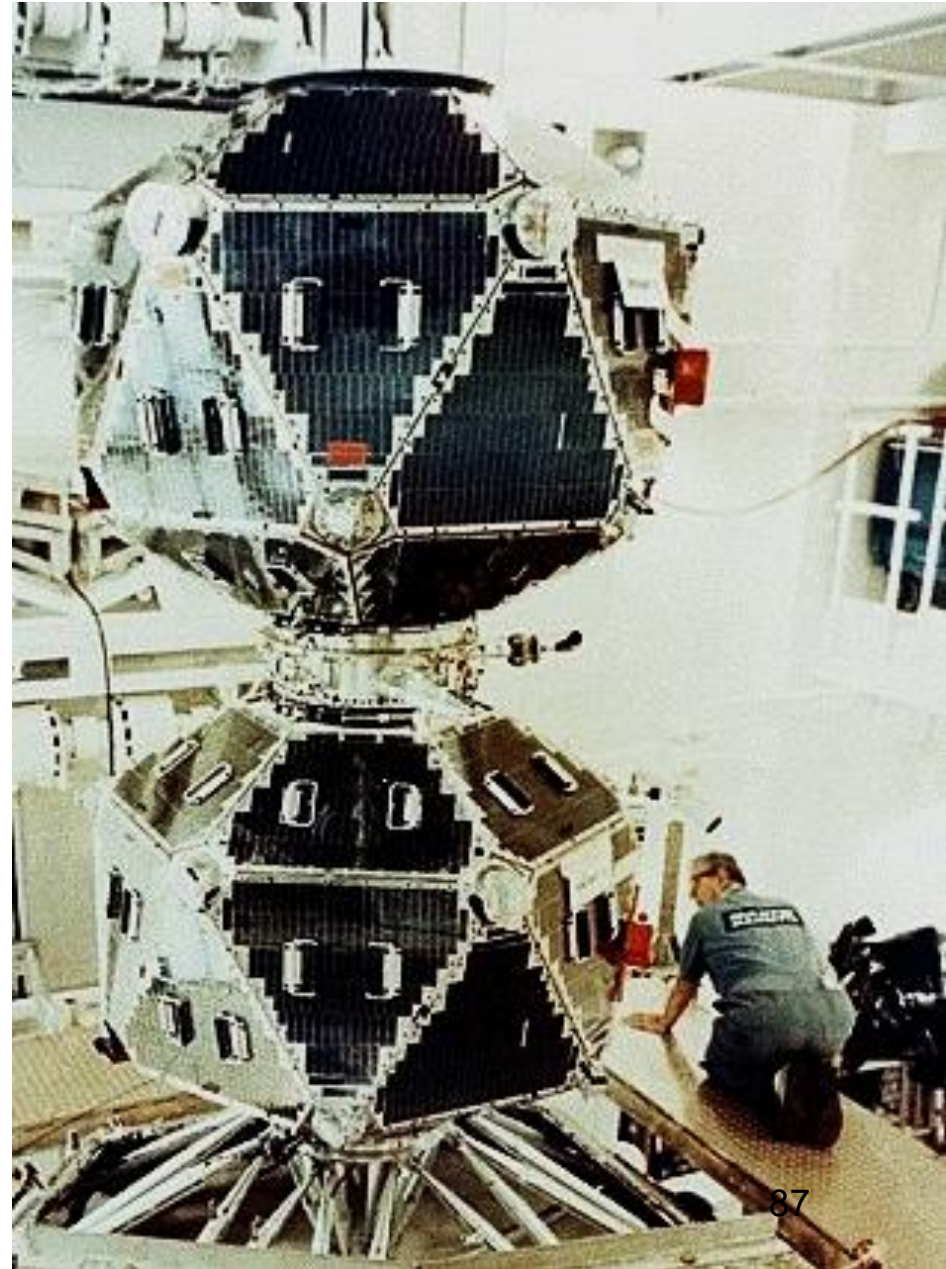


## Vela 5 was launched on May 23, 1969

With **improved sensitivity** and **time resolution**, the Los Alamos team expected these new satellites to detect more gamma-ray bursts.

They found **twelve events** that did not coincide with any solar flares or terrestrial events.

Some of the new detections also showed the same double-peak pattern observed by Vela 4

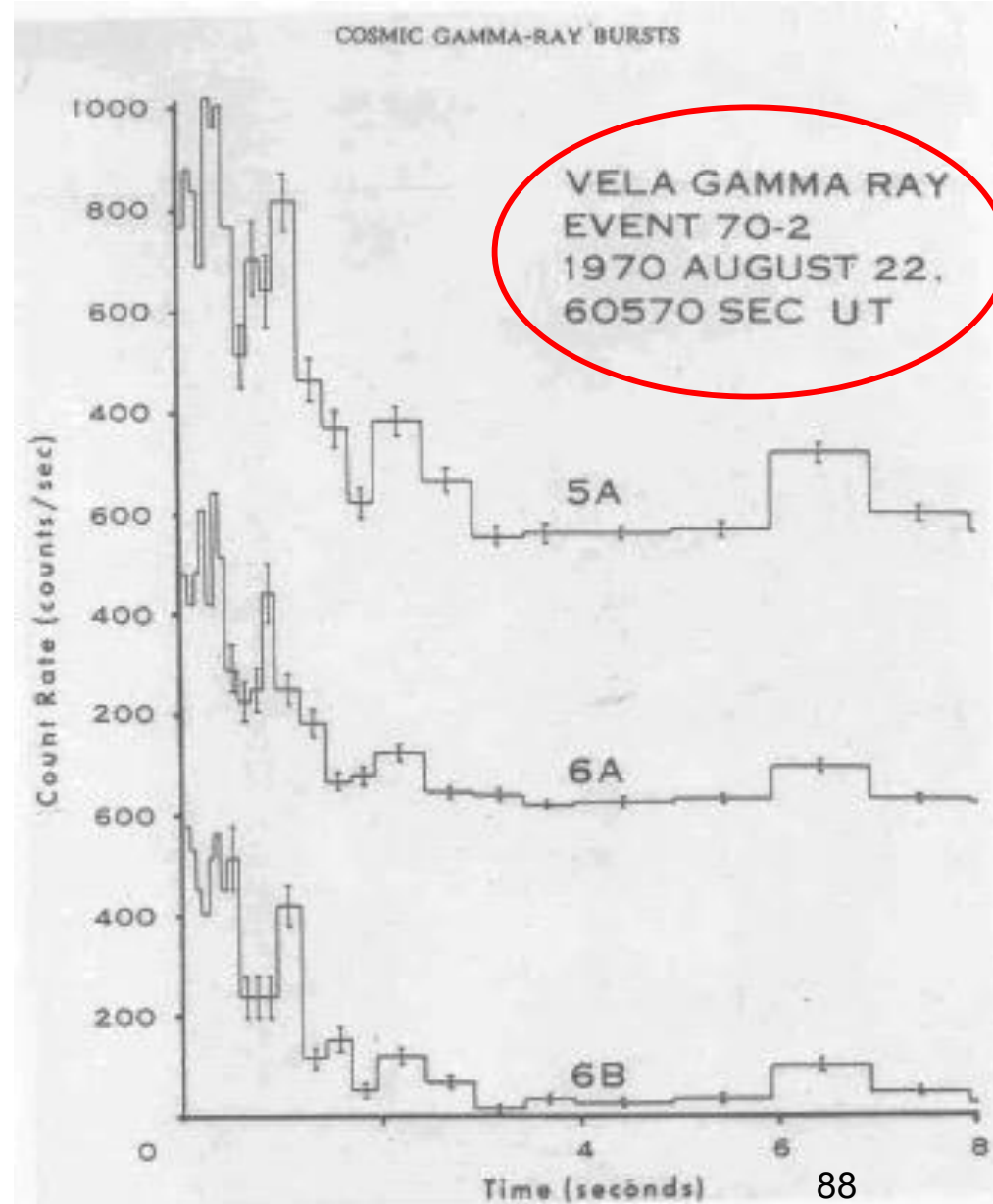


# Vela 6 satellites were launched on April 8, 1970

The Vela 6 satellite orbits were chosen to be as **far away** from **Vela 5** as possible.

This separation meant that, despite gamma rays traveling at the speed of light, a signal would be detected at slightly different times by the satellites.

By analyzing the **arrival times**, the Los Alamos team successfully traced the source direction of **sixteen gamma-ray bursts**.





## Directions of the gamma-ray bursts indicated non-local origin!

The bursts were **not** coming from the **Sun**, **Moon**, or other **planets** in **our solar system**.

In 1973, Ray Klebesadel, Roy Olson, and Ian Strong of the Los Alamos Scientific Laboratory published ***“Observations of Gamma-Ray Bursts of Cosmic Origin”***.

**Abstract:** Sixteen short bursts of photons in the energy range 0.2-1.5 MeV have been observed between 1969 July and 1972 July using widely separated spacecraft. Burst durations ranged from less than 0.1 s to ~30 s, and time-integrated flux densities from  $\sim 10^{-5}$  ergs/cm<sup>2</sup> to  $\sim 2 \times 10^{-4}$  ergs/cm<sup>2</sup> in the energy range given. Significant time structure within bursts was observed. Directional information eliminates the Earth and Sun as sources.

**Amazing science but not recognized. Likely too unexpected.**

Gamma-ray bursts were just “**too unexpected**”.

For some reason there was little press of **Ray Klebesadel’s**, Roy Olson’s, and Ian Strong’s 1973 Gamma-ray Burst publication.

**So far just one photo**



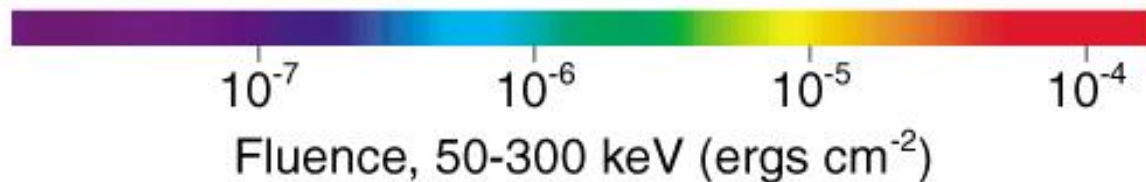
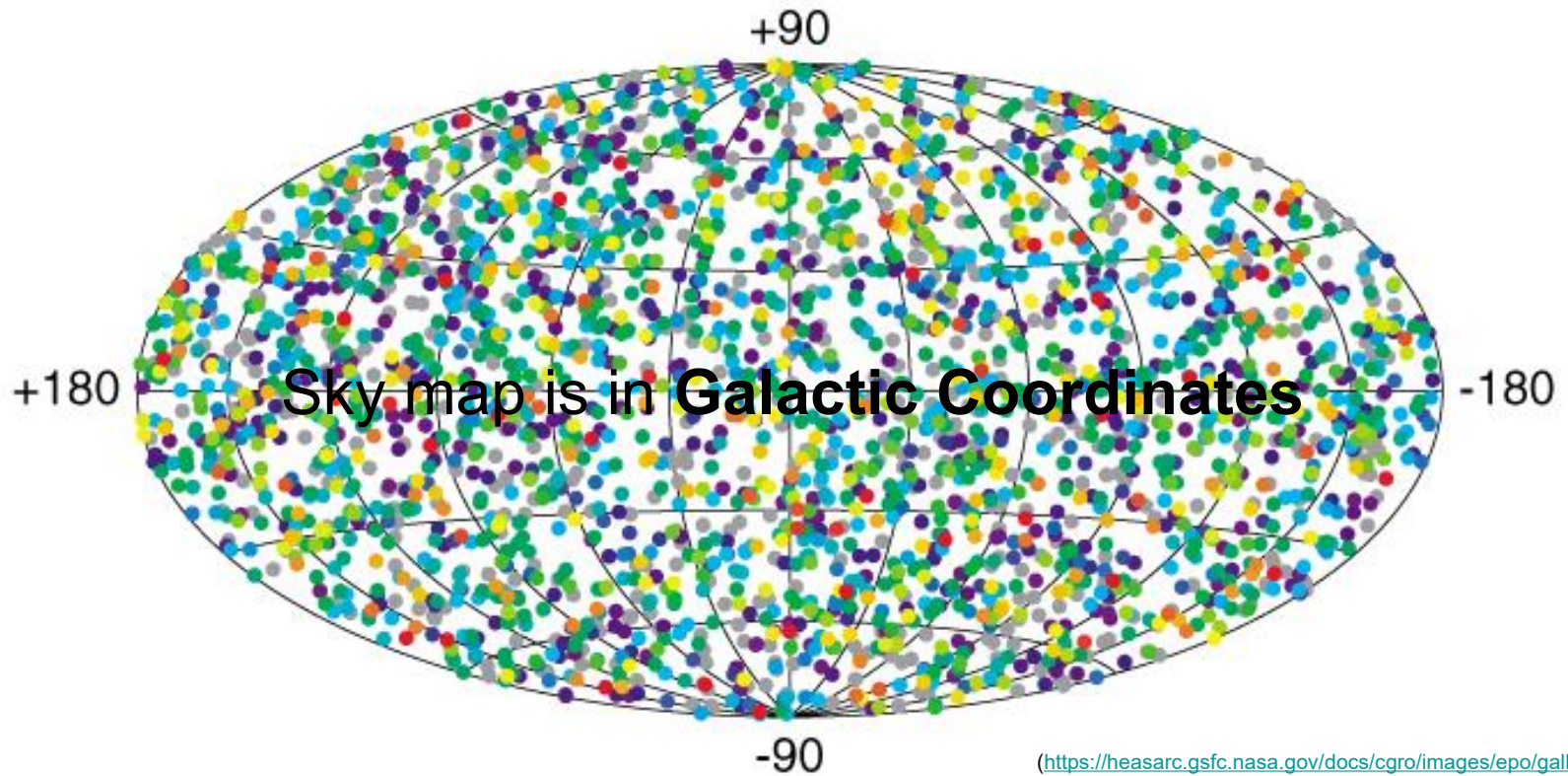
(<http://catdir.foc.gov/catdir/samples/carr031/2002283188.pdf>)

*Ray Klebsadel* (on the right) discovered the existence of gamma ray bursts in 1969 from observations made with the military satellite Vela. To his left are *Graziella Pizzichini* and *Chryssa Kouveliotou*.

# Gamma-ray burst detection is now an *industry*

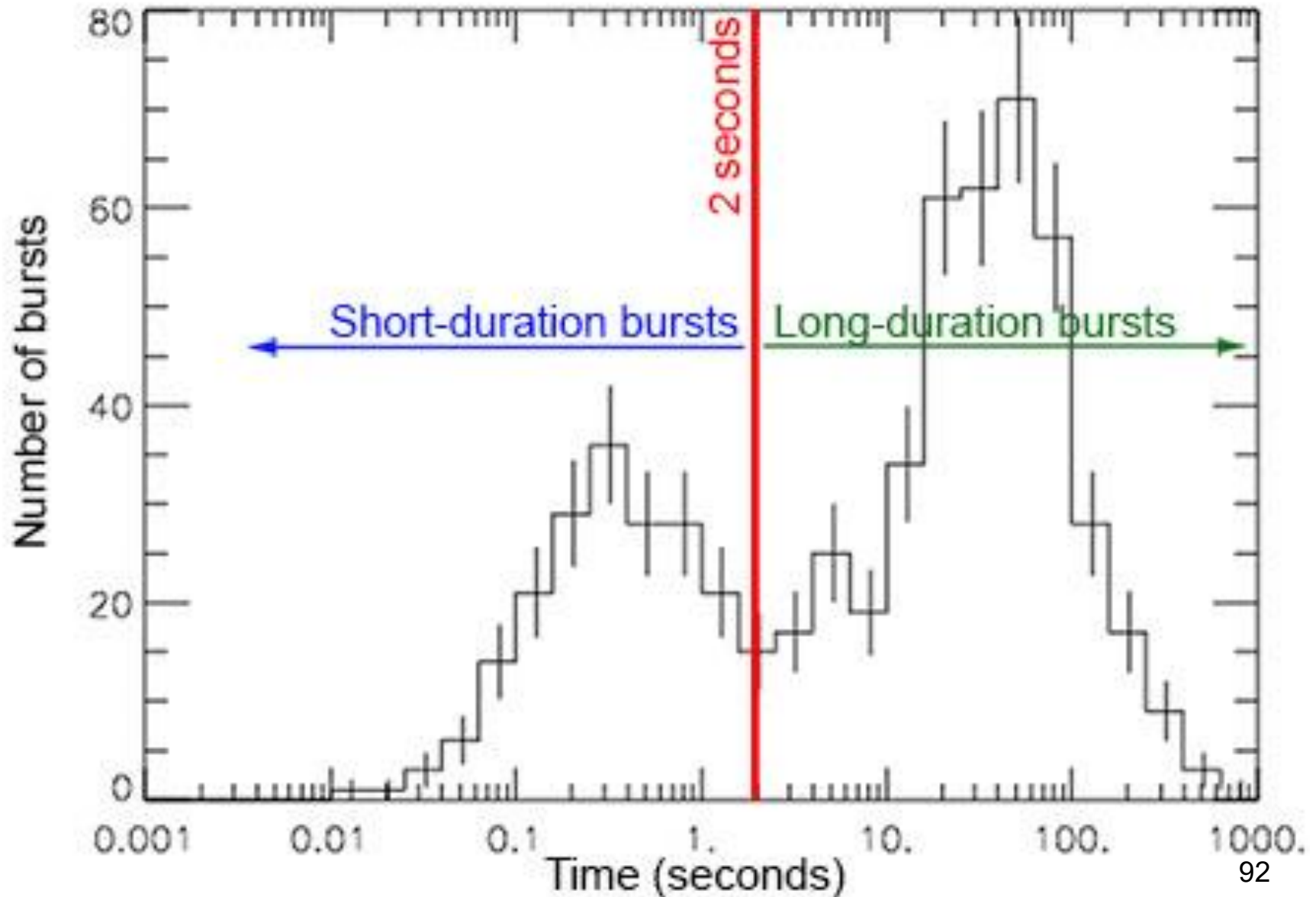
Today the uniform distribution of Gamma Ray Bursts **exclude** our galaxy and support an **extra-galactic** origin.

## 2704 BATSE Gamma-Ray Bursts





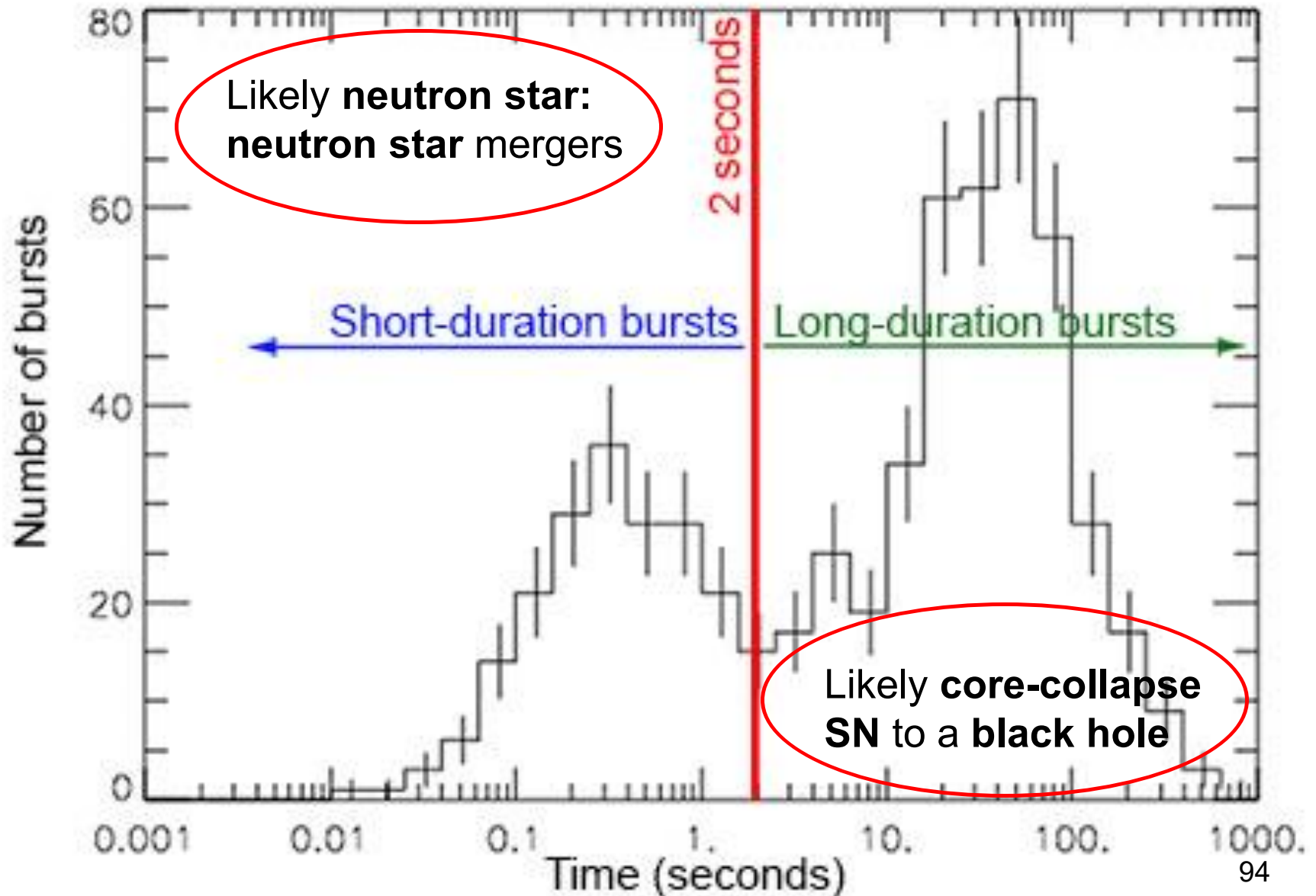
# Gamma-ray bursts' time-extent divide into two categories



## Gamma-ray burst optical counterparts: 24 years later

- 1997 - a gamma-ray burst (GRB 970508) was localized in direction and time for optical follow-up.
- By comparing photographs taken on May 8 and 9, 1997, one object was found to have increased in brightness.
- May 10 and 11, Charles Steidel recorded the spectrum of the variable object from the Keck Observatory.
- Mark Metzger analyzed the spectrum and determined a redshift  $z \geq 0.835$ , placing the burst at a distance of at least 6.6 billion light years: definitely of **non-local origin!**
- Like Quasars, something seen from that **distance** must be **extremely unusual!**

# Gamma-ray bursts remain a “hot” topic

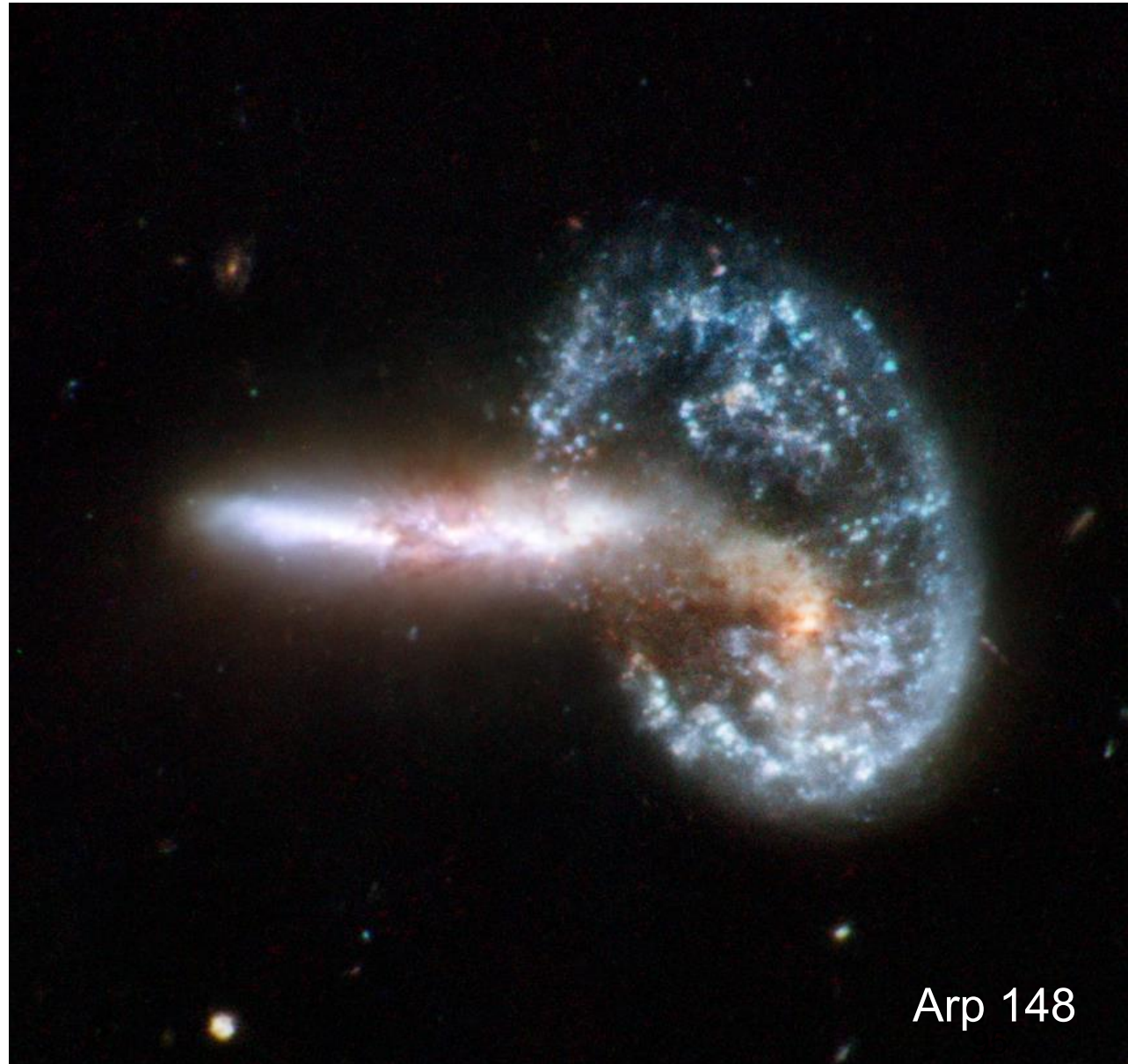




## In 1968, Arp's *peculiar galaxies* were powerful visuals

In 1968, Halton Arp showed images from his 1966 catalog: “Atlas of **Peculiar Galaxies**” at an astronomy symposium at the U. of Toronto.

**His presentations** were the most **electrifying** at the meeting!



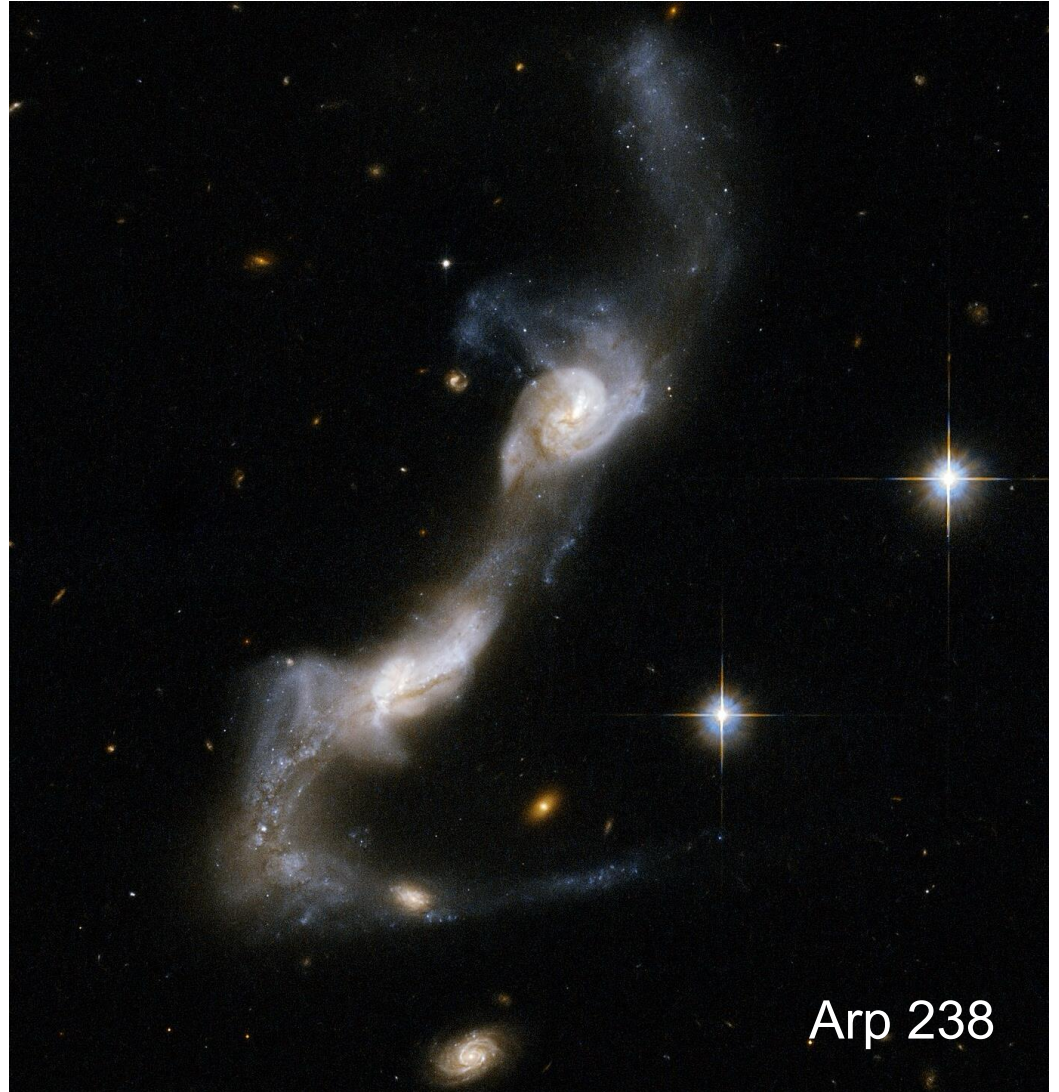
Arp 148

## Arp had ready access to the 200" Hale telescope

Arp focused on several topics including galaxies which showed unusual, or perturbed arms, or filamentary extensions.



(<https://www.astroleague.org/al/obsclubs/arppec/arphalt.html> )

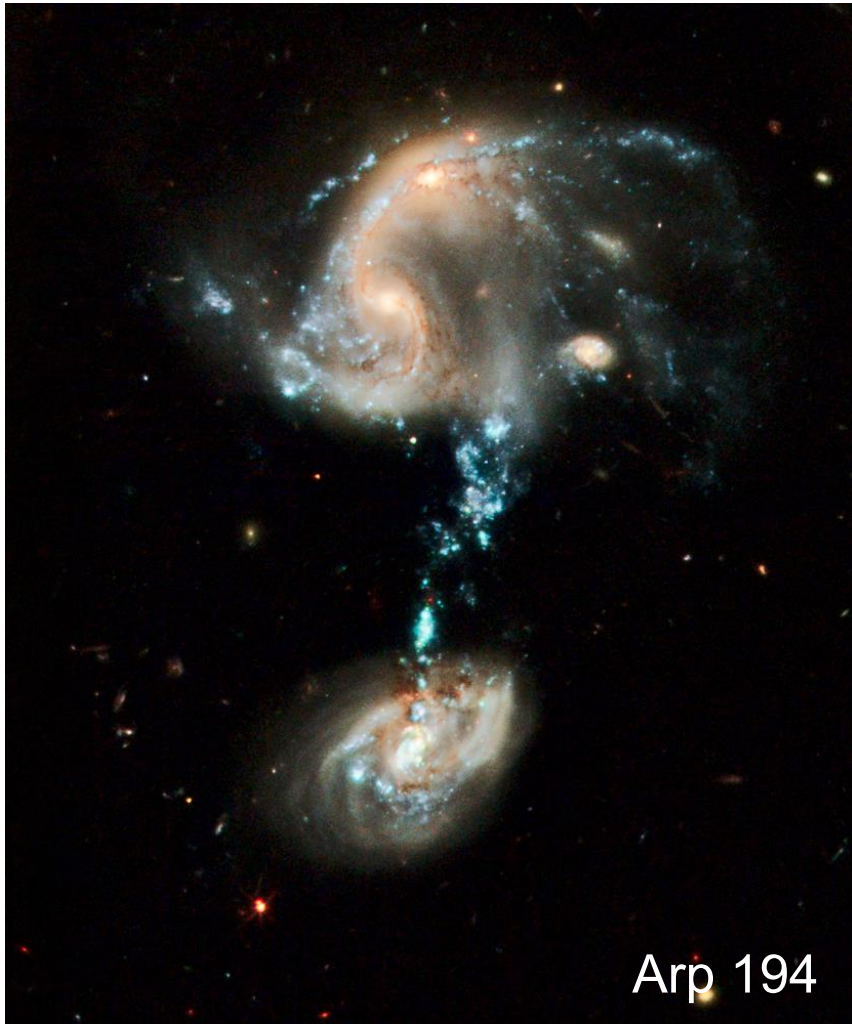


Arp 238



# What are these distorted galaxies?

Many of his images showed what appeared to be streams, or material-flow, linking the galaxies.



Arp 194



Arp 273

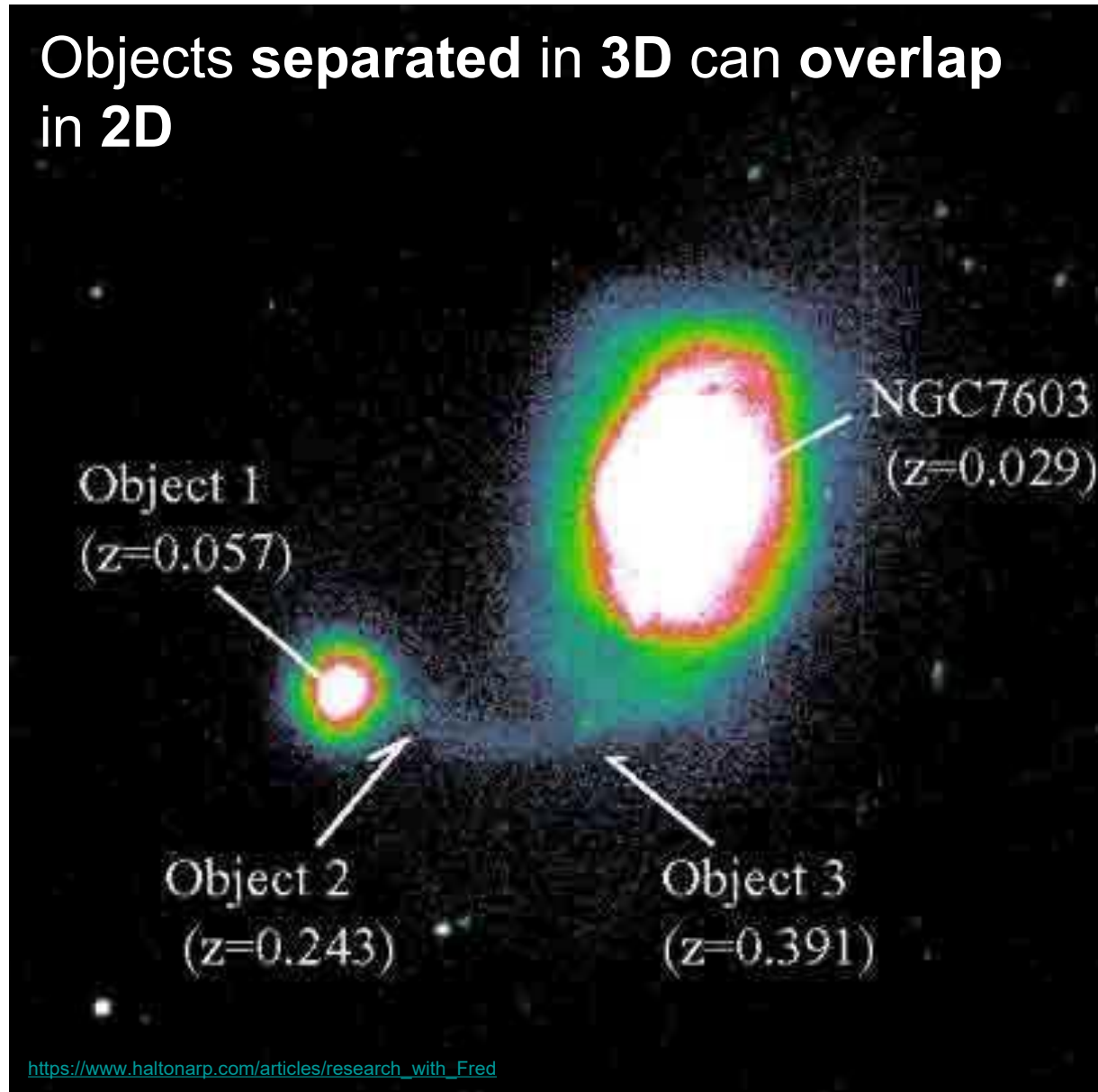


## Is there truly a *Redshift Controversy*? Likely no.

A few of Arp's images show streams, or material-flow, **linking sources at very different red shifts!**

Based on these *pairings* Arp **challenged** the assignment of **redshifts** and the related **expansion** of the **universe** in his 1973 book: *The Redshift Controversy*.

Objects **separated** in 3D can **overlap** in 2D



# Today galaxy collisions are known to be commonplace

Objects thought exotic 60 years ago are now known to be **colliding galaxies**.

**Arp 122**, originally classified as a peculiar galaxy, in fact comprises two galaxies – NGC 6040, the tilted, warped spiral galaxy and LEDA 59642, the round, face-on spiral – that are in the midst of a collision.

Likely 5% ~ 25% of all galaxies are currently merging.

<https://science.nasa.gov/missions/hubble/hubble-captures-a-monster-merger/>

# What's left? For many, Arp's galaxies have a timeless beauty

Arp's "Atlas of Peculiar Galaxies":  
still popular on Amazon



NGC 5395 & NGC 5394, Arp 84  
Gran Telescopio CANARIAS (GTC)



(<http://www.gtc.iac.es/multimedia/imageGallery.php> )



Arp 244

(Arp 244:  
[https://en.wikipedia.org/wiki/Atlas\\_of\\_Peculiar\\_Galaxies](https://en.wikipedia.org/wiki/Atlas_of_Peculiar_Galaxies) )



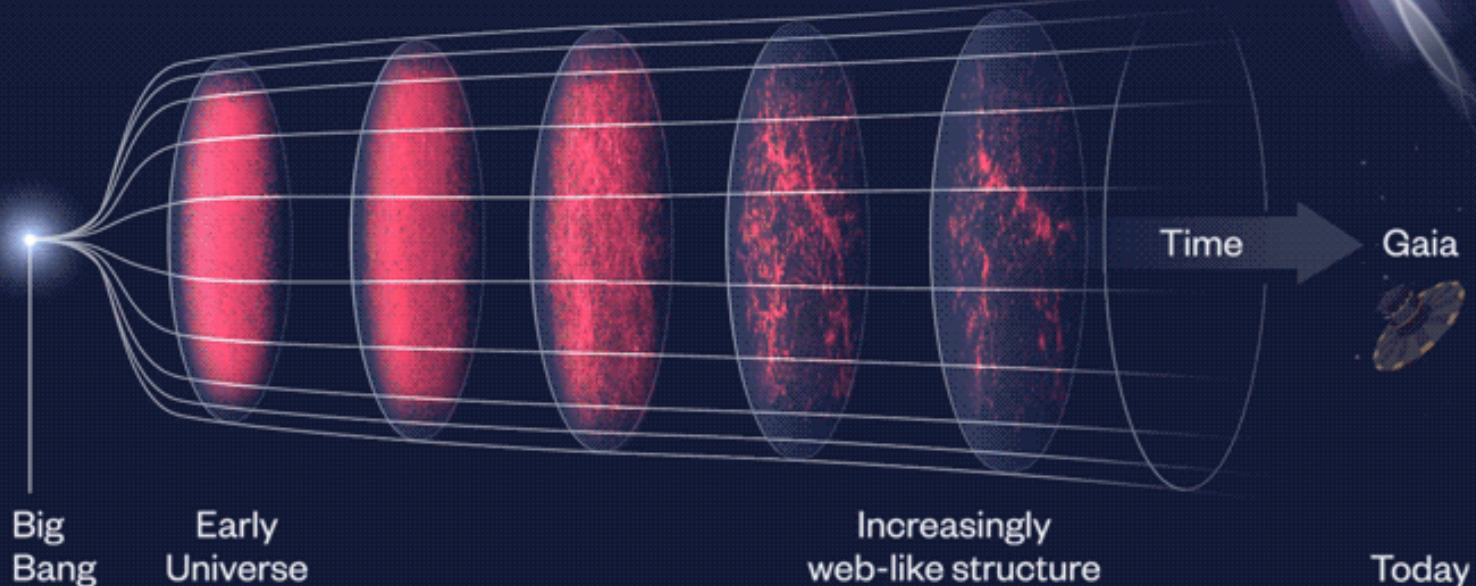
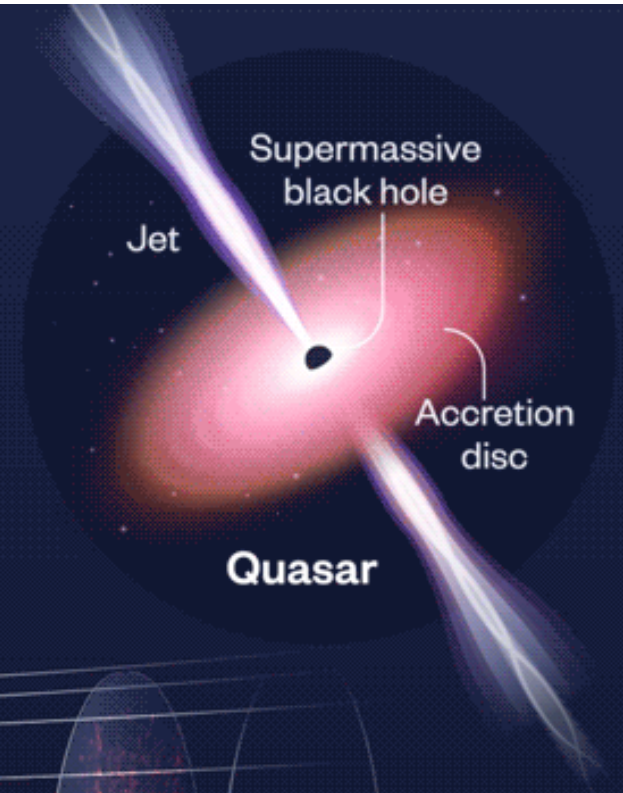
## Summary - 1960s serendipitous discoveries in Astronomy

- **Quasars:** brightest sources in the universe explained by **supermassive black holes** in a *peak feeding period*
- **CMB:** remnant photons from the tiny (1 in  $10^9$ ) matter-anti-matter asymmetry and annihilation about 1 second after the Big Bang. The most concrete evidence for the Big Bang theory
- **Pulsars:** explained by *lighthouse* emissions from **young**, rapidly rotating **neutron stars**
- **Gamma-ray bursts:** likely **neutron star:neutron star** mergers and/or **core collapse hypernovas** to **black holes**. Brightness is likely the result of significant *beamed* emissions.

# Quasars today? Mapping the universe's web-like structure

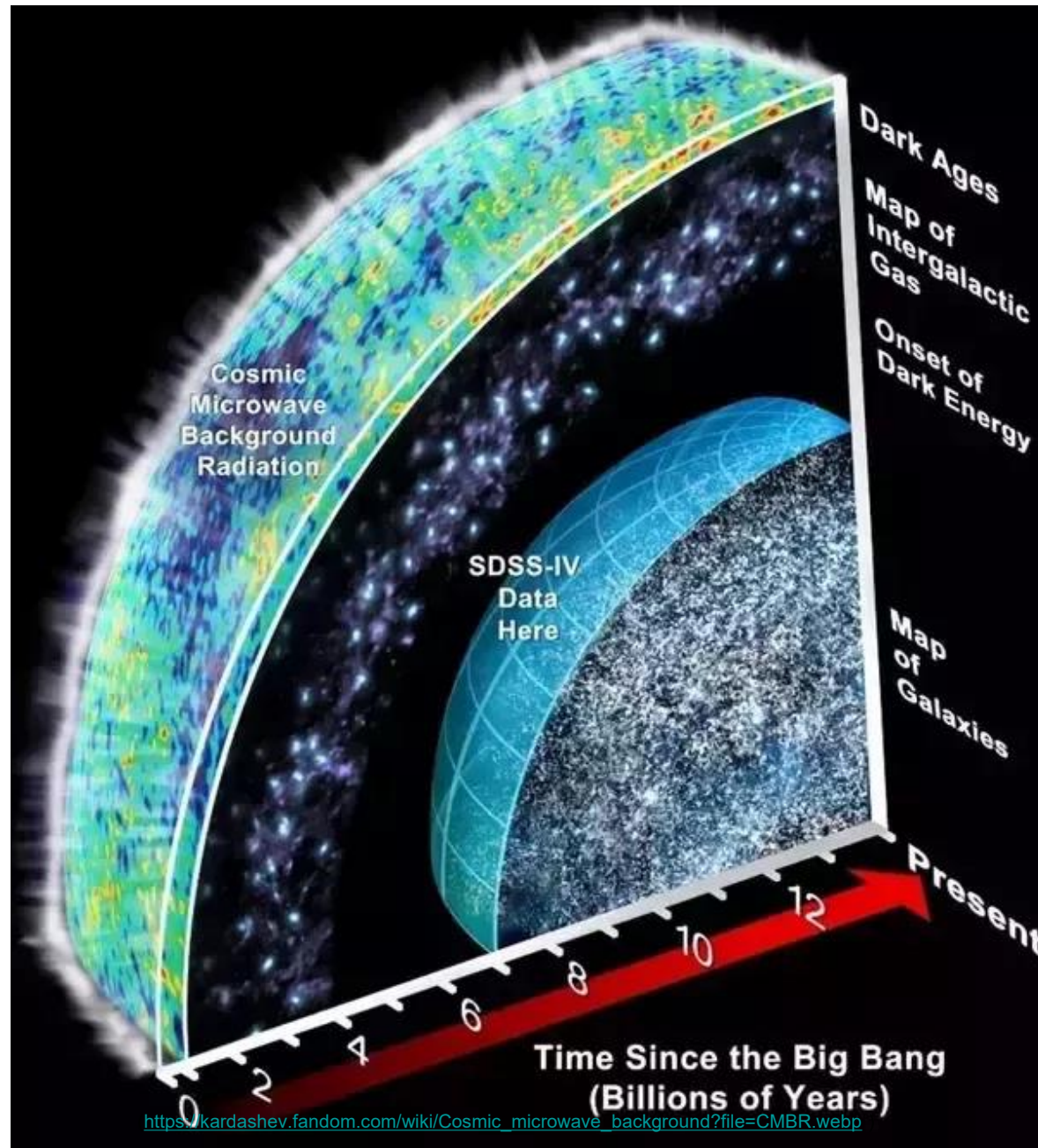
## Mapping the Bright Cosmos

Using data gathered by ESA's Gaia space observatory, astronomers have constructed the largest ever cosmic map of quasars. Thanks to their brightness, quasars can be observed throughout the cosmos, revealing the regions with large amounts of dark matter. Studying how quasars cluster together, and thus how dark matter clusters, can help astronomers gauge how the structure of our Universe formed over time.



[https://www.esa.int/ESA\\_Multimedia/Images/2024/03/Gaia\\_maps\\_largest\\_ever\\_collection\\_of\\_quasars\\_in\\_space\\_and\\_time](https://www.esa.int/ESA_Multimedia/Images/2024/03/Gaia_maps_largest_ever_collection_of_quasars_in_space_and_time)

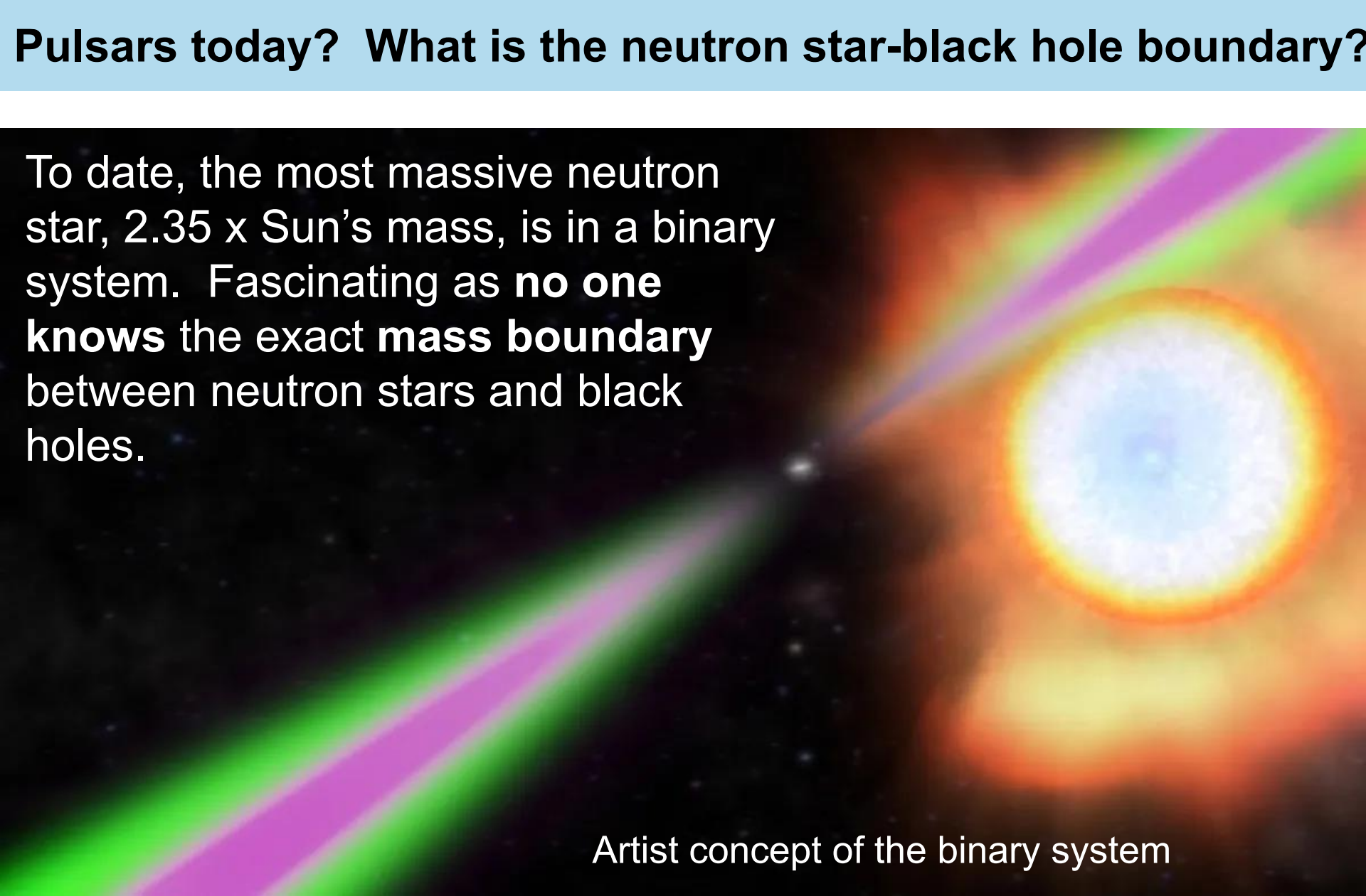
# CMB today? Tiny CMB variations seed today's web of galaxies





# Pulsars today? What is the neutron star-black hole boundary?

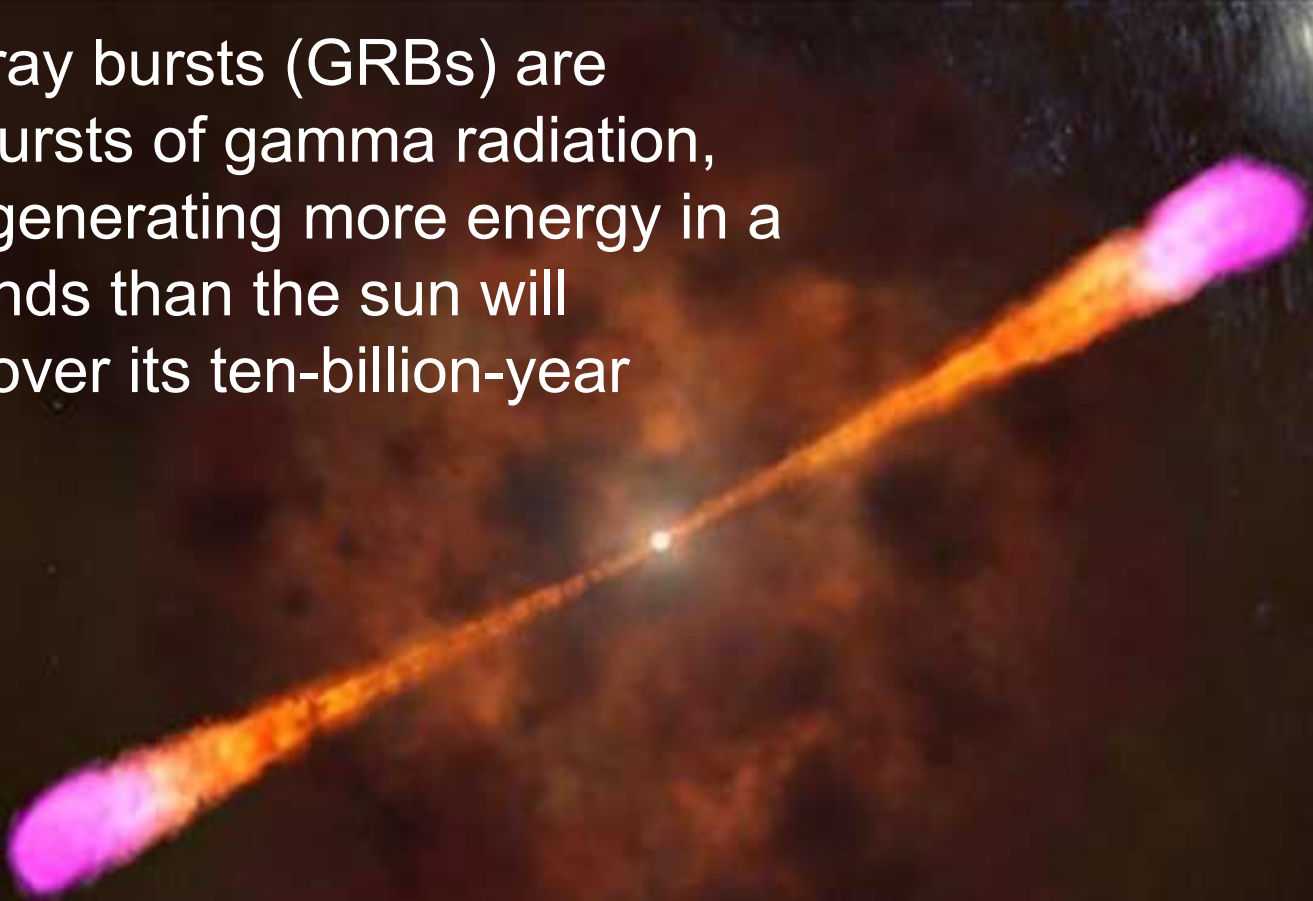
To date, the most massive neutron star,  $2.35 \times$  Sun's mass, is in a binary system. Fascinating as **no one knows** the exact **mass boundary** between neutron stars and black holes.

An artist's concept of a binary system. On the right, a bright, glowing blue and white sphere represents a neutron star, surrounded by a thick, orange and red accretion disk. On the left, a bright green and blue beam of light, representing a pulsar beam, extends towards the center. The background is a dark, star-filled space.

Artist concept of the binary system

# GRBs today? The details are still a puzzle after 50+ years

Gamma-ray bursts (GRBs) are intense bursts of gamma radiation, typically generating more energy in a few seconds than the sun will produce over its ten-billion-year lifetime.



Artist concept of a jet of particles piercing a star as it collapses into a black hole.

1960's-serendipity is still vital science today!



# Thank you

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Professor of Physics  
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[johnm@unm.edu](mailto:johnm@unm.edu)

1960s: Decade of Serendipity (in Astronomy)